

# Advanced Nuclear Physics

Nuclear Theory Group,  
Tohoku University  
**Kouichi Hagino**



## 原子核理論特論

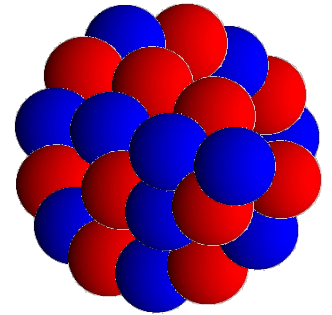
東北大学 → Kyoto U.

原子核理論研究室  
萩野浩一

# Contents

Nuclei: aggregate of nucleons (protons and neutrons)

→ *Nuclear Many-Body Problems*



(Low-energy) Nuclear Physics:

to understand rich nature of atomic nuclei starting from nucleon-nucleon interactions

- size, mass, density, shape
- excitations
- decays
- nuclear reactions

two kinds of particle: protons and neutrons

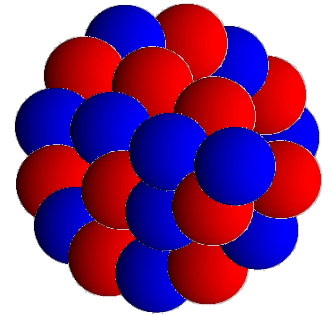
# Contents

Nuclei: aggregate of nucleons (protons and neutrons)

→ *Nuclear Many-Body Problems*

microscopic descriptions of atomic nuclei

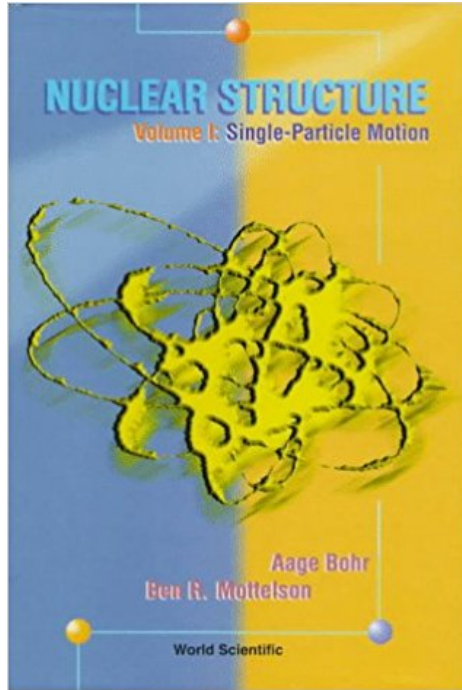
- Liquid drop model
- Single-particle motion and shell structure
- **Hartree-Fock approximation**
- Bruckner theory
- Pairing correlations and superfluid Nuclei
- Angular momentum and number projections
- 1-neutron and 2-neutron halo nuclei
- **Random phase approximation (RPA)**
- **Nuclear reactions**



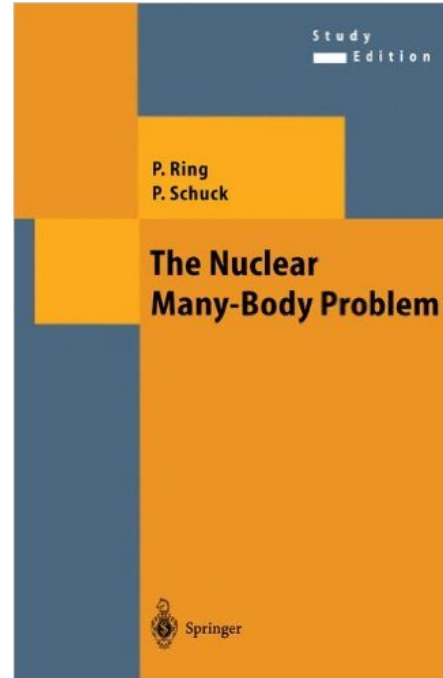
# Contents

1. Introduction
2. Shell energy
3. Mean-Field th. (HF)
4. Deformation
5. Pairing correlations

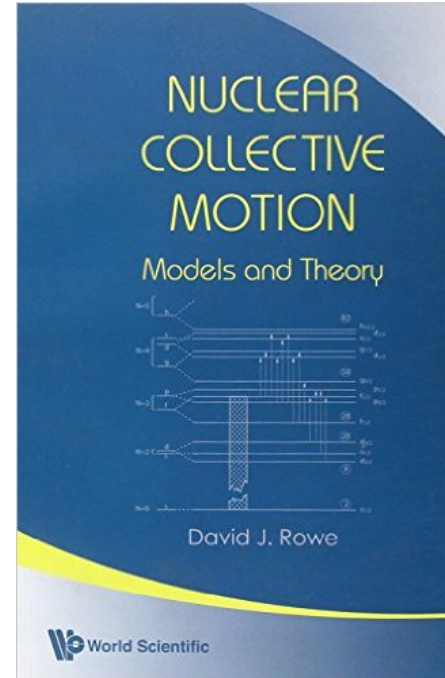
# References



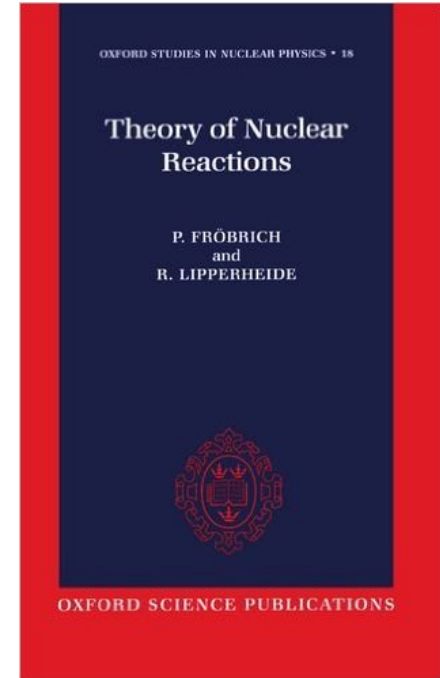
Bohr-Mottelson



Ring-Schuck



Rowe



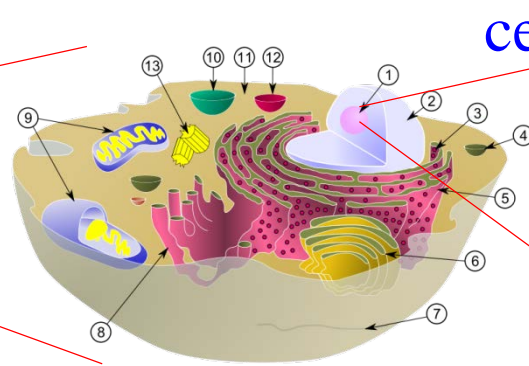
Frobrich  
-Lipperheide

- From Nucleons to Nucleus : Jouni Suhonen
- Introductory nuclear physics : Samuel S. Wong

