Quantum Matter and Quantum Information with Holography August 23 (Sun), 2020 ~ August 31 (Mon), 2020

Quantum chaos in topologically massive gravity

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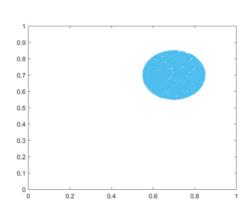
based on: 2005.08508 (with Avinash Raju)



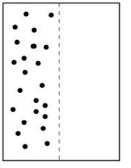
Classical Chaos

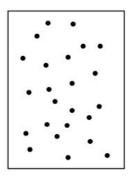
Early time: exponential sensitivity of phase-space trajectories to the initial conditions

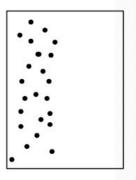
Long time: mixing



Very late time: Poincare recurrences





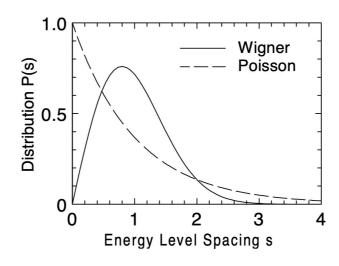


Classical Chaos vs Quantum Mechanics

- In quantum systems, the 'classical' coarse-graining is set by \hbar
- The picture above might not be useful (after Ehrenfest time scale) $t_{\rm Ehrenfest} \sim \frac{1}{\lambda} \log \left(\int_{\Sigma} p dq/\hbar \right)$
- How the previous discussion should be modified due to QM
- Classical chaos for nonzero ħ? Quantum chaos?

Quantum Chaos

Quantize the classical chaotic system: chaotic systems have statistical level repulsion characteristic of random matrices [Review by D'Alessio, Kafri, Polkovikov, Rigol, 1509.06411]



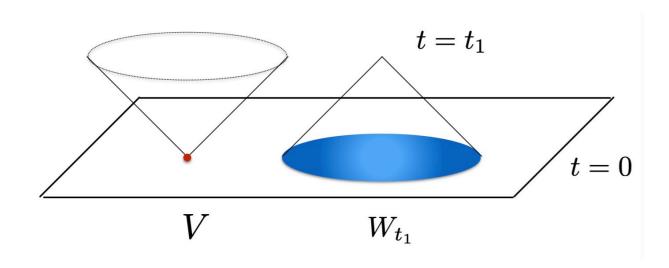
Local chaotic behavior can be generalized on the short time scale (semi-classical intuition)

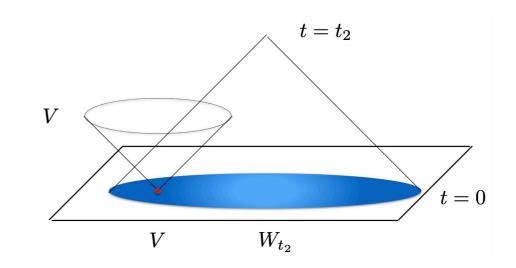
$$\frac{\partial q(t, p_0)}{\partial p_0} = \{q(t), p_0\}_{PB} \qquad \langle [V, W_t]^2 \rangle \sim e^{\lambda t}$$

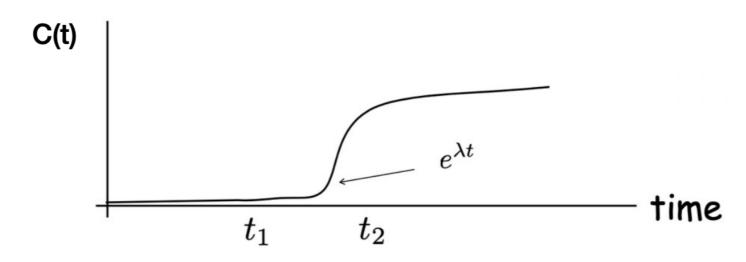
Expectation value of the commutators

[Larkin, Ovchinnikov, JETP (1969); Shenker, Stanford, 1306.0622; Maldcena, Shenker, Stanford, 1503.01409]

$$C(t, \mathbf{x}) = -\langle [W(t, \mathbf{x}), V(0)]^{\dagger} [W(t, \mathbf{x}), V(0)] \rangle_{\beta}$$







