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Physics of Nuclei with an Electron Ion Collider

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Mesons, Baryons, Nuclei, Nuclear Matter, Neutron Stars

- Many-Body systems. Unique combination:
 - Quantum Dynamics
 - Relativitistic
 - Strong coupling
- Described by an elegant microscopic theory:
 - Quantum Chromodynamics (QCD)

Quantum ChromoDynamics(QCD)

- Invariant under `color' SU(3) Gauge Transformations
- Gluon fields G_{μν} are 3x3 color matrices
- Quark fields q_j are 3×1 color vectors

 $\chi = \frac{1}{4g^2} G_{\mu\nu} G_{\mu\nu} + \sum_{j} \overline{g}_j (i \partial^{\mu} D_{\mu} + m_j) q_j$ where $G_{\mu\nu}^{\alpha} \equiv \partial_{\mu} P_{\nu}^{\alpha} - \partial_{\nu} P_{\mu}^{\alpha} + i f_{be}^{\alpha} P_{\mu}^{b} P_{\nu}^{c}$ and Du= du + it An That's it!

 Almost everything we observe about hadrons and nuclei is NOT in the QCD Lagrangian.
 → Emergent phenomena!

Cosmic Puzzles I. Mass

Mass

- Roughly 97% of the mass of the visible matter of the universe comes from the emergent dynamics of QCD.
- The other 3% (quark masses) comes from the Higgs mechanism



Cosmic Puzzles II. Nuclear Binding

Nuclear Binding is a color Van der Waals force

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- np system has bound state, nn system does not.
 Universe has atoms, chemistry, & stars, not just White-Dwarfs, Neutron-Stars and black-holes.
- No bound 5-nucleon or 8-nucleon nuclei.
 → Stars burn for billions of years, not just millions.



Cosmic Puzzles III. Structure and Dynamics of Neutron Stars

- Phases of QCD matter at high density / low T
 - Neutron star
 - Color Superconductivity
- Mass Limits of Neutron Stars



- Measure the neutron radii of heavy nuclei to constrain the Nuclear Matter Equation of State (Pressure-densitytemperature relation)
- Dynamics: Neutron Star Mergers

Neutron Star Merger: The ultimate relativistic heavy ion collision

- LIGO Chirp and Fermi γ -ray burst
 - https://youtu.be/-Yt5EmEgz2w



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Cosmic Puzzles IV: The Vacuum

Dynamic Quark and Gluon condensates

- $\langle 0|\bar{q}(x)q(x)|0\rangle > 0$
- $\langle 0|G(x)G(x)|0\rangle > 0$
- Not static!
- Nearly massless quarks moving through the lumpy gluon condensate acquire qq
 -cloud and mass
 ~ 300 MeV.

D. Leinweber



Unique Features of the Electron Ion Collider

- High Center-of-Mass energy
 - ep Lower than HERA, but first ever eA: ²H to ²³⁸U
- High Luminosity: 100 x HERA
- Polarized Light nuclei: p, d, ³He, ^{6,7}Li ...
 - Longitudinally and transversely polarized, without the dilution of e.g. NH₃, ND₃, Butanol... targets
- Full reconstruction of nuclear final state
 - Effective zero target thickness vis a vis fixed target exp.
 - Far Forward acceptance 50x HERA
 - Spectator and evaporation neutrons are boosted:
 - Precision detection possible

Theory Tools

- Lattice QCD
- Dyson-Schwinger Eq's
- Effective Field Theory
- Light Cone Quantization
- Density Functional Theory



Scattering with Lattice QCD (2016)

Isoscalar $\pi\pi$ scattering and the σ meson resonance from QCD

Raul A. Briceño,^{1, 2, *} Jozef J. Dudek,^{1, 2, †} Robert G. Edwards,^{1, ‡} and David J. Wilson^{3, §} (for the Hadron Spectrum Collaboration)



Dynamical Origin of Mass

 Dyson-Schwinger Eq'ns (C. Roberts)



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Why a high energy EIC to investigate low energy nuclear binding?

- I. Nuclear binding: delicate cancelation of high energy effects.
 - QMC model: scalar (σ) and vector (ρ, ω) mean fields,
 Strengths ~ 500 MeV
 - Nuclear Momentum Distributions have universal tails up to ≥ 1 GeV/c.