# Introduction: atoms and atomic nuclei

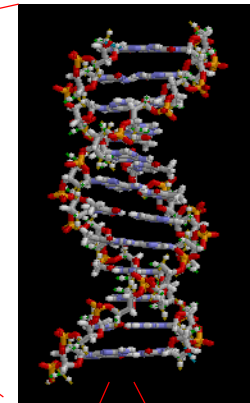


~ 50 cm



cells

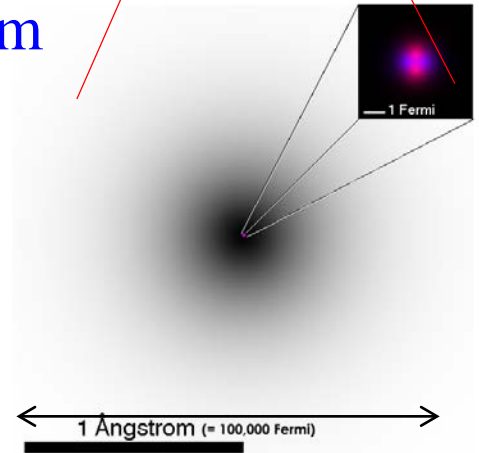
~  $\mu\text{m} = 10^{-6} \text{ m}$



DNA

~  $10^{-8} \text{ m}$

atom



~  $10^{-10} \text{ m}$

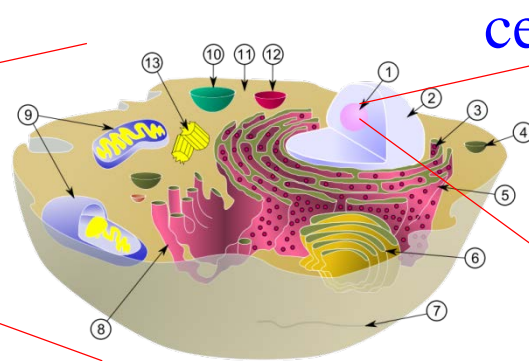
**Everything is made of atoms.**



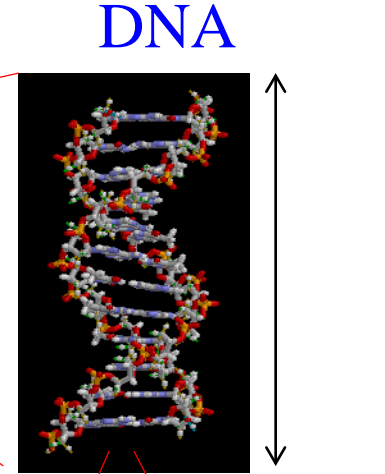
# Introduction: atoms and atomic nuclei



~ 50 cm



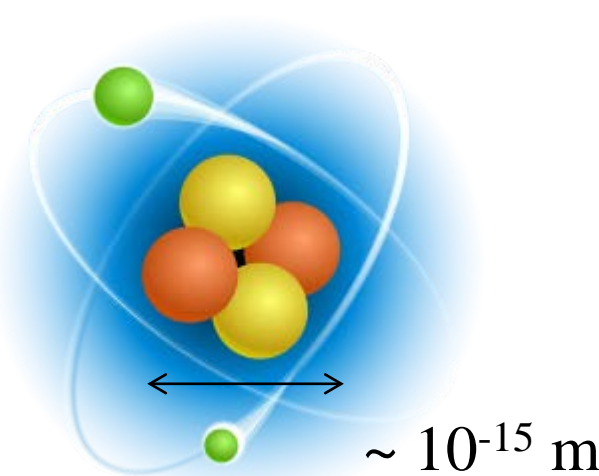
cells



DNA

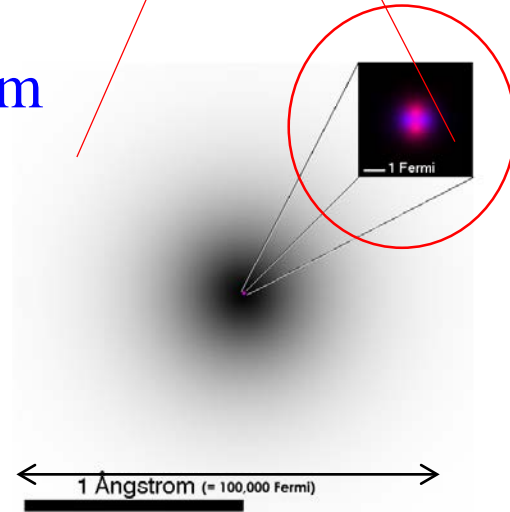
~  $10^{-8}$  m

atomic nucleus



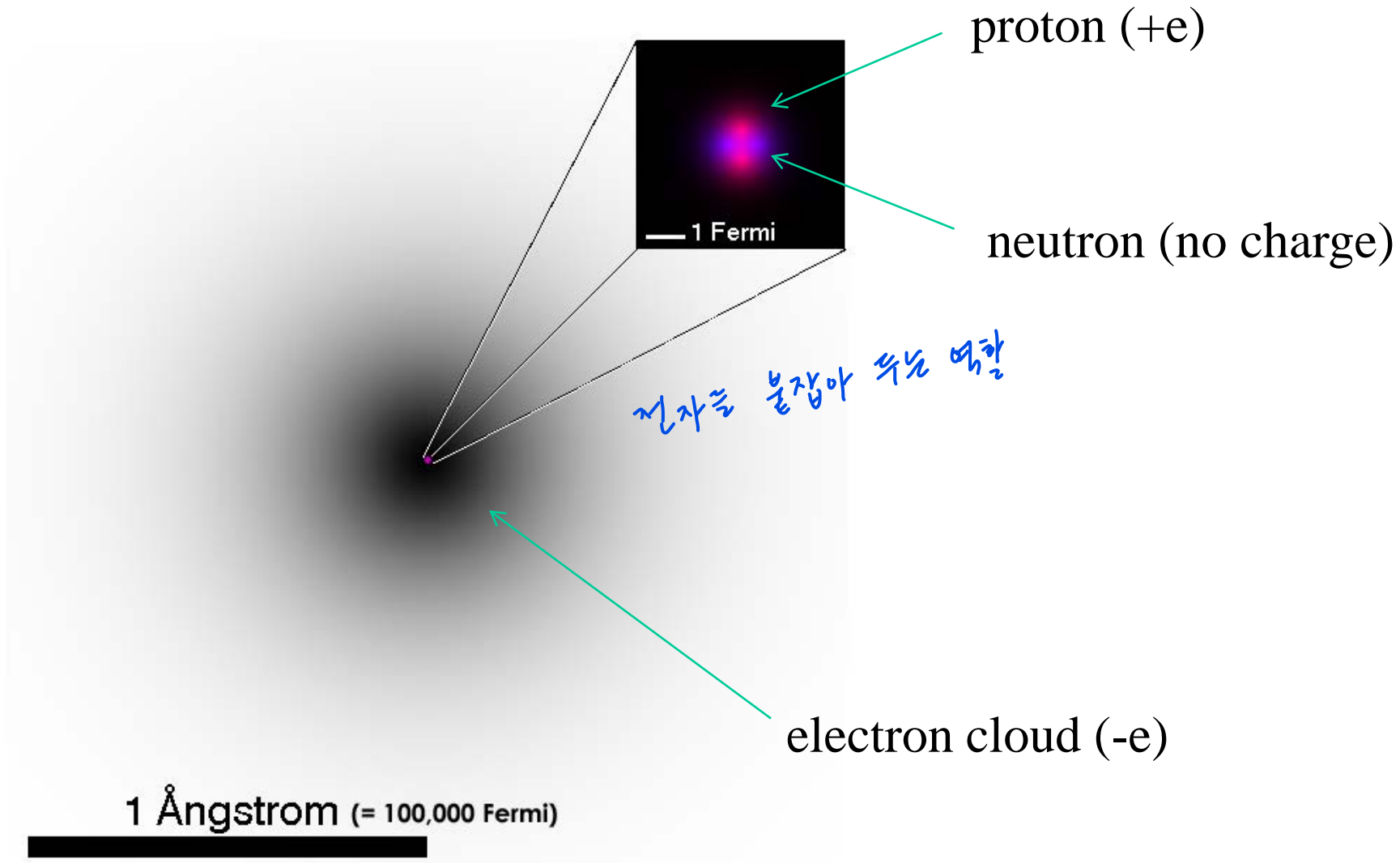
~  $10^{-15}$  m

atom



1 Angstrom (= 100,000 Fermi)

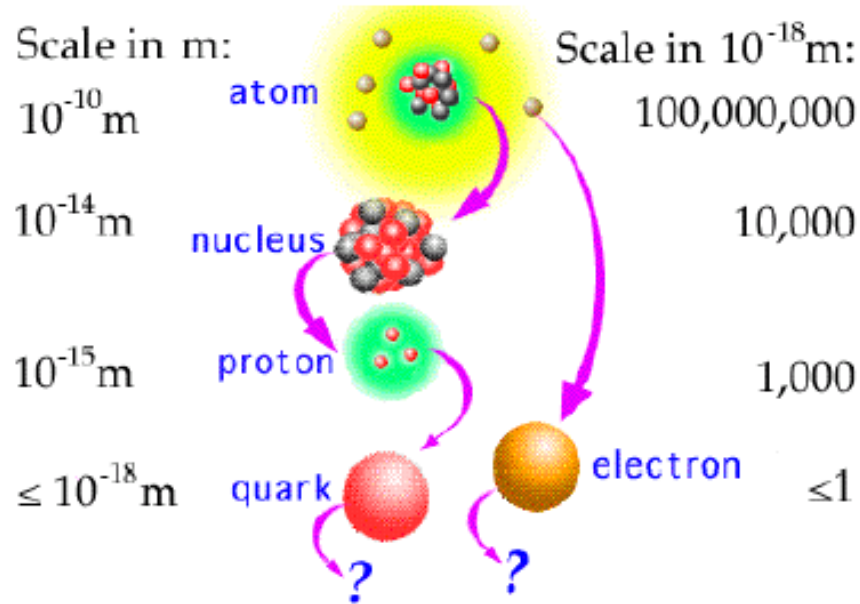
~  $10^{-10}$  m



- Neutral atoms: # of protons = # of electrons
- Chemical properties of atoms → # of electrons
- $M_p \sim M_n \sim 2000 M_e$  → the mass of atom ~ the mass of nucleus  