Out-of-Time-Ordered Correlators (OTOC)

[Larkin, Ovchinnikov, JETP (1969); Shenker, Stanford, 1306.0622; Maldcena, Shenker, Stanford, 1503.01409]

$$C(t, \mathbf{x}) = -\langle [W(t, \mathbf{x}), V(0)]^{\dagger} [W(t, \mathbf{x}), V(0)] \rangle_{\beta}$$

Probing the sensitivity of a system to the initial conditions

$$C_2 = C_1 - C(t, \mathbf{x})$$

$$C_2 = \langle \Psi_1(t) | \Psi_2(t) \rangle + \langle \Psi_2(t) | \Psi_1(t) \rangle \longrightarrow \text{Out-of-time Ordered}$$

$$C_1 = \langle \Psi_1(t) | \Psi_1(t) \rangle + \langle \Psi_2(t) | \Psi_2(t) \rangle \longrightarrow \text{Time ordered}$$

$$|\Psi_1(t)\rangle = V_0 W_t | \text{TFD}\rangle$$

$$|\Psi_2(t)\rangle = W_t V_0 | \text{TFD}\rangle$$

Lyapunov exponents, butterfly velocities: for interacting quantum systems with many degrees of freedom

$$C_2 = 1 - \epsilon e^{\lambda_L \left(t - \frac{x}{v_B} \right)} \qquad t_r \ll t \ll t_*$$

Different (diffusive) spreading might be seen in non-maximally chaotic systems

Holographic OTOC

[Shenker, Stanford, 1306.0622; 1412.6087]

- OTOC = amplitudes for 2-to-2 scatterings of particles dual to W and V in a black hole geometry dual to the thermal state $|{
 m TFD}\rangle$
- In elastic eikonal gravity approximation, the dominate contribution is related to the gravitational shock waves on the horizon of a two-sided black hole
- lacktriangle Universal Lyapunov exponent $\lambda_L=2\pi T$
- The butterfly velocity depends on the details of the black hole geometry

Chaos bound [Maldacena, Shenker, Stanford, 1503.01409]

Related regulated function

$$F(t) = \text{Tr}(yV(0)yW(t)yV(0)yW(t)) \sim 1 - \epsilon e^{\lambda_L t}, \quad y^4 = \frac{e^{-\beta H}}{Z}$$

For systems with large hierarchy between thermalization and scrambling, analyticity in correlation functions demands

$$\lambda_L \le 2\pi T$$

- It holds for very generic quantum many-body systems
- Black holes saturate this bound: maximal chaos
- SYK/AdS₂ [Kitaev, 2015]

Chaos from hydrodynamics via pole skipping

[Grozdanov, Schalm, Scopelliti, 1710.00921; Blake, Lee, Liu, 1801.00010; Blake, Davison, Grozdanov, Liu, 1809.01169;...]

- Naively hydrodynamics has nothing to do with chaos
- **Deep connection from EFT:** Signatures of chaos in energy density two point function of $G_{T^{00}T^{00}}^{R}(\omega,k)$
- There exists a special point $(\omega_*, k_*) = (i\lambda_L, \frac{i\lambda_L}{v_B})$ in $G^R(\omega, k) = \frac{B(\omega, k)}{A(\omega, k)}$ with $A(\omega_*, k_*) = B(\omega_*, k_*) = 0$
- Examples of pole skipping in many maximally chaotic systems: SYK, AdS black holes in Einstein gravity plus matter
- Pole skipping also exists for 2-pt correlators of other operators on the lower half plane

Motivation

- Connection between OTOC and pole skipping, e.g. for systems with multiple Lyapunov exponents or nonmaximally chaotic systems
- What is the role of rotation in holographic chaos
- What is the role of massive graviton in holographic chaos

Why 3D gravity

- A "simple" toy model to understand quantum gravity
- We can learn much from CFT calculations
- In the following, we will talk about
 - Quantum chaos in 3D Einstein gravity
 - Quantum chaos in 3D TMG

3D Einstein gravity

Einstein-Hilbert action

$$S_{\rm EH} = \frac{1}{16\pi G} \int_{\mathcal{M}} d^3x \sqrt{-g} \left(R - 2\Lambda \right)$$

