

# Collective excitations and neutrino-nucleus reactions in relativistic point-coupling models

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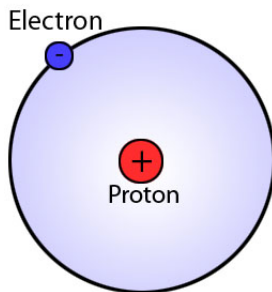
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## Simple system - Hydrogen atom



## Simple system - Hydrogen atom

Dirac Hamiltonian for central field:

$$\hat{H} = c\hat{\alpha} \cdot \hat{\mathbf{p}} + \beta m_0 c^2 + V(r) \quad (1)$$

If the field has spherical symmetry, then total angular momentum operator  $\hat{J}$  and operator  $\hat{P}$  ( $\hat{P} = \exp(i\phi)\hat{\beta}\hat{P}_0$ ,  $P_0 : \vec{x} \rightarrow -\vec{x}$ ) commute with  $\hat{H}$ .

We assume solution:

$$\Psi_{njm} = \begin{pmatrix} \psi_{ljm}(\vec{r}, t) \\ \chi_{\bar{l}jm}(\vec{r}, t) \end{pmatrix} \quad (2)$$

Action of parity operator:

$$\hat{P}_0 \psi_{ljm}(\vec{r}, t) = \lambda_f \psi_{ljm}(\vec{r}, t) \quad (3)$$

$$-\hat{P}_0 \chi_{\bar{l}jm}(\vec{r}, t) = -\lambda_g \chi_{\bar{l}jm}(\vec{r}, t) \quad (4)$$

We must have  $\lambda_f = -\lambda_g$  in order to have good parity of four-spinor.

## Simple system - Hydrogen atom

$$\hat{H}\Psi_{jm}(\vec{r}, t) = -i\hbar \begin{pmatrix} 0 & \hat{\vec{\sigma}} \\ \hat{\vec{\sigma}} & 0 \end{pmatrix} \nabla \Psi_{jm}(\vec{r}, t) + \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} m_0 c^2 \Psi_{jm}(\vec{r}, t) + V(r) \Psi_{jm}(\vec{r}, t) \quad (5)$$

Rewritten in terms of bispinors:

$$-i\hbar c \vec{\sigma} \cdot \vec{\nabla} \chi_{\bar{l}jm} + m_0 c^2 \psi_{ljm} + V(r) \psi_{ljm} = E \psi_{ljm} \quad (6)$$

$$-i\hbar c \vec{\sigma} \cdot \vec{\nabla} \psi_{\bar{l}jm} - m_0 c^2 \chi_{ljm} + V(r) \chi_{ljm} = E \chi_{ljm} \quad (7)$$

For  $\psi$

$$\psi_{ljm} = R_{nl}(r) \sum_{m_l m_s} C_{1/2 m_s m_l}^{jm} Y_{l m_l}(\Omega) \chi_{1/2 m_s}(\vec{\sigma}) \quad (8)$$

and similar for  $\chi$ .

Expansion of bispinors in spherical harmonic oscillator basis:

$$\psi_{nljm} = \sum_n C_n^l f_{nl}(r) \Omega_{ljm}(\hat{\Omega}) \quad (9)$$

$$\chi_{n\bar{l}jm} = \sum_n C_n^{\bar{l}} g_{n\bar{l}}(r) \Omega_{\bar{l}jm}(\hat{\Omega}) \quad (10)$$

## Simple system - Hydrogen atom

Matrix elements of Dirac Hamiltonian:

$$A_{fg} = -\hbar c \int_0^\infty dr r^2 f_{n_a l_a}^*(r) \frac{\partial}{\partial r} g_{n_c \bar{l}_a}(r) \quad (11)$$

$$B_{fg} = -\hbar c \int_0^\infty dr r^2 f_{n_a l_a}^*(r) \frac{1-\kappa}{r} g_{n_c \bar{l}_a}(r) \quad (12)$$

$$C_{ff}^{(S)} = \int dr r^2 f_{n_a l_a}^*(r) S(r) f_{n_c l_a}(r) \quad (13)$$

$$C_{ff}^{(V)} = \int dr r^2 f_{n_a l_a}^*(r) V(r) f_{n_c l_a}(r) \quad (14)$$

$$A_{gf} = \hbar c \int_0^\infty dr r^2 g_{n_a \bar{l}_a}^*(r) \frac{\partial}{\partial r} f_{n_c l_a}(r) \quad (15)$$

$$B_{gf} = \hbar c \int_0^\infty dr r^2 g_{n_a \bar{l}_a}^*(r) \frac{1+\kappa}{r} f_{n_c l_a}(r) \quad (16)$$

$$C_{gg}^{(S)} = - \int dr r^2 g_{n_a \bar{l}_a}^*(r) S(r) g_{n_c \bar{l}_a}(r) \quad (17)$$

$$C_{gg}^{(V)} = \int dr r^2 g_{n_a \bar{l}_a}^*(r) V(r) g_{n_c \bar{l}_a}(r) \quad (18)$$