99.99%    부피 1/10<sup>14</sup>



# Nuclear Physics



$1 \text{ fm} = 10^{-15} \text{ m}$

Nucleus as a *quantum many body system*

*microscopic system of many mutually interacting particles.*

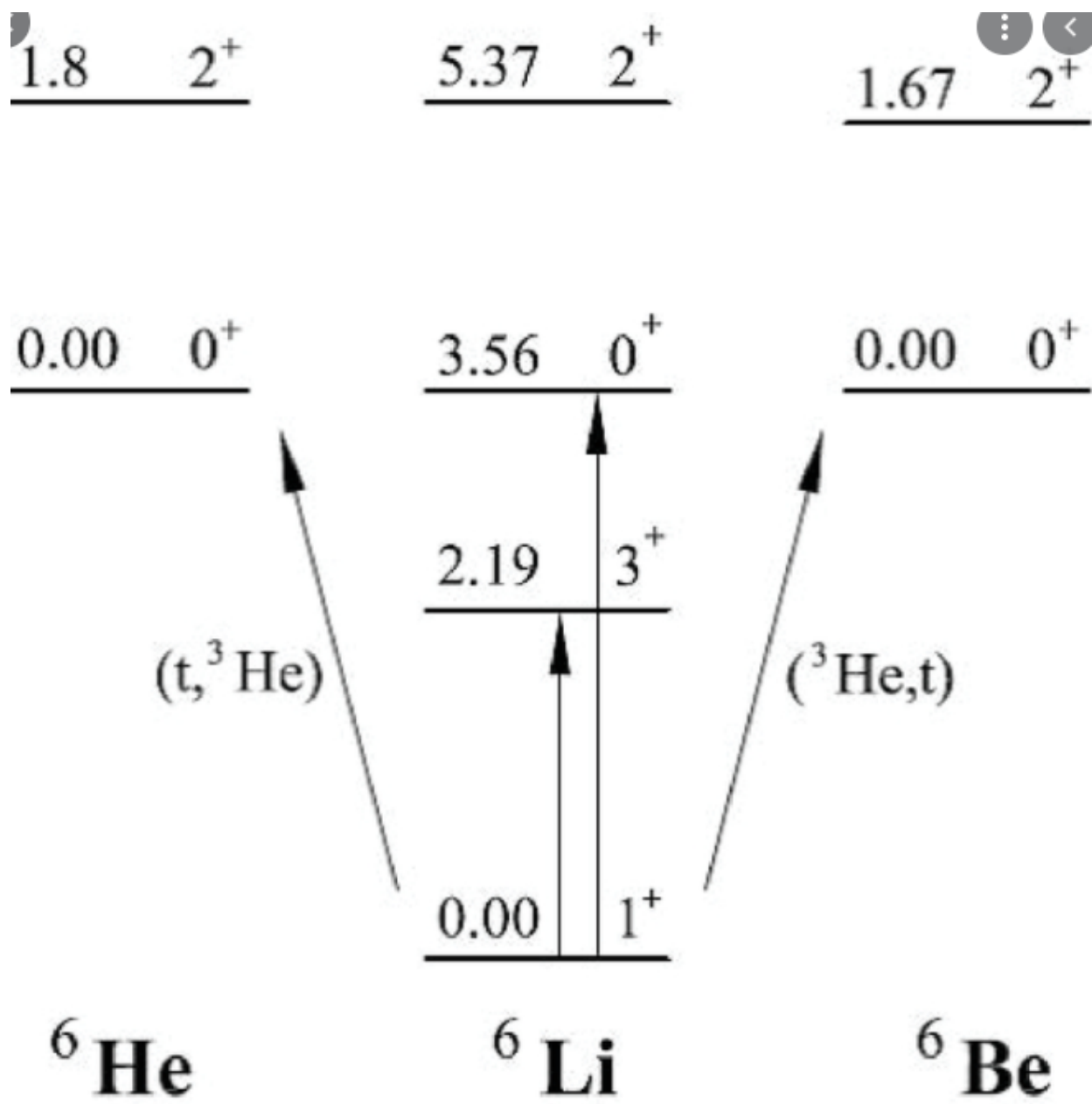
Basic ingredients:

	charge	mass (MeV)	spin
Proton	+e	938.256	1/2+
Neutron	0	939.550	1/2+

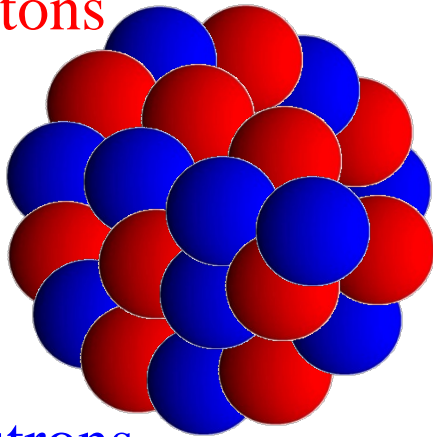
*if, w/o E.M Int.  
Isospin symm.  
n=p*

(note)  $n \rightarrow p + e^- + \bar{\nu}$  (10.4 min)  
*half-life*

*"Isobaric Analogue in same mass"*



protons

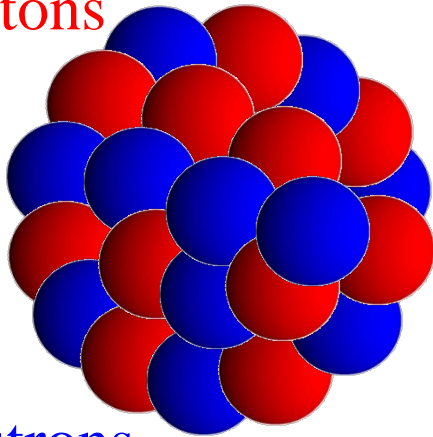


neutrons

- Nucleons are not stopping inside a nucleus.  
(they move relatively freely)
- Yet, they are not completely independent.  
A nucleus keeps its shape while nucleons influence among themselves so that a nucleon does not escape.

