Baryons at finite density

Su Houng Lee

- 1. Few words on Baryons in sum rule approach
- 2. Nucleon in symmetric and asymmetric nuclear medium
- 3. Hyperon in medium
- 4. Delta in medium

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Kiesang Jeong (INT) + former/present collaborators and students

Hyperon puzzle: Slide by Tamura

C. Ishizuka et al., J.Phys. G35 (2008) 085201

"Hyperon puzzle" in neutron stars

Hyperons (Λ at least) should appear at $\rho \sim$ 2-3 ρ_0 EQS's with hyperons or kaons too soft => cannot support $M > 1.5$ M_{sun} Heavy NS's ($^{\circ}$ 2.0 M_{sun}) were observed.

PSR $\sqrt{\frac{614-2230}{2010}}$ (2010) 1.97 ± 0.04 M_{sun} PSR J0348-0432 (2013) 2.01 ± 0.04 M_{sun}

\Rightarrow Unknown repulsion at high ρ

Strong repulsion in three-body force including hyperons,

NNN, YNN, YYN, YYY? Chiral EFT is successful in NNN force. Extension to include hyperons requires

high quality YN scattering data.

■ Phase transition to quark matter ? (quark star or hybrid star)

> We need to know YN, YY, **K**barN interactions both in free space and in nuclear medium

Baryon in nuclear matter : QCD sum rule approach

- Baryon correlation function in medium $u = (1,0,0,0)$ $\Pi (p) = \int d^4x \ e^{ipx} \ \left[\ \left. \left\langle \left(q^T \Gamma q \right) q(x) , \left(\overline{q} \Gamma \overline{q} i \right) \overline{q} \left(0 \right) \right\rangle \right] = \Pi_p \cdot p + \Pi_s \cdot 1 + \Pi_V \cdot \psi$
	- \rightarrow Phenomenological side: Positive parity states

$$
\Pi \Rightarrow \frac{\mathcal{P} + M_{+}^{*} - \Sigma_{\nu} \psi}{\left(\rho_{0} - \mathcal{E}_{\rho}\right)\left(\rho_{0} - \mathcal{E}_{\rho}\right)} \qquad \qquad \mathcal{E}_{\rho} \xrightarrow{\qquad \qquad \mathcal{E}_{=0}} \Rightarrow \Sigma_{\nu} + M_{+}^{*}
$$

 \rightarrow Operator Product Expansion

....

$$
\Pi_p = -\frac{C_p}{64\pi^4} p^4 \ln(-p^2) + ...
$$
\n
$$
\Pi_s = \frac{C_s}{4\pi^2} p^2 \ln(-p^2) \langle \overline{q}q \rangle + ...
$$
\n
$$
\Pi_v = \frac{C_v}{3\pi^2} p^2 \ln(-p^2) \langle \overline{q}\gamma_0 q \rangle + ...
$$

Scalar attraction and vector repulsion Cohen, Griegel, Furnsthal 91

$$
M_{+}^{*} = -\frac{8\pi^{2}C_{S}}{M^{2}C_{P}}\left\langle \overline{q}q \right\rangle + ...
$$

$$
\Sigma_{V} = \frac{64\pi^{2}C_{V}}{3M^{2}C_{P}}\left\langle \overline{q}\gamma_{0}q \right\rangle + ...
$$

$$
C_{P,S,V} = 1 \text{ for } \eta^{\text{loffe}} = \varepsilon^{abc} \left(q_{1a}^T C \gamma_\mu q_{1b} \right) \gamma_5 \gamma^\mu q_{2c}
$$