BTZ Black hole solution

$$ds^{2} = -f(r)dt^{2} + \frac{dr^{2}}{f(r)} + r^{2} \left(d\varphi - \frac{r_{+}r_{-}}{\ell r^{2}} dt \right)^{2},$$

$$f(r) = \frac{(r^{2} - r_{+}^{2})(r^{2} - r_{-}^{2})}{\ell^{2}r^{2}}.$$

- $^{\bullet}$ M, T, Ω are determined by r_+, r_- .
- The dual theory is expected to be a CFT with $\beta_{\pm}=\beta(1\mp\ell\Omega)$ and $c_{+}=c_{-}=\frac{3\ell}{2G}$
- The angular direction is periodic. At high temperature $\frac{\beta}{\ell} \to 0$, we can take a "decompactification" limit (a boosted brane)

Chaos parameters from OTOC

[Jahnke, Kim, Yoon, 1903.09086; (Stikonas 2018; Poojary 2018)]

From shock wave calculations

$$OTOC(t, \varphi_{12}) \simeq 1 + \epsilon e^{\frac{2\pi}{\beta}t} h(\Omega t - \varphi) \simeq 1 + C_1 e^{\frac{2\pi}{\beta_+}(t + \ell \varphi_{12})} + C_2 e^{\frac{2\pi}{\beta_-}(t - \ell \varphi_{12})}$$

Naively we have

$$\lambda_{\pm} = \frac{2\pi}{\beta(1 \mp \Omega\ell)} \qquad v_{\pm} = \mp 1$$

- The chaos bound is violated: $\lambda_- < \frac{2\pi}{\beta} < \lambda_+$
- However, the angular coordinate is periodic, i.e. the profile of shock wave is periodic, therefore the two coefficients C1 and C2 are not independent [Mezei, Sarosi, 1908.03574]

$$OTOC(t, \varphi_{12}) \simeq 1 + \epsilon \left[e^{\frac{2\pi}{\beta_+}(t + \ell \varphi_{12})} + \# e^{\frac{2\pi}{\beta_-}(t - \ell \varphi_{12})} \right]$$

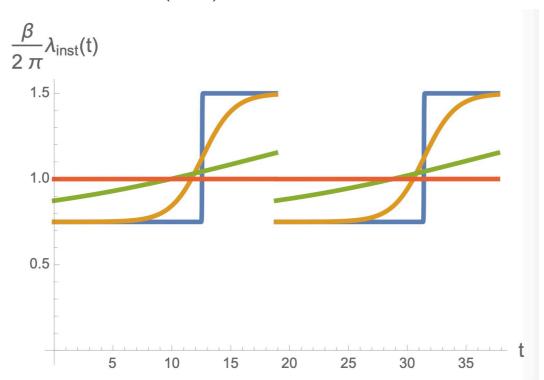
Instantaneous Lyapunov exponent

[Mezei, Sarosi, 1908.03574]

$$OTOC(t, \varphi_{12}) \simeq 1 + \epsilon \left[e^{\frac{2\pi}{\beta_+}(t + \ell \varphi_{12})} + \# e^{\frac{2\pi}{\beta_-}(t - \ell \varphi_{12})} \right]$$

instantaneous Lyapunov exponent

$$OTOC(t, 0) \simeq 1 + \epsilon e^{\lambda_{inst.}t}$$



$$\beta = 0, 2\pi, 16\pi, \infty$$

- In the high temperature limit, the instantaneous Lyapunov exponents behave as step function.
- lacktriangledaw The average of instantaneous Lyapunov exponents is $rac{2\pi}{eta}$

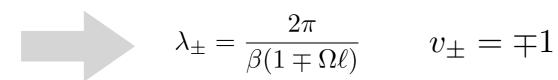
Pole skipping in holography

• From EOM near the horizon $E_{vv} = 0$

Expand
$$h_{ab}=e^{-i\omega v+ik\phi}(r-r_+)^{\gamma}\sum_{n=0}^{\infty}\tilde{h}_{ab}^{(n)}(r-r_+)^n$$
 near horizon,

$$(2\pi i\omega + 4\pi i\Omega k - k^2\beta(1-\Omega^2))\tilde{h}_{vv}^{(0)} = -(2\pi i - \beta\omega)(1-\Omega^2) \left[2k\tilde{h}_{v\phi}^{(0)} + \omega\tilde{h}_{\phi\phi}^{(0)}\right].$$