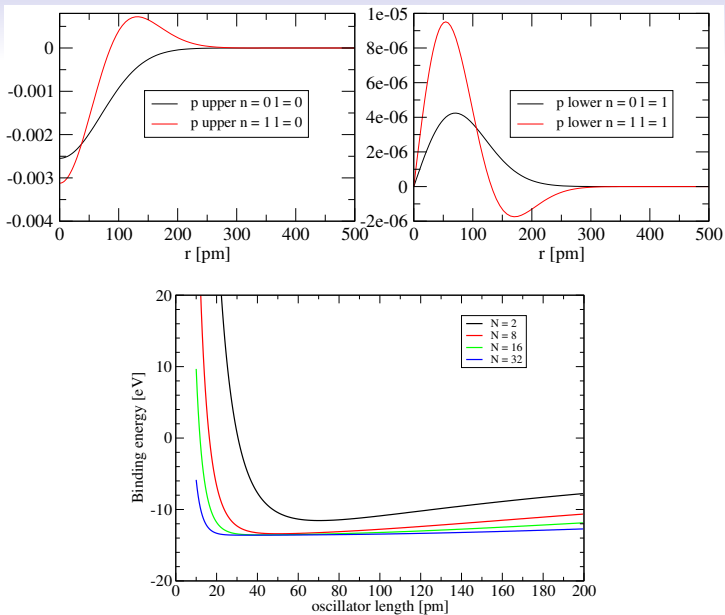
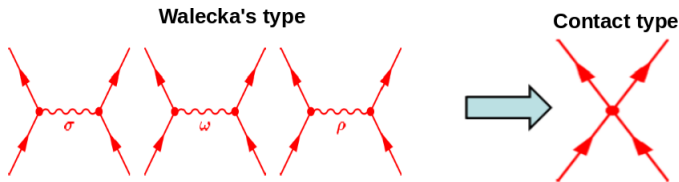


Figure 1: Upper and lower components of spinor  $s_{1/2}$ . Dependence of binding energy of hydrogen atom on harmonic oscillator length for different numbers of major oscillator shells.

## Relativistic nuclear energy density functionals

Two types of nuclear models are usually present in relativistic energy density functionals:

- Walecka's type (nucleons are system of Dirac particles which interacts by exchange of mesons)
- Contact type (point-coupling; where meson propagator is replaced by contact interaction)



In nuclear models of Walecka's type we have meson propagator:

$$\frac{g_m}{-\Delta + m_m^2} \simeq \frac{g_m}{m_m^2} + g_m \frac{\Delta}{m_m^4}, \quad (19)$$

Four-fermion interaction:

$$L_{4f} = -\frac{1}{2}\alpha_S (\bar{\psi}\psi) (\bar{\psi}\psi) - \frac{1}{2}\alpha_V (\bar{\psi}\gamma_\mu\psi) (\bar{\psi}\gamma^\mu\psi) - \frac{1}{2}\alpha_{TV} (\bar{\psi}\vec{\tau}\gamma_\mu\psi) (\bar{\psi}\vec{\tau}\gamma^\mu\psi), \quad (20)$$

Derivative terms:

$$L_{der} = -\frac{1}{2}\delta_S\partial_\nu (\bar{\psi}\psi) \partial^\nu (\bar{\psi}\psi) - \frac{1}{2}\delta_V\partial_\nu (\bar{\psi}\gamma_\mu\psi) \partial^\nu (\bar{\psi}\gamma^\mu\psi) - \frac{1}{2}\delta_{TV}\partial_\nu (\bar{\psi}\vec{\tau}\gamma_\mu\psi) \partial^\nu (\bar{\psi}\vec{\tau}\gamma^\mu\psi), \quad (21)$$

We can add non-linear terms:

$$L_{hot} = -\frac{1}{3}\beta_S (\bar{\psi}\psi)^3 - \frac{1}{4}\gamma_S (\bar{\psi}\psi)^4, \quad (22)$$

Dirac Lagrangian for free nucleon:

$$L_{free} = \bar{\psi} (i\gamma_\mu\partial^\mu - m) \psi \quad (23)$$

Electromagnetic part:

$$L_{em} = -eA_\mu\bar{\psi} [(1 - \tau_3)/2] \gamma^\mu\psi - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \quad (24)$$



## DD-PC1 parametrization

Lagrangian density:

$$\begin{aligned}
 L = & \bar{\psi} (i\gamma_{\mu}\partial^{\mu} - m) \psi - \frac{1}{2}\alpha_S (\bar{\psi}\psi) (\bar{\psi}\psi) - \frac{1}{2}\alpha_V (\bar{\psi}\gamma_{\mu}\psi) \\
 & \times (\bar{\psi}\gamma^{\mu}\psi) - \frac{1}{2}\alpha_{TV} (\bar{\psi}\vec{\tau}\gamma_{\mu}\psi) (\bar{\psi}\vec{\tau}\gamma^{\mu}\psi) \\
 & - \frac{1}{2}\delta_S\partial_{\nu} (\bar{\psi}\psi) \partial^{\nu} (\bar{\psi}\psi) - e\bar{\psi}\gamma_{\mu}A^{\mu} \frac{1-\tau_3}{2}\psi. \quad (25)
 \end{aligned}$$

Coupling density dependence in case of DD-PC1 parametrization:

$$f_i [\rho] = a_i + (b_i + c_i x) \exp(-d_i x), \quad (26)$$

Variation of Lagrangian with respect to  $\bar{\psi}$  we obtain one-nucleon Dirac equation:

$$[\gamma_{\mu} (i\partial^{\mu} - \Sigma^{\mu} - \Sigma_{rearr}^{\mu}) - (m + \Sigma_S)] \psi = 0, \quad (27)$$

$$\Sigma^{\mu} = \Sigma_V^{\mu} + \vec{\tau} \cdot \vec{\Sigma}_{TV}^{\mu}, \quad (28)$$

$$\Sigma_S = \alpha_S(\rho)\rho_S + \delta_S\Delta\rho_S, \quad (29)$$

$$\Sigma_V^{\mu} = \alpha_V(\rho)j^{\mu} + e\frac{1-\tau_3}{2}A^{\mu}, \quad (30)$$

$$(31)$$

$$\vec{\Sigma}_{TV}^{\mu} = \alpha_{TV}(\rho) \vec{j}^{\mu}, \quad (32)$$

$$\Sigma_{rearr}^{\mu} = \frac{1}{2} \frac{j^{\mu}}{\rho} \left[ \frac{\partial \alpha_S(\rho)}{\partial \rho} \rho_S^2 + \frac{\partial \alpha_V(\rho)}{\partial \rho} j_{\nu} j^{\nu} + \frac{\partial \alpha_{TV}(\rho)}{\partial \rho} \vec{j}_{\nu} \vec{j}^{\nu} \right]. \quad (33)$$