a self-bound system

protons

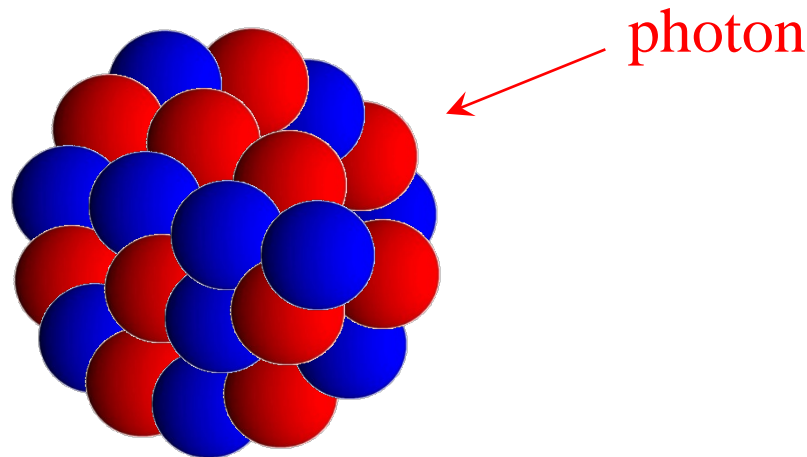


neutrons

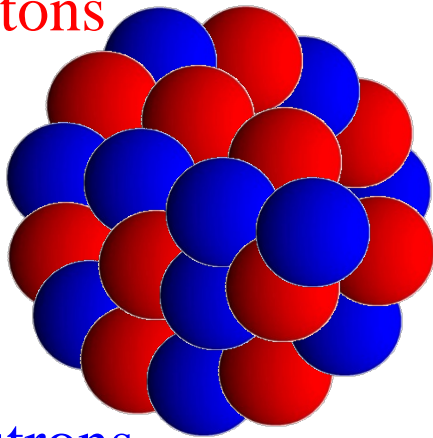
- Nucleons are not stopping inside a nucleus.  
(they move relatively freely)
- Yet, they are not completely independent.  
A nucleus keeps its shape while nucleons influence among themselves so that a nucleon does not escape.

a self-bound system

What happens if a photon is absorbed into a nucleus?  
- one nucleon simply starts moving faster?



protons



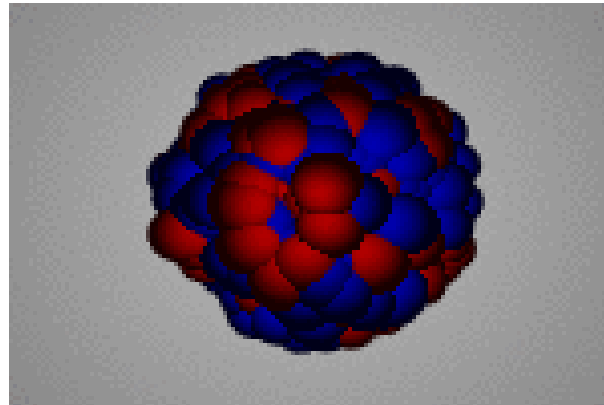
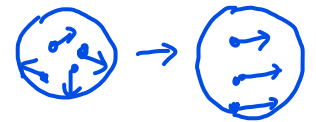
neutrons

- Nucleons are not stopping inside a nucleus.  
(they move relatively freely)
- Yet, they are not completely independent.  
A nucleus keeps its shape while nucleons influence among themselves so that a nucleon does not escape.

a self-bound system

What happens if a photon is absorbed into a nucleus?

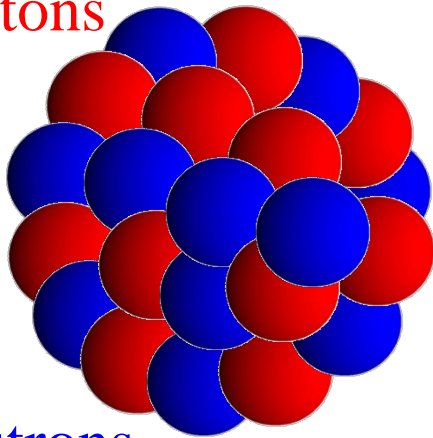
- one nucleon simply starts moving faster?



Very coherent  
motion can happen

Collective motions

protons

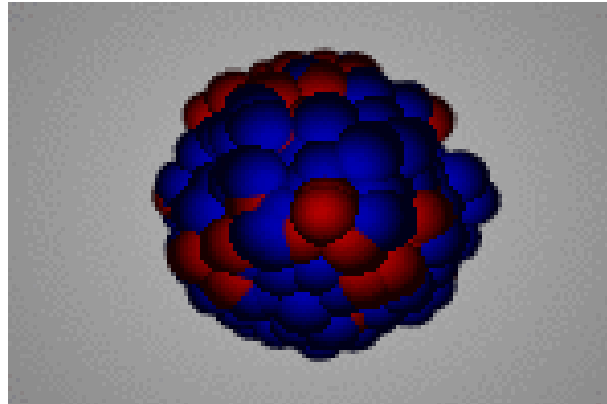
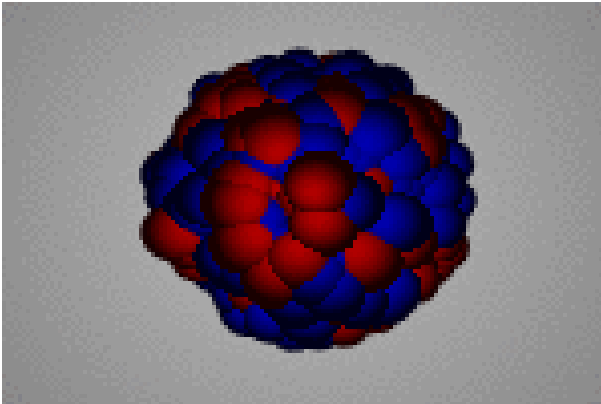


neutrons

- Nucleons are not stopping inside a nucleus.  
(they move relatively freely)
- Yet, they are not completely independent.  
A nucleus keeps its shape while nucleons influence among themselves so that a nucleon does not escape.

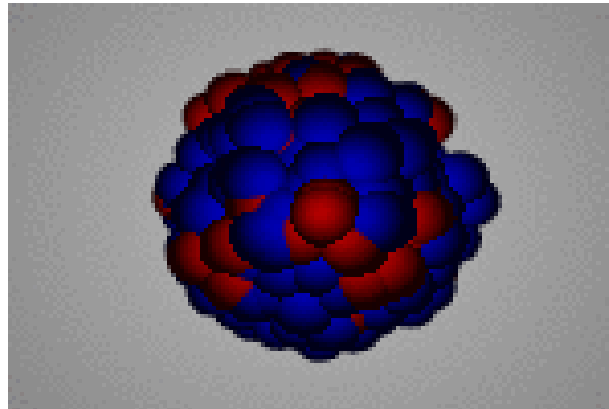
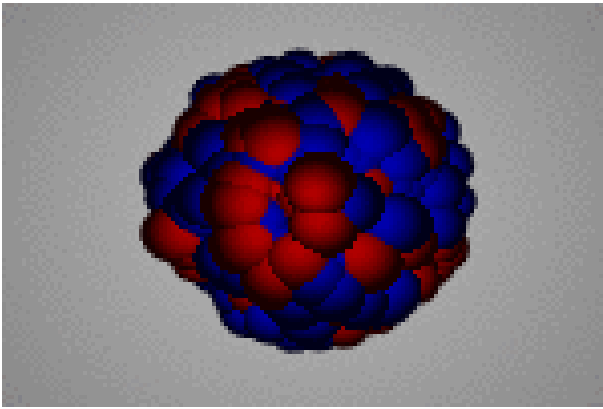
a self-bound system

What happens if a photon is absorbed into a nucleus?  
- one nucleon simply starts moving faster?