At
$$(\omega, k) = \left(\frac{2\pi i}{\beta(1 \mp \Omega)}, \pm \frac{2\pi i}{\beta(1 \mp \Omega)}\right)$$
 both solutions are regular



Correlators of energy density from holography

$$\langle T^{\tau\tau}(\omega_E, k) T^{\tau\tau}(-\omega_E, -k) \rangle \propto \frac{\delta^2 S_{\text{ren.}}}{\delta \tilde{h}_{\tau\tau}^{(0)} \delta \tilde{h}_{\tau\tau}^{*(0)}} = \frac{k^2 (4 + k^2)}{2(\omega_E^2 + k^2)}$$

The pole skipping point is

$$(\omega, k) = \left(\frac{2\pi i}{\beta(1 \mp \Omega)}, \pm \frac{2\pi i}{\beta(1 \mp \Omega)}\right)$$

Pole skipping in CFT

For CFT on cylinder

$$G_R(\omega, k) = \frac{c_L}{6} \left(\frac{2\pi}{\beta_L}\right)^3 \left(\frac{2i}{\omega - k} + \pi\delta\left(\frac{k - \omega}{2}\right)\right) \sinh\left[\frac{\beta_L k}{2}\right] \left|\Gamma\left(2 + \frac{i\beta_L}{2\pi}k\right)\right|^2 - \frac{c_R}{6} \left(\frac{2\pi}{\beta_R}\right)^3 \left(\frac{2i}{\omega + k} + \pi\delta\left(\frac{k + \omega}{2}\right)\right) \sinh\left[\frac{\beta_R k}{2}\right] \left|\Gamma\left(2 + \frac{i\beta_R}{2\pi}k\right)\right|^2.$$

- From the first term $(\omega, k) = \left(\pm \frac{2\pi i}{\beta(1-\Omega)}, \pm \frac{2\pi i}{\beta(1-\Omega)}\right)$
- From the second term $(\omega, k) = \left(\pm \frac{2\pi i}{\beta(1+\Omega)}, \ \mp \frac{2\pi i}{\beta(1+\Omega)}\right)$

[see also Haehl, Rozali, 1808.02898]

Pole skipping is a generic feature of any CFT, including chaotic CFTs and non-chaotic CFTs.

Topologically Massive Gravity (TMG)

A gravitational Chern-Simons deformation to Einstein gravity [Deser, Jackiw, Templeton, 1988]

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{-g} \left(R + 2 + \frac{1}{2\mu} \varepsilon^{abc} \Gamma^d_{ae} \left(\partial_b \Gamma^e_{cd} + \frac{2}{3} \Gamma^e_{bf} \Gamma^f_{cd} \right) \right)$$

- Any solution of Einstein gravity is a solution of TMG
- Thermodynamics for rotating BTZ black holes [Krause, Larsen, hep-th/0508218]

$$M(\mu) = M + \frac{J}{\mu}, \quad J(\mu) = J + \frac{M}{\mu}$$

- The angular direction and the decompactification limit
- * The dual field theory for TMG on rotating BTZ is a CFT with $~eta_{\pm}=eta(1\mp\ell\Omega)$

and
$$(c_L, c_R) = \frac{3\ell}{2G} \left(1 - \frac{1}{\mu}, 1 + \frac{1}{\mu} \right)$$

- When $\mu\ell < 1$: negative central charge; Black hole instability [Park, hep-th/0608165]
- ightharpoonup Chiral point $\mu\ell=1$ [Li, Song, Stronminger, 0801.4566]

Chaos in TMG from OTOC

 ** Profile for shock wave and OTOC ($\mu \neq 1$)

$$h(\phi) = c_1 e^{-\frac{2\pi\phi}{\beta(1+\Omega)}} + c_2 e^{\frac{2\pi\phi}{\beta(1-\Omega)}} + c_3 e^{\frac{2\pi(\Omega-\mu)\phi}{\beta(1-\Omega^2)}}$$
$$OTOC(t,\varphi) = 1 - \varepsilon e^{\frac{2\pi}{\beta}t} h(\Omega t - \varphi)$$