• *Consequence of chiral symmetry restoration* $\Pi(\rho) = \Pi_{\rho} \cdot \rho + \Pi_{s} \cdot 1 + \Pi_{u} \cdot \psi$ $(\rho_{_0} - \overline{\varepsilon}_{_\rho})\! (\rho_{_0} - \overline{\varepsilon}_{_\rho})$ v ρ_{0} – E_{p} $\left|D_{0}\right\rangle$ – E_{p} $p\llap/\,\,+\mathcal{M}^*_+-\Sigma_{\scriptscriptstyle V}\mu$ $- E_{\rho}$ (ρ_{0} – $\Pi \Rightarrow \frac{\not{D} + M_{+}^{*} - \Sigma_{\nu} \mu}{\sigma \nu}$ $_{0}$ \sim $_{p}$ N μ_{0} * $(\rho_{_0} - \overline{\varepsilon}_{_\rho})\! (\rho_{_0} - \overline{\varepsilon}_{_\rho})$ v ρ_{0} – E_{p} $\left|D_{0}\right\rangle$ – E_{p} $\rlap{\,/}D - M^*_{-} - \Sigma_{\scriptscriptstyle V} \rlap{\,/} \mu$ $- E_{\rho}$ (ρ_{0} – $\mathcal{D} - \mathcal{M}^\star_{\scriptscriptstyle{-}} - \Sigma_{\scriptscriptstyle{V}} \mathcal{U}$ $+\frac{W}{V}$ $\frac{W}{V}$ $\mathcal{O} \subset L_p \wedge \mathcal{O}$ *

$$
\Pi_s \propto \text{Tr} \big[S(x,0) \big] = \langle \overline{q} q \rangle + \ldots \to 0 \qquad M_{\pm}^* \to 0 \quad \text{or} \ \big(M_{-}^* - M_{+}^* \big) \to 0
$$

• *Finite T Lattice result (Aarts et al. arXiv:1710.00566)*

Nucleons in symmetric and asymmetric matter (K. Jeong, SHL 2013)

• *Quark operators at finite density with isospin asymmetry*

$$
\left\langle \phi_{u,d} \right\rangle_{\rho,I} = \left\langle \phi_{u,d} \right\rangle_{\text{vac}} + \left\langle \phi_{0} \right\rangle_{p} \rho \pm \left\langle \phi_{1} \right\rangle_{p} I \rho\n\qquad\n\begin{aligned}\n&\phi_{0} = \frac{1}{2} (\phi_{u} + \phi_{d}), & \phi_{1} = \frac{1}{2} (\phi_{u} - \phi_{d}) \\
&\phi = \rho_{n} + \rho_{p}, & I \rho = \rho_{n} - \rho_{p}\n\end{aligned}
$$

 \rightarrow Important operators

symmetric\n
$$
\left\{\n\begin{aligned}\n\left\langle \overline{q}q \right]_{0}\n\right\rangle_{p} &= \frac{\sigma_{N}}{2m_{q}}: 4.5 \\
\left\langle \overline{u}\gamma_{0}u \right\rangle_{p} &= 2\n\end{aligned}\n\right.
$$
\n
$$
\left\{\n\begin{aligned}\n\left\langle \overline{q}q \right]_{1}\n\right\rangle_{p} &= \frac{1}{2} \left[\left\langle \overline{u}u \right\rangle_{p} - \left\langle \overline{d}d \right\rangle_{p}\right] = \frac{1}{2} \left[\left\langle \frac{m_{n} - m_{p}}{m_{d} - m_{u}} \right] : 0.1 \left\langle \left[\overline{q}q \right]_{0} \right\rangle_{p} < 0.5\n\end{aligned}\n\right.
$$
\n
$$
\left\{\n\left[\overline{q}\gamma_{0}q \right]_{1}\n\right\rangle_{p} = \frac{1}{2} \left[\left\langle \overline{u}\gamma_{0}u \right\rangle_{p} - \left\langle \overline{d}\gamma_{0}d \right\rangle_{p}\right] = \frac{1}{3} \left\langle \left[\overline{q}\gamma_{0}q \right]_{0}\right\rangle_{p} = 0.5\n\right.
$$
\n
$$
\left\langle \left[\overline{q}\gamma_{0}D_{0}q \right]_{1}\n\right\rangle_{p} = 0.35 \left\langle \left[\overline{q}\gamma_{0}D_{0}q \right]_{0}\n\right\rangle_{p}
$$

Nucleons in symmetric and asymmetric matter

• *Symmetry energy from QCD sum rule:* K. Jeong, SHL, PRC87, 015204 (2013)