RPA matrix elements can be easily obtained from:

$$V_{abcd} = \frac{\delta h_{ac}}{\delta \rho_{db}} = \frac{\delta^2 E[\rho]}{\delta \rho_{ac} \delta \rho_{bd}}, \quad (34)$$

where

$$h = -i\alpha \nabla + \beta(m + \Sigma) + \beta\gamma_{\mu} \Sigma^{\mu} \quad (35)$$

## General particle-hole matrix elements

$$V_{abcd}^{(1)} = \int d^3 r_1 \int d^3 r_2 \sum_j \bar{\psi}_a(\vec{r}_1) \Gamma_{j\mu}^{(1)} O_\tau^{(1)} \psi_c(\vec{r}_1) D_j(\vec{r}_1, \vec{r}_2) \\ \times \bar{\psi}_b(\vec{r}_2) \Gamma_j^{(2)\mu} O_\tau^{(2)} \psi_d(\vec{r}_2) \quad (36)$$

$$V_{abcd}^{(2)} = \int d^3 r_1 \int d^3 r_2 \sum_j \bar{\psi}_a(\vec{r}_1) \Gamma_{j\mu}^{(1)} O_\tau^{(1)} \psi_c(\vec{r}_1) \frac{\partial D_j(\vec{r}_1, \vec{r}_2)}{\partial \rho(\vec{r}_1)} \\ \times \rho_j^\mu(\vec{r}_2) \bar{\psi}_b(\vec{r}_2) 1^{(2)} \psi_d(\vec{r}_2) \quad (37)$$

$$V_{abcd}^{(3)} = 1/2 \int d^3 r_1 \int d^3 r_2 \sum_j \bar{\psi}_a(\vec{r}_1) 1^{(1)} \psi_c(\vec{r}_2) \frac{\partial D_j^2(\vec{r}_1, \vec{r}_2)}{\partial \rho^2(\vec{r}_1)} \\ \times \rho_{j\mu}(\vec{r}_1) \rho_j^\mu(\vec{r}_2) \bar{\psi}_b(\vec{r}_2) 1^{(2)} \psi_d(\vec{r}_2) \quad (38)$$

$$V_{abcd}^{(4)} = \int d^3 r_1 \int d^3 r_2 \sum_j \bar{\psi}_a(\vec{r}_1) 1^{(1)} \psi_c(\vec{r}_2) \frac{\partial D_j(\vec{r}_1, \vec{r}_2)}{\partial \rho(\vec{r}_1)} \\ \times \rho_{j\mu}(\vec{r}_1) \bar{\psi}_b(\vec{r}_1) \Gamma_j^{(2)\mu} O_\tau^{(2)} \psi_d(\vec{r}_2), \quad (39)$$

## Separable pairing matrix elements

In 2009 Y. Tian, Z. Y. Ma and P. Ring introduced separable pairing in  $^1S_0$  channel gap equation of symmetric nuclear matter:

$$\Delta(k) = - \int_0^\infty \frac{k'^2 dk'}{2\pi^2} \langle k | V_{sep}^{1S_0} | k' \rangle \frac{\Delta(k')}{2E(k')}, \quad (40)$$

where

$$\langle k | V_{sep}^{1S_0} | k' \rangle = - G p(k) p(k'), \quad (41)$$

with Gaussian ansatz

$$p(k) = e^{-a^2 k^2}. \quad (42)$$

These two parameters  $a$  and  $G$  were fitted to density dependence of the gap at the Fermi surface in nuclear matter. After transformation the separable force from momentum to coordinate space we obtained:

$$V(\vec{r}_1, \vec{r}_2, \vec{r}'_1, \vec{r}'_2) = - G \delta(\vec{R}_1 - \vec{R}_2) P(r) P(r') \frac{1 - \hat{P}^\sigma}{2} \quad (43)$$

with  $\vec{r} = 1/\sqrt{2}(\vec{r}_1 - \vec{r}_2)$  and  $\vec{R} = 1/\sqrt{2}(\vec{r}_1 + \vec{r}_2)$ . Values of  $G = 728 \text{ MeV fm}^3$  and  $a_0 = 0.644 \text{ fm}$

## Separable pairing matrix elements

After Fourier transformation:

$$P(r) = \frac{e^{-r^2/(2a^2)}}{(4\pi a^2)^{3/2}} \quad (44)$$

Due to coordinate transformation from laboratory to center of mass and relative coordinates we need to use Talmi-Moschinsky brackets:

$$|n_1 l_1, n_2 l_2; \lambda \mu\rangle = \sum_{NLnl} M_{n_1 l_1 n_2 l_2}^{NLnl} |NL, nl; \lambda \mu\rangle \quad (45)$$

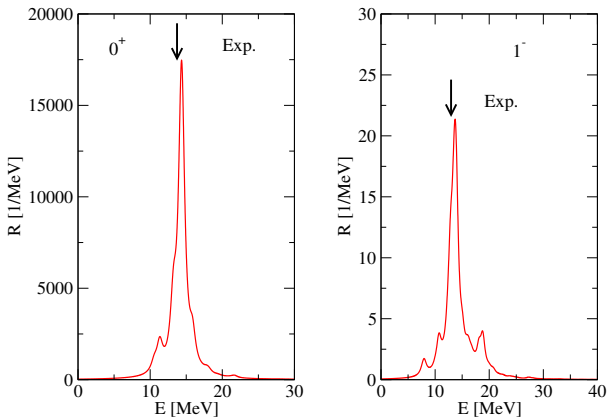
If we work in the basis of spherical harmonic oscillator:

$$I_n = \sqrt{4\pi} \int R_{nl}(r) P(r) r^2 dr = \frac{1}{2^{2/3} \pi^{3/4} b^{3/2}} \frac{(1 - \alpha^2)^n}{(1 + \alpha^2)^{n+3/2}} \frac{\sqrt{2n+1}}{2^n n!}, \quad (46)$$

the coupled matrix element for  $T = 1$  pairing:

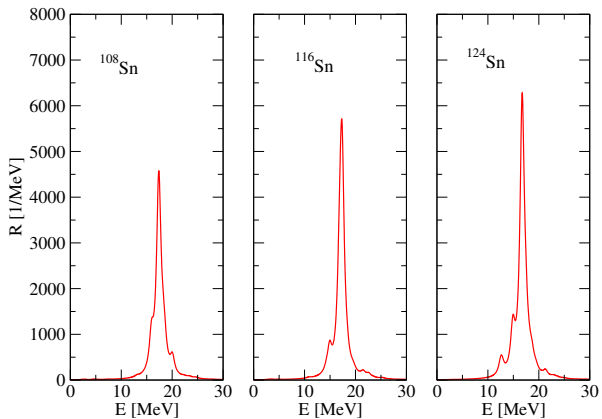
$$V_{abcd}^{JM} = G \hat{j}_a \hat{j}_b \hat{j}_c \hat{j}_d (-1)^{l_b + l_d + j_a + j_c} \left\{ \begin{array}{ccc} l_a & j_a & 1/2 \\ j_b & l_b & J \end{array} \right\} \left\{ \begin{array}{ccc} l_c & j_c & 1/2 \\ j_d & l_d & J \end{array} \right\} \sum_{Nnn'} I_n I_{n'} M_{n_a l_a n_b l_b}^{NJn'0} M_{n_c l_c n_d l_d}^{NJn'0} \quad (47)$$

## Collective excitations



**Figure 2:** The Isoscalar monopole (left) and isovector dipole (right) strength distribution for  $^{208}\text{Pb}$ . ISGMR centroid energy is on 14.2 (exp.  $13.96 \pm 0.20$  MeV, while for RPA IVGDR predicts the excitation energy on 13.7 (exp. 13.3 MeV).

## Isoscalar monopole collective excitation



**Figure 3:** The Isoscalar monopole strength distribution for even-even nuclei  $A = 108 - 124$  with  $\Delta A = 8$ .

## Proton-neutron (R)QRPA

RPA equations:

$$\begin{pmatrix} A & B \\ B^* & A^* \end{pmatrix} \begin{pmatrix} X^\lambda \\ Y^\lambda \end{pmatrix} = E_\lambda \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \quad (48)$$

$$\begin{aligned} A_{pn,p'n'} &= H_{pp'}^{11} \delta_{nn'} + H_{nn'}^{11} \delta_{pp'} + (u_p v_n u_{p'} v_{n'} + v_p u_n v_{p'} u_{n'}) \\ &\quad \times V_{pn'np'}^{ph} + (u_p u_n u_{p'} u_{n'} + v_p v_n v_{p'} v_{n'}) V_{pnp'n'}^{pp}, \quad (49) \end{aligned}$$

$$\begin{aligned} B_{pn,p'n'} &= (u_p v_n v_{p'} u_{n'} + v_p u_n u_{p'} v_{n'}) V_{pp'nn'}^{ph} \\ &\quad + (u_p u_n v_{p'} v_{n'} + v_p v_n u_{p'} u_{n'}) V_{pnp'n'}^{pp}. \quad (50) \end{aligned}$$

Non-diagonal  $H^{11}$  is defined as

$$H^{11} = (u_i u_j + v_i v_j) h_{ij} + (u_i v_j - v_i u_j) \Delta_{ij} \quad (51)$$



## Proton-neutron (R)QRPA

Isovector-vector part of PN-(R)QRPA survives only for  $V^{(1)}$  type of matrix elements. In the case of spacelike components we have:

$$V_{abcd}^{(1s)} = - \int d^3 r_1 \int d^3 r_2 \psi_a^\dagger(\vec{r}_1) (\vec{\tau} \gamma_0 \gamma_\mu)^{(1)} \psi_c(\vec{r}_1) \\ \times \alpha_{TV}(\rho) \delta(\vec{r}_1 - \vec{r}_2) \psi_b^\dagger(\vec{r}_2) (\vec{\tau} \gamma_0 \gamma^\mu)^{(2)} \psi_d(\vec{r}_2), \quad (52)$$

and timelike components:

$$V_{abcd}^{(1t)} = \int d^3 r_1 \int d^3 r_2 \psi_a^\dagger(\vec{r}_1) \vec{\tau}^{(1)} \psi_c(\vec{r}_1) \alpha_{TV}(\rho) \\ \times \delta(\vec{r}_1 - \vec{r}_2) \psi_b^\dagger(\vec{r}_2) \vec{\tau}^{(2)} \psi_d(\vec{r}_2). \quad (53)$$

## Isvector-vector matrix elements

If the radial part of wave function is given by:

$$\psi(r) = \begin{pmatrix} f(r) \\ ig(r) \end{pmatrix} \quad (54)$$

by coupling of angular momentum, we get coupled matrix elements for spacelike part:

$$\begin{aligned} V_{abcd}^{(1s)J} = & 2\hat{J}^{-2} \sum_L \int dr r^2 \alpha_{TV}(f_a(r)g_c(r) \langle j_a || (\sigma_S Y_L)_J || \bar{j}_c \rangle \\ & - g_a(r)f_c(r) \langle \bar{j}_a || (\sigma_S Y_L)_J || j_c \rangle) (f_b(r)g_d(r) \langle \bar{j}_d || (\sigma_S Y_L)_J || j_b \rangle \\ & - g_b(r)f_d(r) \langle j_d || (\sigma_S Y_L)_J || \bar{j}_b \rangle), \quad (55) \end{aligned}$$

and timelike part

$$\begin{aligned} V_{abcd}^{(1t)J} = & 2J^{-2} \int dr r^2 \alpha_{TV}(f_a(r)f_c(r) + g_a(r)g_c(r)) \\ & (f_b(r)f_d(r) + g_b(r)g_d(r)) \langle j_a || Y_J || j_c \rangle \langle j_d || Y_J || j_b \rangle, \quad (56) \end{aligned}$$