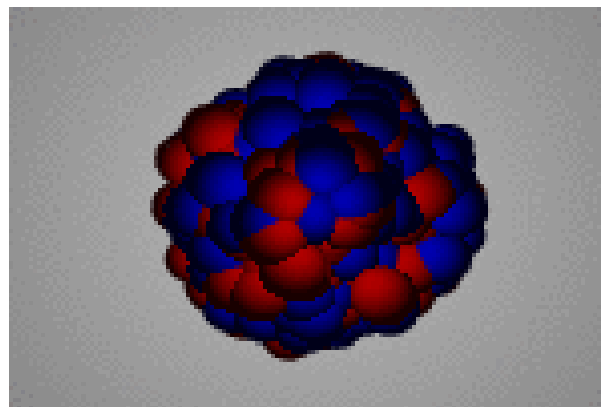
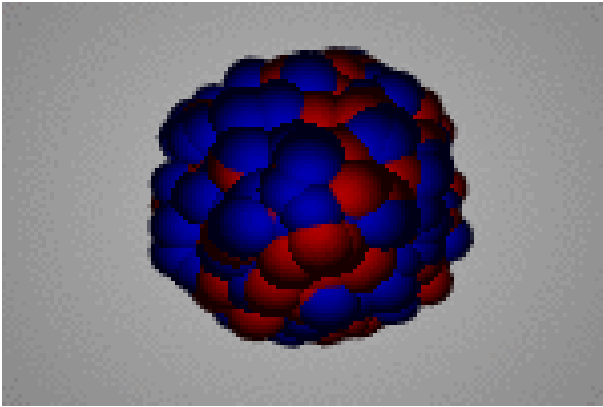
Very coherent  
motion can happen

Collective motions

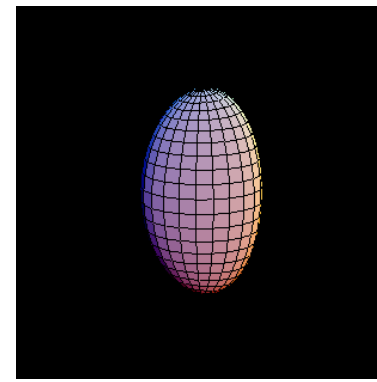
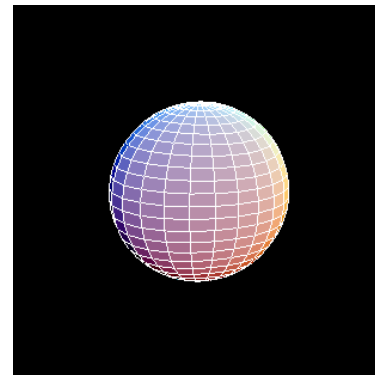
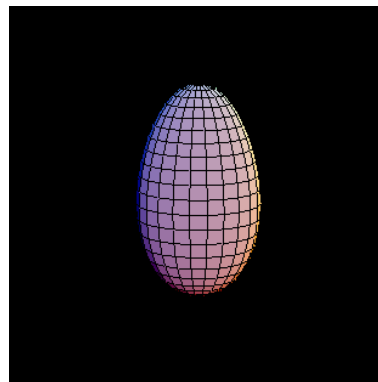
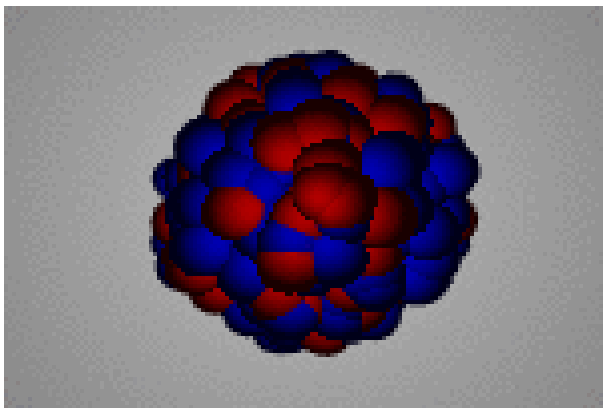


Very coherent  
motion can happen

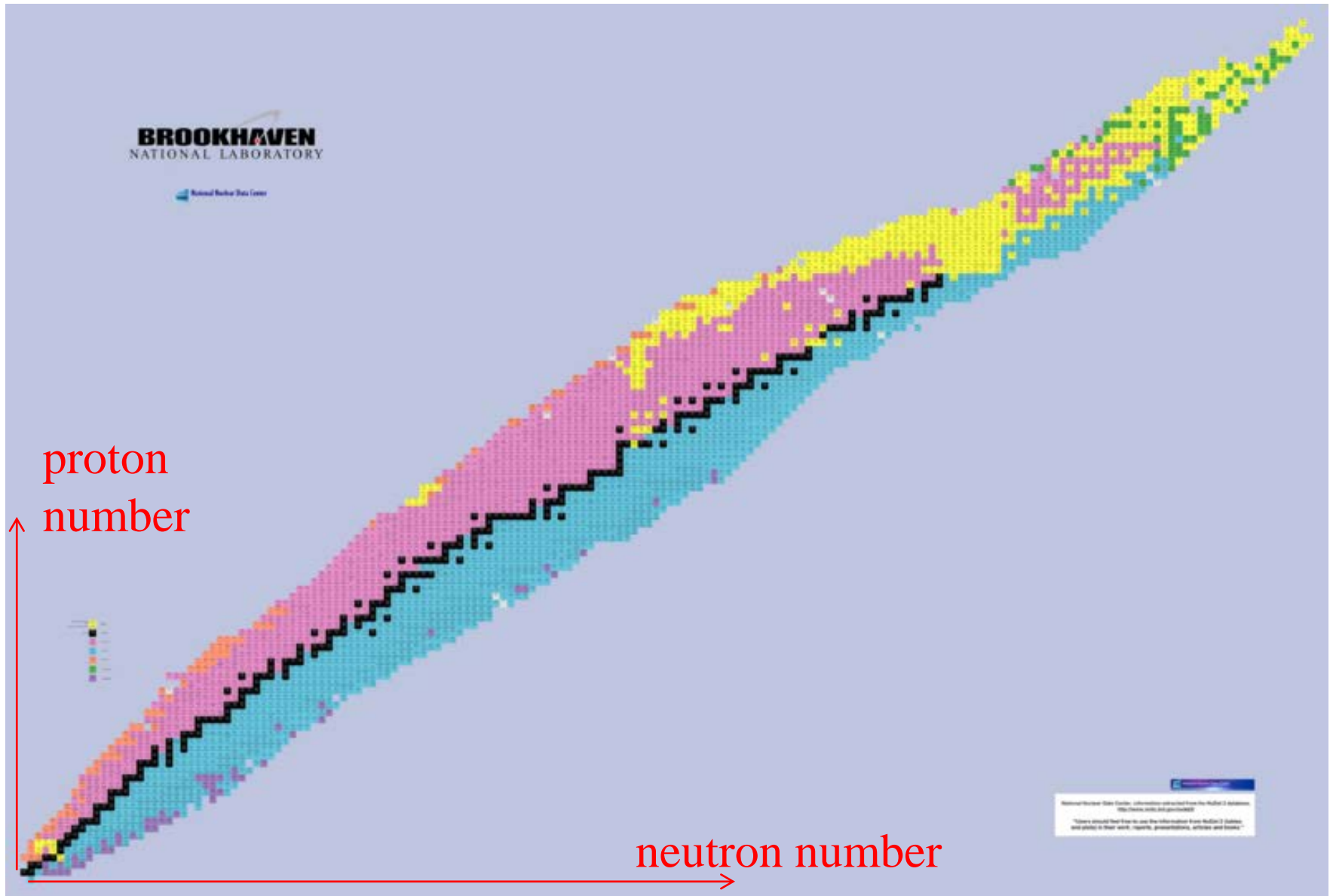
Collective motions



a variety of  
motions  
→ very rich!