Naively, we have three Lyapunov exponents (non-maximal chaos?)

$$\lambda_{\pm} = \frac{2\pi}{\beta(1 \mp \Omega)}, \quad \lambda_m = \frac{2\pi(1 - \mu\Omega)}{\beta(1 - \Omega^2)} \qquad v_{\pm} = \pm 1, \quad v_m = \frac{1 - \mu\Omega}{\Omega - \mu}$$

Periodicity in $\phi \colon h(\phi) \to h(\phi \bmod 2\pi)$ there is a constraint equation among c_i

OTOC
$$(t, \varphi) = 1 - \varepsilon_{VW} e^{\frac{2\pi}{\beta}t} \left[\alpha_1 h_1(\Omega t - \varphi) + \alpha_2 h_2(\Omega t - \varphi) \right]$$

There are two independent "instantaneous Lyapunov exponents" $\lambda_{\text{I,inst.}}(t) = \frac{2\pi}{\beta} + \frac{\Omega \partial_t h_1(\Omega t)}{h_1(\Omega t)}, \quad \lambda_{\text{II,inst.}}(t) = \frac{2\pi}{\beta} + \frac{\Omega \partial_t h_2(\Omega t)}{h_2(\Omega t)}$

High T limit of instantaneous Lyapunov exponents

ightharpoonup When $\mu>1$

$$\lambda_{\text{I, inst.}} = \begin{cases} \lambda_{-}, & \text{if } t \in \left[0, \frac{\pi(1+\Omega)}{\Omega}\right) \\ \lambda_{+}, & \text{if } t \in \left[\frac{\pi(1+\Omega)}{\Omega}, \frac{2\pi}{\Omega}\right) \end{cases}; \quad \lambda_{\text{II, inst.}} = \begin{cases} \lambda_{m}, & \text{if } t \in \left[0, \frac{2\pi(1+\Omega)}{\Omega(1+\mu)}\right) \\ \lambda_{+}, & \text{if } t \in \left[\frac{2\pi(1+\Omega)}{\Omega(1+\mu)}, \frac{2\pi}{\Omega}\right) \end{cases}$$

$$\langle \lambda_{\rm I, inst.} \rangle = \langle \lambda_{\rm II, inst.} \rangle = \frac{2\pi}{\beta}$$

ightharpoonup When $\mu < 1$

$$\Omega < \mu \qquad \lambda_{\text{I, inst.}} = \lambda_{\text{II, inst.}} = \begin{cases} \lambda_m \,, & \text{if } t \in \left[0, \, \frac{2\pi(1+\Omega)}{\Omega(1+\mu)}\right) \\ \lambda_+ \,, & \text{if } t \in \left[\frac{2\pi(1+\Omega)}{\Omega(1+\mu)}, \, \frac{2\pi}{\Omega}\right) \end{cases} \qquad \langle \lambda_{\text{I, inst.}} \rangle = \langle \lambda_{\text{II, inst.}} \rangle = \frac{2\pi}{\beta}$$

$$\mu < \Omega \qquad \lambda_{\text{I, inst.}} = \begin{cases} \lambda_{-}, & \text{if } t \in \left[0, \frac{2\pi(\Omega - \mu)}{\Omega(1 - \mu)}\right) \\ \lambda_{m}, & \text{if } t \in \left[\frac{2\pi(\Omega - \mu)}{\Omega(1 - \mu)}, \frac{2\pi}{\Omega}\right) \end{cases}; \quad \lambda_{\text{II, inst.}} = \lambda_{m}, \quad \text{if } t \in \left[0, \frac{2\pi}{\Omega}\right) \\ \langle \lambda_{\text{I, inst.}} \rangle = \frac{2\pi}{\beta} \qquad \langle \lambda_{\text{II, inst.}} \rangle = \lambda_{m} > \frac{2\pi}{\beta} \end{cases}$$

$$\Omega = \mu \qquad \lambda_{\text{I, inst.}} = \begin{cases} \lambda_{-}, & \text{if } t \in \left[0, \frac{\pi(1+\Omega)}{\Omega}\right) \\ \lambda_{+}, & \text{if } t \in \left[\frac{\pi(1+\Omega)}{\Omega}, \frac{2\pi}{\Omega}\right) \end{cases}; \quad \lambda_{\text{II, inst.}} = \lambda_{m}, \text{ if } t \in \left[0, \frac{2\pi}{\Omega}\right),$$

$$\langle \lambda_{\rm I, \; inst.} \rangle = \langle \lambda_{\rm II, \; inst.} \rangle = \frac{2\pi}{\beta}$$