## Isvector-vector matrix elements

Angular part is given by expressions:

$$\langle j_a || Y_J || j_c \rangle = \frac{1 + (-1)^{l_a + l_c + J}}{2} \frac{\hat{j}_a \hat{j}_c \hat{J}}{\sqrt{4\pi}} (-1)^{j_a - 1/2} \begin{pmatrix} j_a & J & j_c \\ -1/2 & 0 & 1/2 \end{pmatrix}, \quad (57)$$

and

$$\begin{aligned} \langle j_a || (\sigma_S Y_L)_J || j_c \rangle &= \frac{1 + (-1)^{l_a + l_c + J}}{2} \frac{\hat{j}_a \hat{j}_c \hat{J} \hat{L}}{\sqrt{4\pi}} (-1)^{l_a + L} \\ &\times \left[ (-1)^{l_c + j_c + 1/2} \begin{pmatrix} 1 & L & J \\ 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} j_a & J & j_c \\ -1/2 & 0 & 1/2 \end{pmatrix} - \right. \\ &\quad \left. \sqrt{2} \begin{pmatrix} 1 & L & J \\ -1 & 0 & 1 \end{pmatrix} \begin{pmatrix} j_a & J & j_c \\ 1/2 & -1 & 1/2 \end{pmatrix} \right]. \quad (58) \end{aligned}$$

## Isovector-pseudovector matrix elements

In order to describe the effects of pions in multipolar transitions in PN-RQRPA we use isovector-pseudovector coupling. However, we expect that the strength of pion-nucleon in nuclei coupling should be somewhat reduced by factor of  $g'$ :

$$V_{PV} = -g' \left( \frac{f_\pi}{m_\pi} \right)^2 \delta(\vec{r}_1 - \vec{r}_2) (\gamma_0 \gamma_5 \gamma_\mu \vec{\tau})^{(1)} (\gamma_0 \gamma_5 \gamma^\mu \vec{\tau})^{(2)}, \quad (59)$$

where  $f_\pi$  decay constant and  $m_\pi$  mass of pion. The value of  $g'$  we don't know *a priori* and should be deduced from experiment.

For timelike part pseudovector matrix elements look like:

$$V_{abcd}^{PV(t)J} = 2g' \hat{J}^{-2} \left( \frac{f_\pi}{m_\pi} \right)^2 \int dr r^2 (f_a(r) g_c(r) - g_a(r) f_c(r)) \\ \times (f_b(r) g_d(r) - g_b(r) f_d(r)) \langle j_a || Y_J || j_c \rangle \langle j_d || Y_J || j_c \rangle, \quad (60)$$

**Timelike matrix elements are non-zero only in case of unnatural parity transitions, like Gamow-Teller transition.**

## Isovector-pseudovector matrix elements

For spacelike part isovector-pseudovector matrix elements look like:

$$\begin{aligned}
 V_{abcd}^{PV(s)J} = & 2g' \hat{J}^{-2} \left( \frac{f_\pi}{m_\pi} \right)^2 \sum_L \int dr r^2 (f_a(r) f_c(r)) \\
 & \times \langle j_a || (\sigma_S Y_L)_J || j_c \rangle + g_a(r) g_c(r) \langle \bar{j}_a || (\sigma_S Y_L)_J || \bar{j}_c \rangle \langle f_b(r) f_d(r) \\
 & \times \langle j_d || (\sigma_S Y_L)_J || j_b \rangle + g_b(r) g_d(r) \langle \bar{j}_d || (\sigma_S Y_L)_J || \bar{j}_b \rangle. \quad (61)
 \end{aligned}$$

**Spacelike isovector-pseudovector matrix elements contribute to both unnatural and natural parity transitions, but in opposite manner compared to isovector-vector matrix elements.**

## Separable pairing in the case of PN-(R)QRPA

Coupled pp-elements for PN-(R)QRPA are given by:

$$V_{abcd}^{JM} = -G \hat{j}_a \hat{j}_b \hat{j}_c \hat{j}_d \sum_{LS} \left(1 - (-1)^{S+T}\right) \hat{S}^2 \hat{L}^2 \left\{ \begin{array}{ccc} l_b & 1/2 & j_b \\ l_a & 1/2 & j_a \\ L & S & J \end{array} \right\} \left\{ \begin{array}{ccc} l_d & 1/2 & j_d \\ l_c & 1/2 & j_c \\ L & S & J \end{array} \right\} \sum_{nn'} I_n I_{n'} M_{n_a l_a n_b l_b}^{NLn0} M_{n_c l_c n_d l_d}^{NLn'0} \quad (62)$$

**This is non-vanishing only for  $S = 0$  and  $T = 1$  or  $S = 1$  and  $T = 0$ .** Assumed form of pp-matrix elements:

$$V_{abcd}^{JM} = \langle ab|V(T=0; S=1)|cd\rangle^{JM} \times V_0 + \langle ab|V(T=1; S=0)|cd\rangle^{JM} \quad (63)$$

We don't know *a priori* the value of factor  $V_0$ . It can be somewhat reduced or enhanced compared to the case  $T = 1$ .