# Nuclear Chart: 2D map of atomic nuclei





# Periodic table: protons only, no neutrons

## 元素の周期表

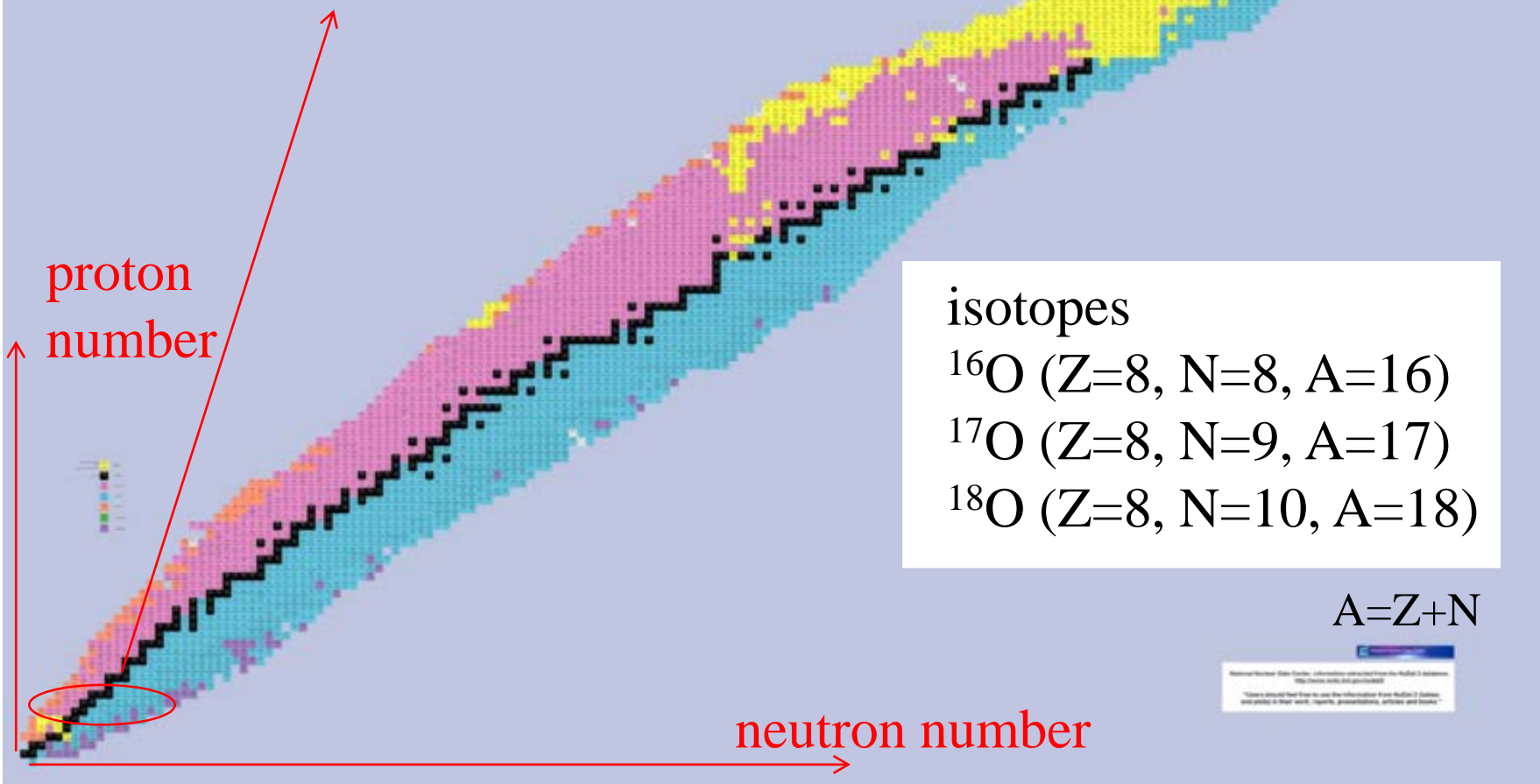
	1A	2A	3A	4A	5A	6A	7A	8	1B	2B	3B	4B	5B	6B	7B	0		
1	<sup>1</sup> H															<sup>2</sup> He		
2	<sup>3</sup> Li	<sup>4</sup> Be									<sup>5</sup> B	<sup>6</sup> C	<sup>7</sup> N	<sup>8</sup> O	<sup>9</sup> F	<sup>10</sup> Ne		
3	<sup>11</sup> Na	<sup>12</sup> Mg									<sup>13</sup> Al	<sup>14</sup> Si	<sup>15</sup> P	<sup>16</sup> S	<sup>17</sup> Cl	<sup>18</sup> Ar		
4	<sup>19</sup> K	<sup>20</sup> Ca	<sup>21</sup> Sc	<sup>22</sup> Ti	<sup>23</sup> V	<sup>24</sup> Cr	<sup>25</sup> Mn	<sup>26</sup> Fe	<sup>27</sup> Co	<sup>28</sup> Ni	<sup>29</sup> Cu	<sup>30</sup> Zn	<sup>31</sup> Ga	<sup>32</sup> Ge	<sup>33</sup> As	<sup>34</sup> Se	<sup>35</sup> Br	<sup>36</sup> Kr
5	<sup>37</sup> Rb	<sup>38</sup> Sr	<sup>39</sup> Y	<sup>40</sup> Zr	<sup>41</sup> Nb	<sup>42</sup> Mo	<sup>43</sup> Tc	<sup>44</sup> Ru	<sup>45</sup> Rh	<sup>46</sup> Pd	<sup>47</sup> Ag	<sup>48</sup> Cd	<sup>49</sup> In	<sup>50</sup> Sn	<sup>51</sup> Sb	<sup>52</sup> Te	<sup>53</sup> I	<sup>54</sup> Xe
6	<sup>55</sup> Cs	<sup>56</sup> Ba	<sup>L</sup>	<sup>72</sup> Hf	<sup>73</sup> Ta	<sup>74</sup> W	<sup>75</sup> Re	<sup>76</sup> Os	<sup>77</sup> Ir	<sup>78</sup> Pt	<sup>79</sup> Au	<sup>80</sup> Hg	<sup>81</sup> Tl	<sup>82</sup> Pb	<sup>83</sup> Bi	<sup>84</sup> Po	<sup>85</sup> At	<sup>86</sup> Rn
7	<sup>87</sup> Fr	<sup>88</sup> Ra	<sup>A</sup>															
		<sup>L</sup>	<sup>57</sup> La	<sup>58</sup> Ce	<sup>59</sup> Pr	<sup>60</sup> Nd	<sup>61</sup> Pm	<sup>62</sup> Sm	<sup>63</sup> Eu	<sup>64</sup> Gd	<sup>65</sup> Tb	<sup>66</sup> Dy	<sup>67</sup> Ho	<sup>68</sup> Er	<sup>69</sup> Tm	<sup>70</sup> Yb	<sup>71</sup> Lu	
		<sup>A</sup>	<sup>89</sup> Ac	<sup>90</sup> Th	<sup>91</sup> Pa	<sup>92</sup> U	<sup>93</sup> Np	<sup>94</sup> Pu	<sup>95</sup> Am	<sup>96</sup> Cm	<sup>97</sup> Bk	<sup>98</sup> Cf	<sup>99</sup> Es	<sup>100</sup> Fm	<sup>101</sup> Md	<sup>102</sup> No	<sup>103</sup> Lr	

- 典型金属元素
- 半金属元素
- 非金属元素
- 遷移金属元素
- 希ガス

# Nuclear Chart: 2D map of atomic nuclei

isotope - same Z  
 isotone - same N  
 isobar - same A  
 isomer - same element

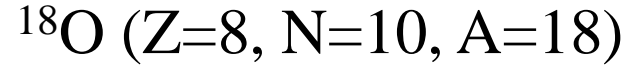
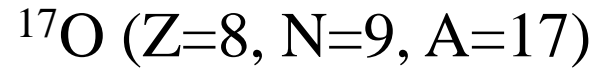
	$^{12}\text{O}$	$^{13}\text{O}$	$^{14}\text{O}$	$^{15}\text{O}$	$^{16}\text{O}$	$^{17}\text{O}$	$^{18}\text{O}$	$^{19}\text{O}$	$^{20}\text{O}$	$^{21}\text{O}$	$^{22}\text{O}$	$^{23}\text{O}$	$^{24}\text{O}$
		$^{12}\text{N}$	$^{13}\text{N}$	$^{14}\text{N}$	$^{15}\text{N}$	$^{16}\text{N}$	$^{17}\text{N}$	$^{18}\text{N}$	$^{19}\text{N}$	$^{20}\text{N}$	$^{21}\text{N}$	$^{22}\text{N}$	$^{23}\text{N}$
$^9\text{C}$	$^{10}\text{C}$	$^{11}\text{C}$	$^{12}\text{C}$	$^{13}\text{C}$	$^{14}\text{C}$	$^{15}\text{C}$	$^{16}\text{C}$	$^{17}\text{C}$	$^{18}\text{C}$	$^{19}\text{C}$	$^{20}\text{C}$		$^{22}\text{C}$