Lyapunov exponent and butterfly velocities from OTOC

- ** In the high temperature limit and $|\Omega t \varphi| \ll 1$
 - \blacktriangleright When $\mu > 1$

$$OTOC(t,\varphi) = 1 - \varepsilon \begin{cases} \# e^{\lambda_{+}(t-\varphi)}, & \text{if } \Omega t < \varphi \\ \#_{1} e^{\lambda_{-}(t+\varphi)} + \#_{2} e^{\lambda_{m}\left(t-\frac{\varphi}{v_{m}}\right)}, & \text{if } \Omega t > \varphi \end{cases}$$

 \blacktriangleright When $\mu < 1$

$$\Omega < \mu$$
 OTOC $(t, \varphi) = 1 - \varepsilon \begin{cases} \#_1 e^{\lambda_+(t-\varphi)}, & \text{if } \Omega t < \varphi \\ \#_2 e^{\lambda_m \left(t - \frac{\varphi}{v_m}\right)}, & \text{if } \Omega t > \varphi \end{cases}$

$$\mu < \Omega \qquad \text{OTOC}(t,\varphi) = 1 - \varepsilon \begin{cases} \# \ e^{\frac{4\pi^2}{\beta(1-\Omega)}} \big) e^{\lambda_m \left(t - \frac{\varphi}{v_m}\right)}, & \text{if } \Omega t < \varphi \\ \#_1 \ e^{\lambda_-(t+\varphi)} + \#_2 e^{\lambda_m \left(t - \frac{\varphi}{v_m}\right)}, & \text{if } \Omega t > \varphi \end{cases} \quad \text{violate the chaos} \quad \text{bound; Vm>C}$$

$$\Omega = \mu \qquad \text{OTOC}(t, \varphi) = 1 - \varepsilon \begin{cases} \#_1 e^{\lambda_+(t-\varphi)} + \#_2 e^{\frac{2\pi}{\beta}t}, & \text{if } \Omega t < \varphi \\ \#_1 e^{\lambda_-(t+\varphi)} + \#_2 e^{\frac{2\pi}{\beta}t}, & \text{if } \Omega t > \varphi \end{cases}$$

Chaos in TMG from OTOC

• When $\mu = 1$, the profile of the shock wave is

$$h(\phi) = c_1 e^{-\frac{2\pi}{\beta(1+\Omega)}\phi} + c_2 e^{\frac{2\pi}{\beta(1-\Omega)\phi}} + c_3 \phi e^{-\frac{2\pi}{\beta(1+\Omega)}\phi}$$

$$\lambda_{\text{I, inst.}} = \lambda_{\text{II, inst.}} = \lambda_{-}, \quad \text{for } t \in \left[0, \frac{2\pi}{\Omega}\right)$$

$$\text{OTOC}(t, \varphi) = 1 - \varepsilon \begin{cases} \#_1 e^{\lambda_{-}(t+\varphi)}, & \text{if } \Omega t < \varphi \\ \#_2 e^{\lambda_{-}(t+\varphi)}, & \text{if } \Omega t > \varphi \end{cases}$$

- At high temperature, if we impose the chaos bound on the average of instantaneous Lyapunov exponent, only $\mu \ge 1$ is allowed; If we lower the temperature and impose the chaos bound, only the chiral point is allowed
- \blacksquare At high temperature, if we impose the chaos bound on the Lyapunov exponent, only $\mu \ge 1$ is allowed