## Determination of parameter $g'$

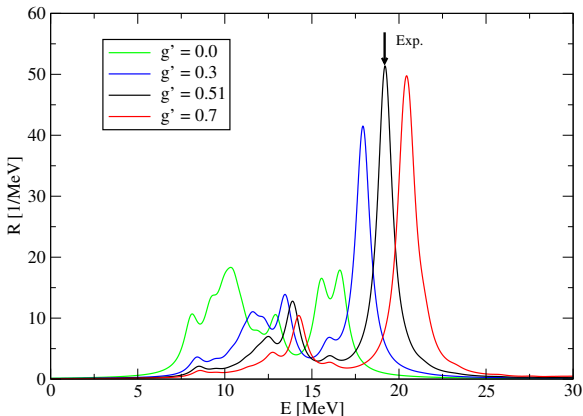
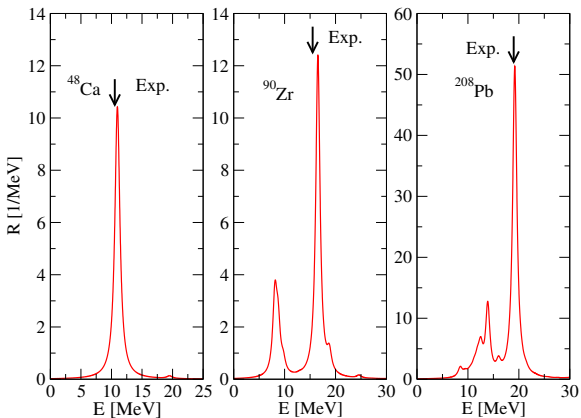


Figure 4: The Gammow-Teller strength distribution for  $^{208}\text{Pb}$  for different values of parameter  $g'$ , calculated with DD-PC1 interaction. The experimental value of the position of  $\text{GT}^-$  main peak in  $^{208}\text{Pb}$  is on 19.2 MeV. This corresponds to  $g' = 0.51$ . The width of Lorentzian is 1 MeV.

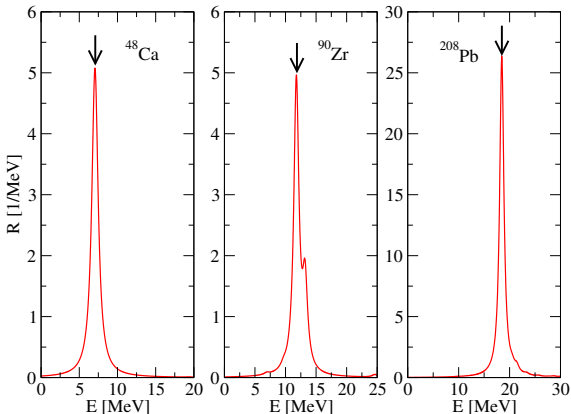
## Gammow-Teller resonance



**Figure 5:** The Gammow-Teller strength distribution for  $^{48}\text{Ca}$ ,  $^{90}\text{Zr}$  and  $^{208}\text{Pb}$ , calculated by DD-PC1 interaction. The theoretical (experimental) values of central position GTR main peak are 11.0 (10.5) MeV for  $^{48}\text{Ca}$  and 16.6 (15.6) MeV for  $^{90}\text{Zr}$ .



## Isobaric analog resonance



**Figure 6:** The isobar analog state strength distribution for  $^{48}\text{Ca}$ ,  $^{90}\text{Zr}$  and  $^{208}\text{Pb}$ , calculated by DD-PC1. The theoretical (experimental) values of central position GTR main peak are 7.05 (7.17) MeV for  $^{48}\text{Ca}$  and 11.7 ( $12.0 \pm 0.9$ ) MeV for  $^{90}\text{Zr}$  and 18.47 ( $18.83 \pm 0.02$ ) MeV for  $^{208}\text{Pb}$ .

## Isobaric analog resonance in Sn isotopes

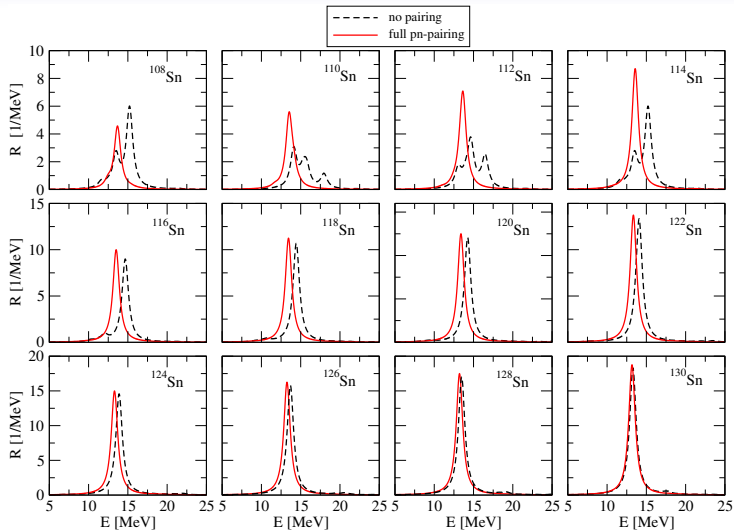


Figure 7: PN-(R)QRPA strength distribution for even-even nuclei with  $A = 108 - 130$ , calculated by DD-PC1 with separable pairing interaction.

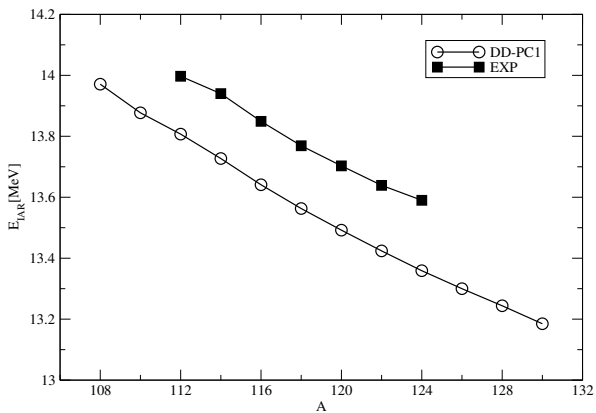


Figure 8: Calculated and experimental IAR position for Sn isotopes.