proton number

neutron number

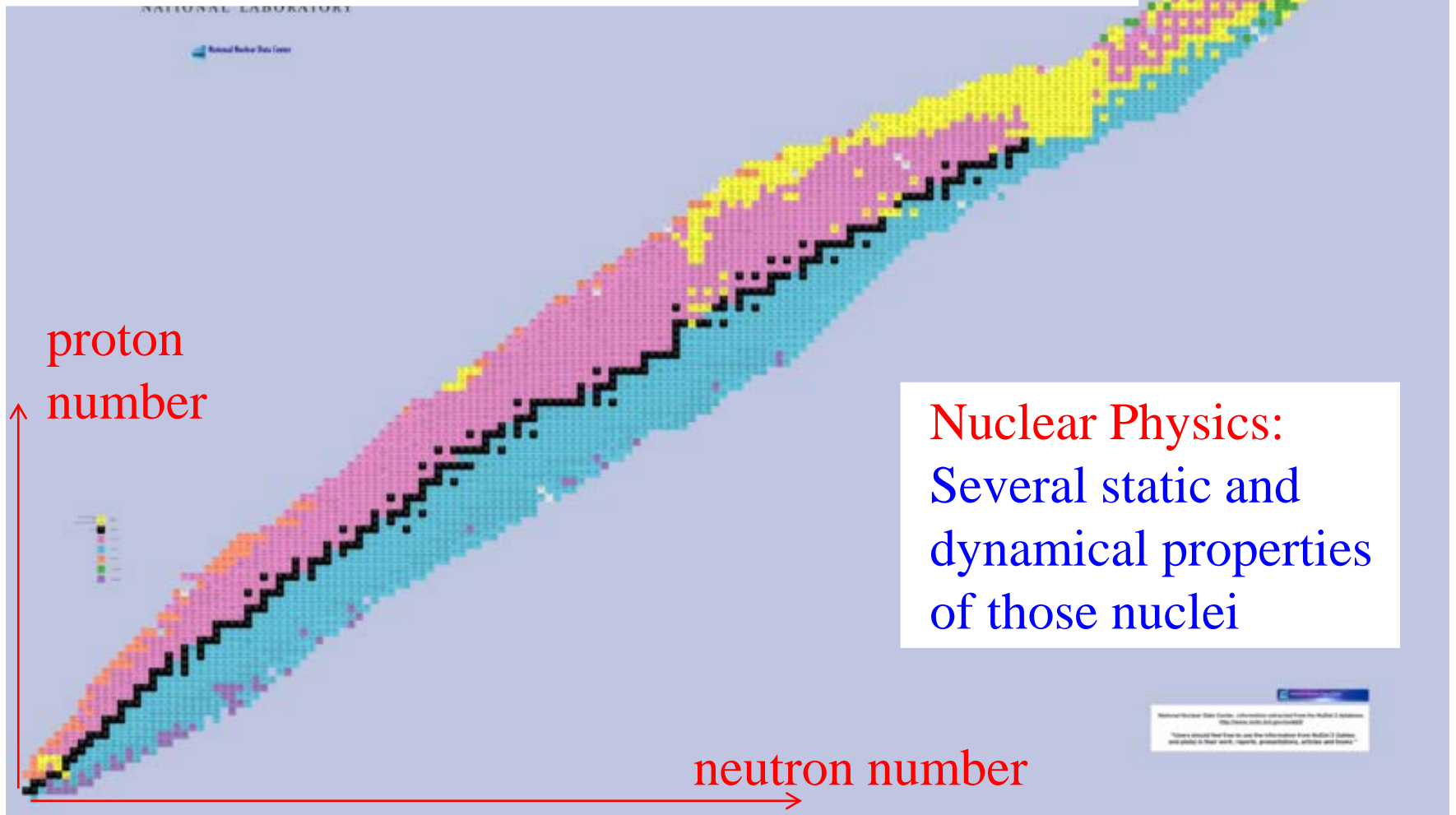
isotopes



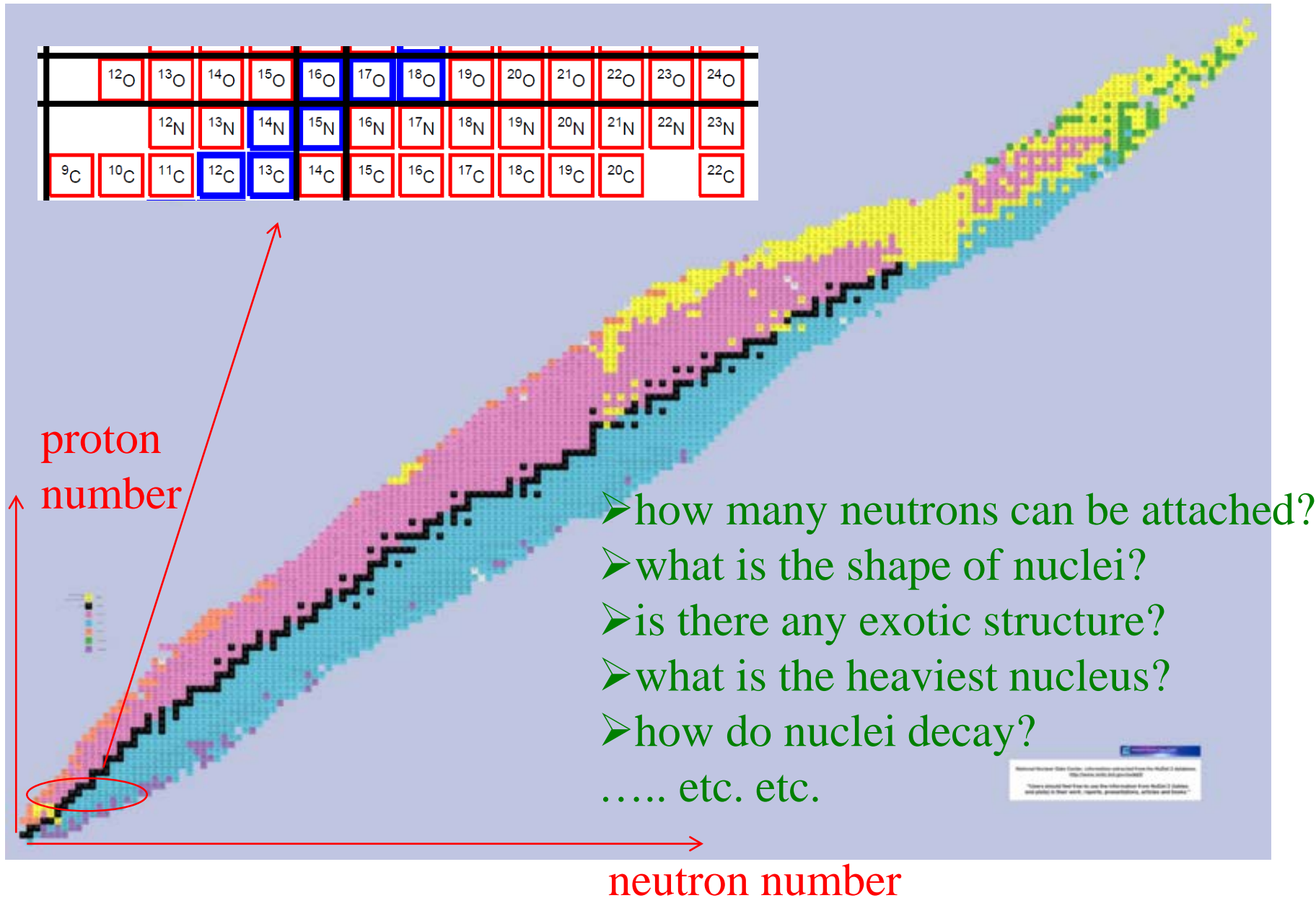
$A=Z+N$

Revised Nuclear Data Tables, information extracted from the NNDC database  
 (http://www.nndc.gov/)  
 "There should be no fee to use the information from NNDC's database, and please to their work, reports, presentations, articles and books."