 \rightarrow Feature similar to Relativistic mean field models

$$
L = \overline{q}_N \left[\gamma_\mu \left(i \partial^\mu - g_{\omega N} \omega^\mu - g_{\rho N} \rho^\mu \right) - \left(m_N - g_{\sigma N} \sigma - g_{\delta N} \delta \right) \right] \rho_B
$$

$$
E_{V,\rho}^{\text{sym}} = \frac{1}{2} \left[\frac{g_{\rho N}^2}{m_\rho^2} - \frac{g_{\delta N}^2}{m_\delta^2} \left(\frac{m^*}{E_F^*} \right) \right] \rho_B
$$

Hyperons in symmetric and asymmetric matter

- *QCD sum rule: Interpolating field determines how well OPE converges*
- *Needed additional input*

 $\sum_{p} = \frac{\sigma_{sN}}{m} \rightarrow$ Will take $0.1 \times \langle \overline{u}u \rangle_{p}$ *s* $\langle \overline{s} s \rangle_n = \frac{\partial s_N}{\partial x}$ \rightarrow Will take $0.1 \times \langle \overline{u} u \rangle$ *m* $=\frac{\sigma_{sN}}{\sigma} \rightarrow$ Will take 0.1 × Lattice $\sigma_{sN} \rightarrow 17 \text{ MeV (JLQCD)}$ \cdot 105 MeV (BMW)

 $\left. s\gamma_0^{\vphantom{0}} s\right\rangle_p=0$

$$
\left\langle q_a^a \overline{q}_\beta^b q_\gamma^c \overline{q}_\delta^d \right\rangle_{\rho,I} = \left\langle q_a^a \overline{q}_\beta^b \right\rangle_{\rho,I} \left\langle q_\gamma^c \overline{q}_\delta^d \right\rangle_{\rho,I} - \left\langle q_a^a \overline{q}_\delta^d \right\rangle_{\rho,I} \left\langle q_\gamma^c \overline{q}_\beta^b \right\rangle_{\rho,I}
$$

^Λ*in symmetric nuclear matter*

• ^Λ *in nuclear matter* (Jeong, Gye, SHL PRC94 (2016)065201)

Operator Product Expansion when the current has a strong u-d scalar diquark

Nucleon with Ioffe current

$$
M_{+}^{*} = -\frac{8\pi^{2}}{M^{2}} \langle \overline{q}q \rangle + ...
$$

$$
\Sigma_{V} = \frac{64\pi^{2}2}{3M^{2}} \langle \overline{q}_{1} \gamma_{0} q_{1} \rangle + ...
$$

Λ with strong u-d diquark current

$$
M_{+}^{*} = -\frac{8\pi^{2}}{M^{2}} \left(2\left\langle \overline{s}s \right\rangle + \text{small} \left\langle \overline{q}q \right\rangle \right) + \dots
$$

$$
\Sigma_{V} = \frac{64\pi^{2}2}{3M^{2}} \left(\frac{1}{4} \left\langle \overline{q}_{1} \gamma_{0} q_{1} \right\rangle \dots \right) + \dots
$$

 \rightarrow Both Smaller than 2/3 from Quark counting but similar total binding

Vector repulsion < Scalar attraction

$$
\frac{\Sigma_V(\Lambda)}{\Sigma_V(N)} \approx 0.26
$$

$$
\frac{\Sigma_{s}\left(\Lambda\right)}{\Sigma_{s}\left(N\right)} \approx 0.3
$$

^Σ*in symmetric nuclear matter*

• ^Σ *in nuclear matter* (Jeong, Gye, SHL PRC94 (2016)065201)

Operator Product Expansion is similar with Ioffe current with $d\rightarrow s$

Nucleon with Ioffe current

$$
M_{+}^{*} = -\frac{8\pi^{2}}{M^{2}} \langle \overline{q}q \rangle + ...
$$

$$
\Sigma_{V} = \frac{64\pi^{2}2}{3M^{2}} \langle \overline{q}_{1} \gamma_{0} q_{1} \rangle + ...
$$