Pole skipping from holography

Pole skipping from near horizon EOM

$$e_{vv}^{(0)}h_{vv}^{(0)} + e_{vr}^{(0)}h_{vr}^{(0)} + e_{v\phi}^{(0)}h_{v\phi}^{(0)} + e_{r\phi}^{(0)}h_{r\phi}^{(0)} + e_{\phi\phi}^{(0)}h_{\phi\phi}^{(0)} + e_{vv}^{(1)}h_{vv}^{(1)} + e_{v\phi}^{(1)}h_{v\phi}^{(1)} = 0$$

$$\left(\frac{2\pi i}{\beta(1\mp\Omega)}\,,\quad\mp\frac{2\pi i}{\beta(1\mp\Omega)}\right)\quad\&\quad\left(\frac{2\pi i(1-\Omega\mu)}{\beta(1-\Omega^2)}\,,\quad\frac{2i\pi(\Omega-\mu)}{\beta(1-\Omega^2)}\right)$$

Pole skipping from holographic massive mode

$$h_{ij}(\rho) = e^{-i\omega T + ikX} \left[h_{ij}^{(0)} + \rho h_{ij}^{(1)} + \rho^2 h_{ij}^{(2)} + \rho^{-\delta} \left(b_{ij}^{(0)} + \rho b_{ij}^{(1)} + \rho^2 b_{ij}^{(2)} + \cdots \right) + \rho^{\delta+1} \left(c_{ij}^{(0)} + \rho c_{ij}^{(1)} + \rho^2 c_{ij}^{(2)} + \cdots \right) \right]$$

$$\left. + \rho^{\delta+1} \left(c_{ij}^{(0)} + \rho c_{ij}^{(1)} + \rho^2 c_{ij}^{(2)} + \cdots \right) \right]$$

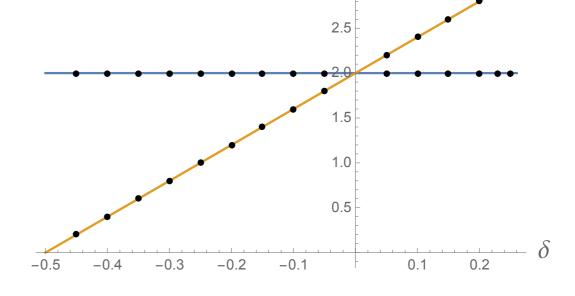
$$\left(\operatorname{Im} \omega, \operatorname{Im} k \right)$$

$$\mu = 2\delta + 1$$

$$G_R^{t^{00}t^{00}}(\omega,k) \propto \frac{c_{tt}^{(0)}}{b_{tt}^{(0)}} \quad \begin{array}{c} \operatorname{Im}\omega = -\operatorname{Im}k + 4(1-\omega) \\ \operatorname{Im}\omega = \operatorname{Im}k - 4\delta \end{array}$$

$$Im\omega = -Imk + 4(1+\delta)$$

$$Im\omega = Imk - 4\delta$$



3.0

Pole skipping from CFT

- The massive graviton is dual to an operator with conformal dimension $(2 + \delta, \delta)$
- The retarded Green's function

$$G_R(\omega, k) \propto \sin\left[\delta + \frac{i\beta_R}{2\pi} \left(\frac{\omega - k}{2}\right)\right] \sin\left[2 + \delta + \frac{i\beta_L}{2\pi} \left(\frac{\omega + k}{2}\right)\right] \times \left|\Gamma\left(\delta + \frac{i\beta_R}{2\pi} \left(\frac{\omega - k}{2}\right)\right)\right|^2 \left|\Gamma\left(2 + \delta + \frac{i\beta_L}{2\pi} \left(\frac{\omega + k}{2}\right)\right)\right|^2$$

Pole-skipping point

$$(\omega, k) = \left(\frac{2\pi i(1 - \Omega\mu)}{\beta(1 - \Omega^2)}, \frac{2i\pi(\Omega - \mu)}{\beta(1 - \Omega^2)}\right)$$

Conclusion

- OTOC and pole-skipping (from near horizon dynamics, holographic correlators, CFT calculations) are two features of quantum chaos
- For rotating BTZ in 3D Einstein gravity, we find a match between the two methods in the high temperature limit
- For rotating BTZ in 3D TMG, we find a match between these two methods in the high temperature limit and $\mu \ge 1$
- $\mu \geq 1$ is also the limit that the chaos bound is satisfied
- It would be interesting to study other systems with the non-maximal chaos (from CFT or holography)

Thank you!