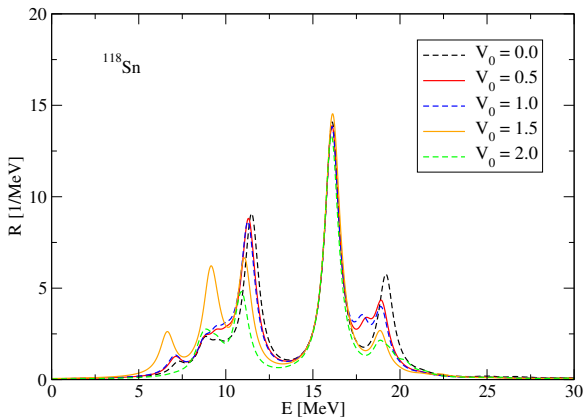
Gamow-Teller resonance in  $^{118}\text{Sn}$ 

Figure 9: The Gamow-Teller strength distribution for  $^{118}\text{Sn}$  for different values of parameter  $V_0$ . Notice, only  $T = 0$  pp matrix elements are nonvanishing.

## Standard model

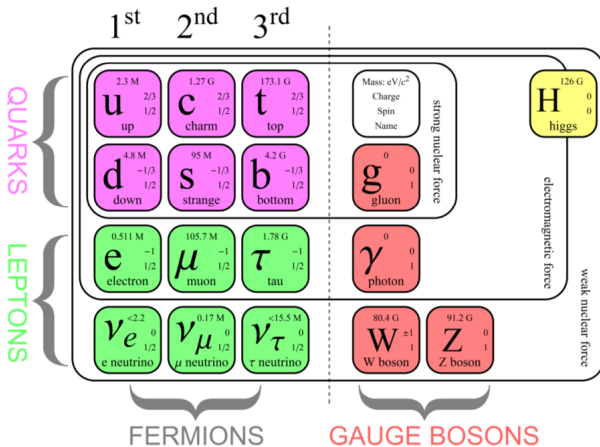


Figure 10: Standard model.

## Why neutrino-nucleus reactions?

Neutrino mass hierarchy still remains an open problem due to insensitivity of previous and nowadays experiments on sign of difference between squares of neutrino masses.

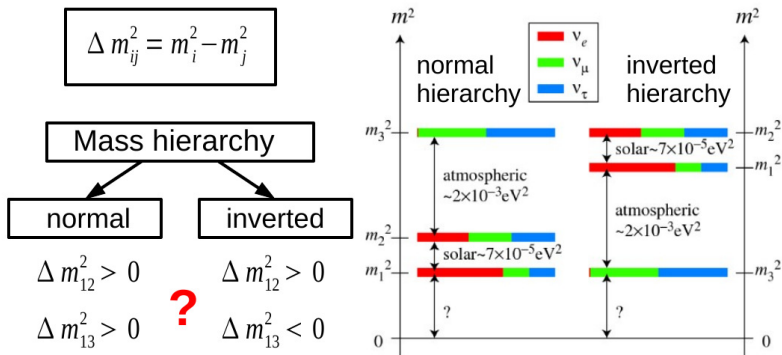


Figure 11: Possible neutrino mass ordering (hierarchy). Degenerate case excluded from figure.

## Why neutrino-nucleus reactions?

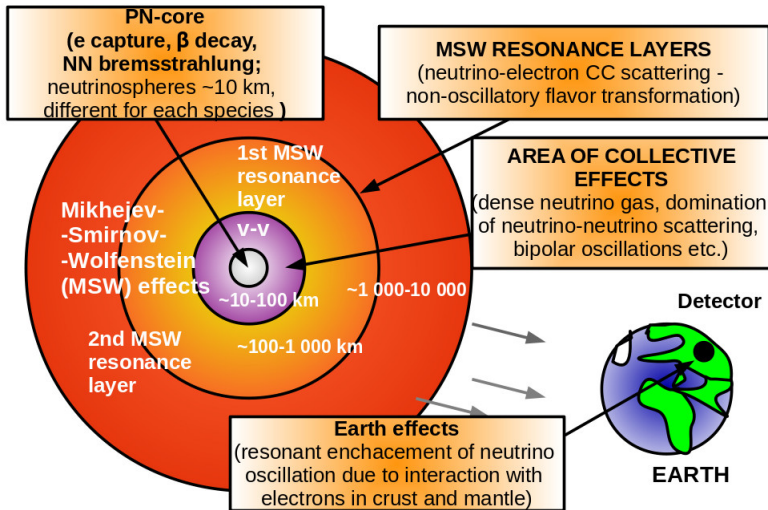


Figure 12: Supernova from neutrino perspective.

## Why neutrino-nucleus reactions?

We need to choose adequate target material for detector in order to achieve good statistics and include or exclude particular neutrino species in a specific energy range.

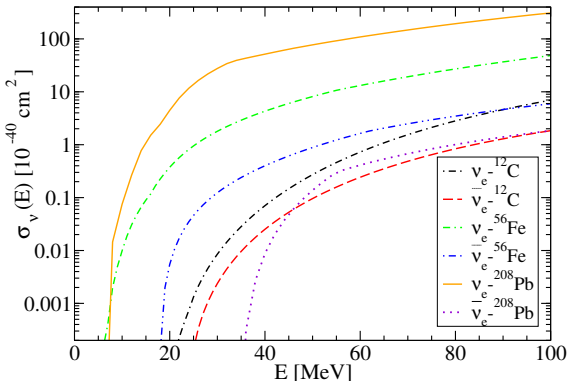


Figure 13: The inclusive  $\nu_e(\bar{\nu}_e)$  cross sections with  $^{12}\text{C}$ ,  $^{56}\text{Fe}$  and  $^{208}\text{Pb}$  as a function of  $\nu_e(\bar{\nu}_e)$  energy (charged current reactions only). (see refs. D. Vale, N. Paar AIP Conference Proceedings 1681 (1), 050011; N. Paar, T. Marketin, D. Vale, D. Vretenar International Journal of Modern Physics E 24 (09), 1541004.