 Λ with loffe current d \rightarrow s

$$
M_{+}^{*} = -\frac{8\pi^{2}}{M^{2}} \left(\langle \overline{s}s \rangle + \text{small} \langle \overline{q}q \rangle \right) + ...
$$

$$
\Sigma_{V} = \frac{64\pi^{2} 2}{3M^{2}} \left(\frac{7}{8} \langle \overline{q}_{1} \gamma_{0}q_{1} \rangle ... \right) + ...
$$

Vector repulsion Scalar attraction

$$
\frac{\Sigma_V(\Lambda)}{\Sigma_V(N)} \approx 1 \qquad \Longrightarrow \qquad
$$

$$
\frac{\Sigma_{S}\left(\Lambda\right)}{\Sigma_{S}\left(N\right)} \approx 0.3
$$

 \rightarrow Total repulsion is larger than 100 MeV at nuclear matter

Delta in nuclear matter

• *Baryon in Relativistic Mean Field models*

$$
L = \overline{N} \Big[\gamma_{\mu} \Big(i \partial^{\mu} - g_{\omega N} \omega^{\mu} - g_{\rho N} \tau \rho^{\mu} \Big) - \Big(m_N - g_{\sigma N} \sigma \Big) \Big] N
$$

+
$$
\overline{\Delta}_{\nu} \Big[\gamma_{\mu} \Big(i \partial^{\mu} - g_{\omega \Delta} \omega^{\mu} - g_{\rho \Delta} \tau \rho^{\mu} \Big) - \Big(m_{\Delta} - g_{\sigma \Delta} \sigma \Big) \Big] \Delta^{\nu}
$$

 \rightarrow Effects in neutron star [Cai, Fattoyev, Bao-An Li, Newton (PRC92 (2015) 015802)] $x_{\sigma} = g_{\sigma \Delta}/g_{\sigma N}$, $x_{\omega} = g_{\omega \Delta}/g_{\omega N}$, $x_{\rho} = g_{\rho \Delta}/g_{\rho N}$

Experimental results on Delta mass shift

- *At symmetric Nuclear matter in* ^γ *-A*
	- \rightarrow Broadening
	- \rightarrow No clear observable Mass change?
- *Heavy Ion Collisions*

possible mass shift at higher density

Analysis of pion Pt spectra (Rafelski 2007) BNL: Au+Au

• *Raon*

Can produce through subthreshold production in heavy ion collision

• [∆] *interpolating currents* (Marques, SHL, A. Park, Matheus, Jeong PRC98 (2018)025206)

Nucleon Ioffe current Nucleon Ioffe current

$$
\eta_{\mu}^{\text{Ioffe}} = \varepsilon^{abc} \left(q_{1a}^T C \gamma_{\mu} q_{1b} \right) \gamma_5 \gamma^{\mu} q_{2c}
$$
\n
$$
\Pi_p = -\frac{1}{64\pi^4} p^4 \ln(-p^2) + \dots
$$
\n
$$
\Pi_s = \frac{1}{4\pi^2} p^2 \ln(-p^2) \langle \overline{q}q \rangle + \dots
$$
\n
$$
\Pi_s = \frac{1}{3\pi^2} p^2 \ln(-p^2) \langle \overline{u} \gamma_0 u \rangle + \dots
$$
\n
$$
\Pi_p = \frac{1}{160\pi^4} p^4 \ln(-p^2) \ln(-p^2) \langle \overline{u} \gamma_0 u \rangle + \dots
$$
\n
$$
\Pi_p = \frac{1}{3\pi^2} p^2 \ln(-p^2) \ln(-p^2) \langle \overline{u} \gamma_0 u \rangle + \dots
$$
\n
$$
\Pi_p = \frac{1}{3\pi^2} p^2 \ln(-p^2) \ln(-p^2) \langle \overline{u} \gamma_0 u \rangle + \dots
$$
\n
$$
\Pi_p = \frac{1}{4\pi^2} p^2 \ln(-p^2) \ln(-p^2) \langle \overline{u} \gamma_0 u \rangle + \dots
$$
\n
$$
\Pi_p = \frac{1}{4\pi^2} p^2 \ln(-p^2) \ln(-p^2) \langle \overline{u} \gamma_0 u \rangle + \dots
$$

 Π_{V} = $\frac{8}{3} \times$ (condensates) 3 *V s* $\rightarrow \frac{\Pi_{V}}{\Pi_{s}} = \frac{8}{3} \times (\text{condensates}) \rightarrow \frac{\Pi_{V}}{\Pi_{s}} = \frac{3}{4} \times (\text{condensates})$

$$
\eta_{\mu}^{\text{left}} = \varepsilon^{abc} \left(q_{1a}^T C \gamma_{\mu} q_{1b} \right) q_{1c}
$$
\n
$$
\Pi_p = \frac{1}{160\pi^4} p^4 \ln \left(-p^2 \right) + \dots
$$
\n
$$
\Pi_s = -\frac{1}{3\pi^2} p^2 \ln \left(-p^2 \right) \langle \overline{u}u \rangle + \dots
$$
\n
$$
\Pi_V = -\frac{1}{4\pi^2} p^2 \ln \left(-p^2 \right) \langle \overline{u} \gamma_0 u \rangle + \dots
$$

$$
\rightarrow \frac{\Pi_{V}}{\prod_{s} = \frac{3}{4} \times (\text{condensates})}
$$

- [∆]*Has smaller vector repulsion compared to nucleon*
- ∆ *scalar attraction : 150 MeV,* [∆] *Vector repulsion: 75 MeV*

Delta in asymmetric nuclear matter

• [∆] *vector self energy* (Marques, SHL, A. Park, Matheus, Jeong PRC98 (2018)025206)

 $\Sigma_V(\rho, I) = (0.075 - 0.01 \tau_A^3 I) \text{ GeV}$ $\tau_{\Delta^{++}}^3 = 3$, $\tau_{\Delta^+}^3 = 1$, $\tau_{\Delta^0}^3 = -1$, $\tau_{\Delta^-}^3 = -3$,

• *Nucleon vector self energy* (Marques, SHL, A. Park, Matheus, Jeong PRC98 (2018)025206)

 $\Sigma_{V}(\rho, I) = (0.32 - 0.08 I)$ GeV

Smaller vector repulsion for Delta in quark model

- *Nucleon Nucleon interaction at short distance is dominated by color spin interaction* (arXiv:1907.06351 A Park, SHL, Inoue, Hatsuda)
- Color-spin interaction

$$
H = \sum_{i=1}^N \left(m_i + \frac{p_i^2}{2m_i} \right) - \sum_{i < j}^N \left(\lambda_i^c \lambda_j^c \right) V_{ij}^C \left(r_{ij} \right) - \sum_{i < j}^N \frac{\left(\lambda_i^c \lambda_j^c \right) \left(\sigma_i \sigma_j \right)}{m_i m_j} V_{ij}^{SS} \left(r_{ij} \right)
$$

☞ Color-spin interaction for 2 body: **^q ^q ^q ^q**

$$
K=-\sum_{i
$$

- Comparison with lattice (HAL QCD) NN interaction
- ☞ NN force in SU(2) spin 1 vs spin 0 channel (W.Park, A. Park, Lee 2015)

☞ H dibaryon channel: Flavor 1 vs Flavor 27

- K factors in NN vs N∆
- **■** Spin and Isospin averaged NN interaction: repulsive with $K_{2-N} \approx 2.37$

■ Spin and Isospin averaged N∆ interaction: repulsive with $K_{2-N} \approx 1.9$

Summary

- 1. QCD sum rule approach offers a way to calculate the potential part of the nucleon, delta and hyperons in medium \rightarrow similar to RMF models.
- 2. In sum rule, Symmetry energy is positive and due to larger vector repulsion and scalar attraction.
- 3. Lambda has smaller self energies than 2/3 of nucleon, but repulsion is even smaller
- 4. Sigma repulsion in matter is larger than naïve expectation. 100 MeV
- 5. Delta has smaller vector repulsion compared to nucleon
- 6. Further studies with different currents will be useful