## Why neutrino-nucleus reactions?

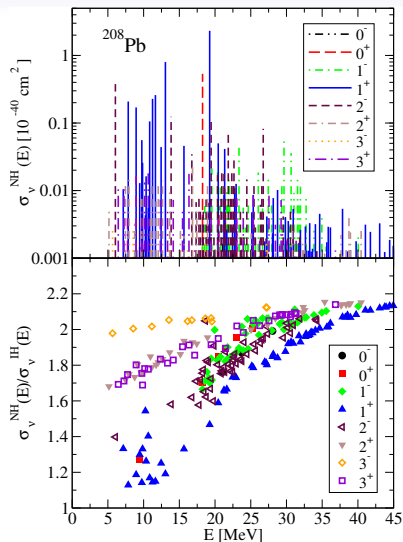


Figure 14: Example of multipole decomposition of the flux averaged cross sections for  $\nu_e - {}^{208}\text{Pb}$  charged current reaction. see ref. D Vale, T Rauscher, N Paar *Journal of Cosmology and Astroparticle Physics* 2016 (02), 007

## Why neutrino-nucleus reactions?

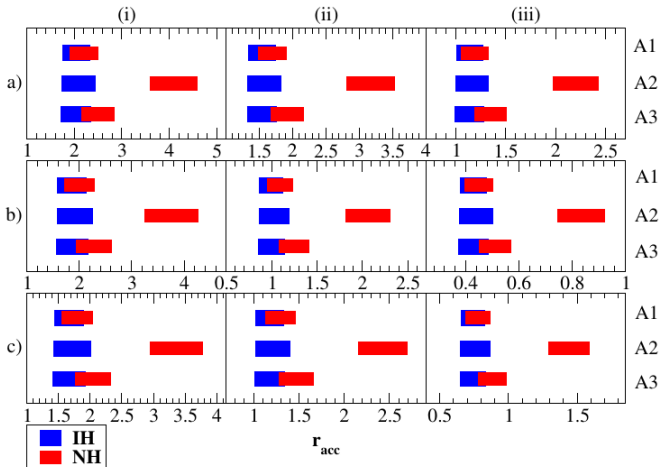


FIG. 6: The ratios  $r_{acc}$  of the number of the detector events induced in mineral oil ( $\text{CH}_2$ ), water ( $\text{H}_2\text{O}$ ), and  $^{208}\text{Pb}$  for the incoming (anti)neutrino fluxes of the accretion phase in normal (NH) and inverted (IH) mass hierarchy: (a)  $r_{e^+, \text{Pb}}^{\text{free p, Pb}} [10^{-1}]$  (b)  $r_{e^+, 2n}^{\text{free p, Pb}}$  (c)  $r_{e^+, \text{tot n}}^{\text{free p, Pb}} [10^{-1}]$ . The three columns correspond to (i)  $\langle E_{\nu_e}^0 \rangle = 8$  MeV,  $\langle E_{\bar{\nu}_e}^0 \rangle = 11$  MeV,  $\langle E_{\nu_x}^0 \rangle = 13$  MeV; (ii)  $\langle E_{\nu_e}^0 \rangle = 10$  MeV,  $\langle E_{\bar{\nu}_e}^0 \rangle = 13$  MeV,  $\langle E_{\nu_x}^0 \rangle = 15$  MeV; and (iii)  $\langle E_{\nu_e}^0 \rangle = 12$  MeV,  $\langle E_{\bar{\nu}_e}^0 \rangle = 15$  MeV,  $\langle E_{\nu_x}^0 \rangle = 19$  MeV. A1, A2, A3 denote the combinations of the luminosity ratios for the accretion phase given in Tab. [I](#).

## Why neutrino-nucleus reactions?

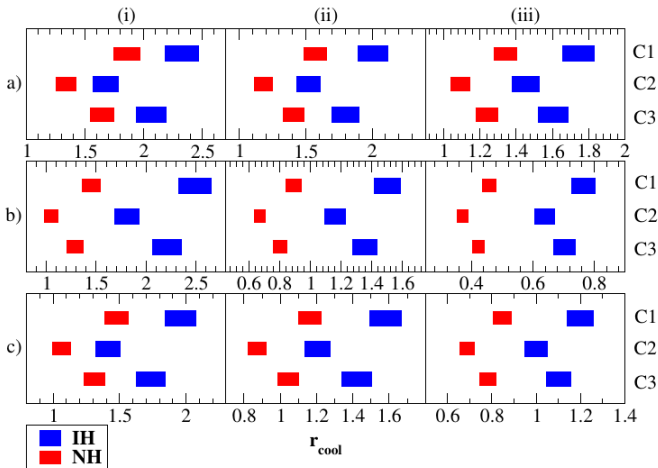


FIG. 7: The ratios  $r_{cool}$  of the number of the detector events induced in mineral oil ( $\text{CH}_2$ ), water ( $\text{H}_2\text{O}$ ), and  $^{208}\text{Pb}$  for the incoming (anti)neutrino fluxes of the cooling phase in normal (NH) and inverted (IH) mass hierarchy for three configurations of initial  $\nu$  average energies. The same notation as in Fig. 6 applies. C1, C2, C3 denote the combinations of the luminosity ratios for the cooling phase given in Tab. I.