- Stable nuclei in nature: 287
- Nuclei artificially synthesized : about 3,000
- Nuclei predicted : about 7,000 ~ 10,000

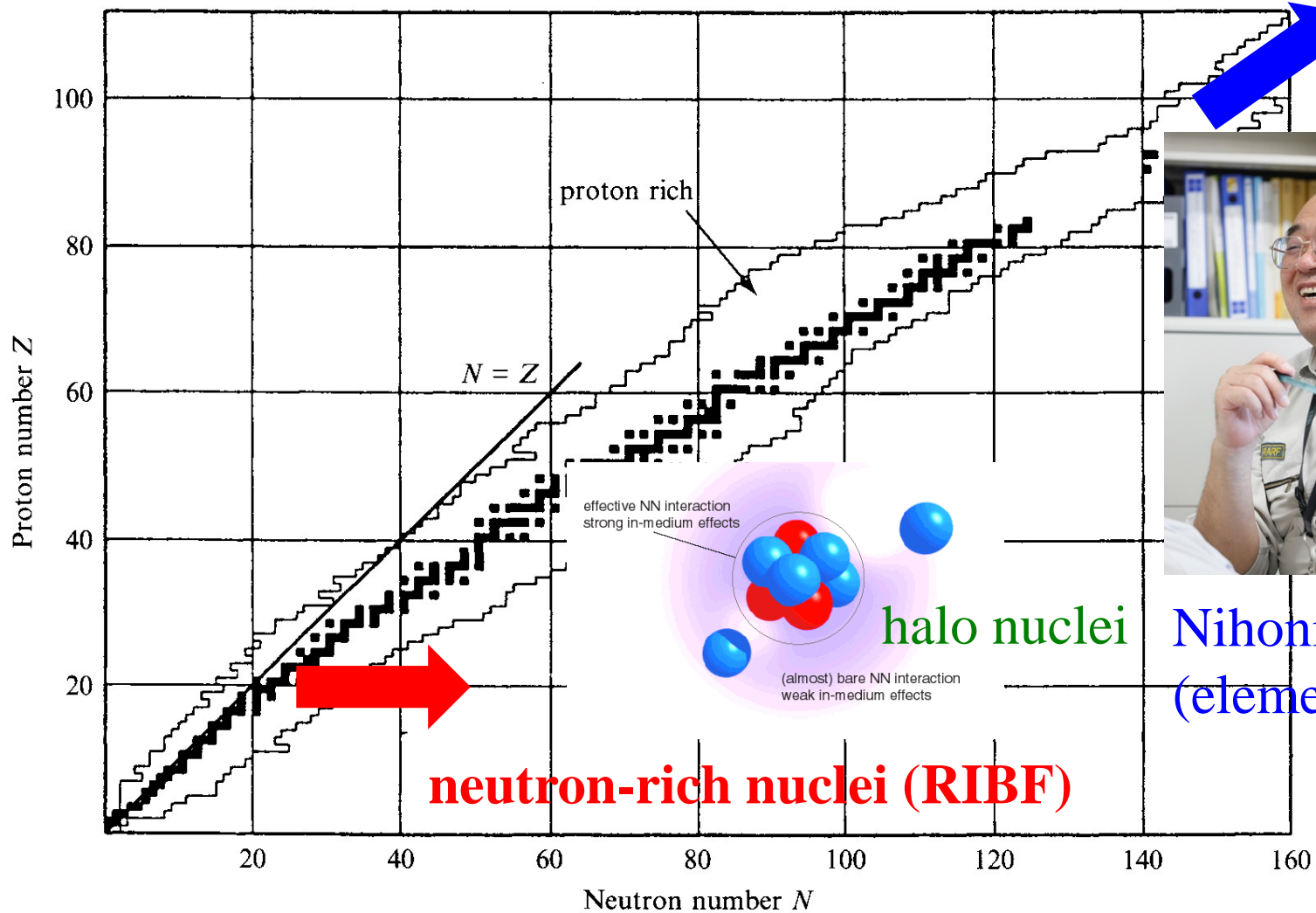


# Nuclear Chart: 2D map of atomic nuclei

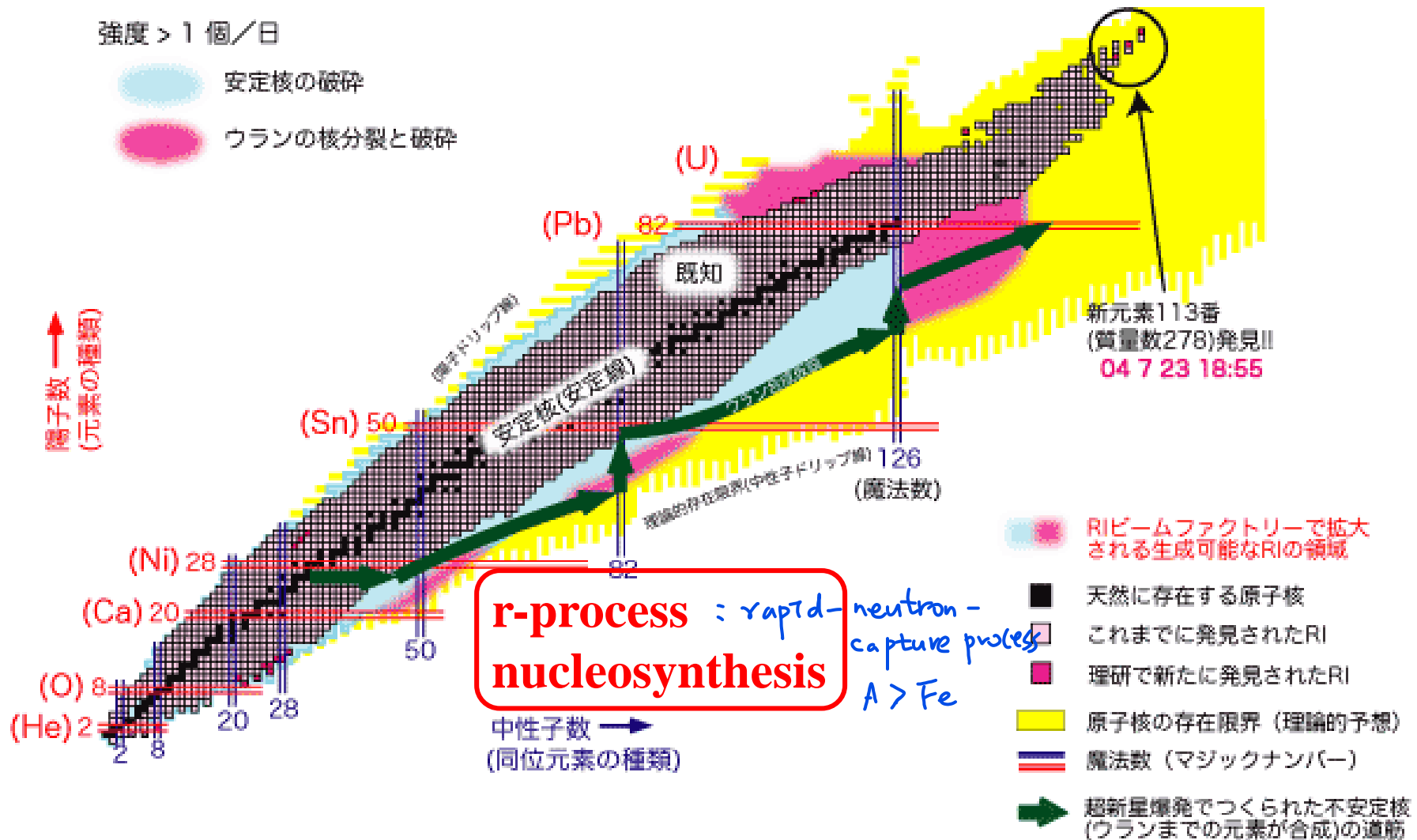


# Extension of nuclear chart: frontier of nuclear physics

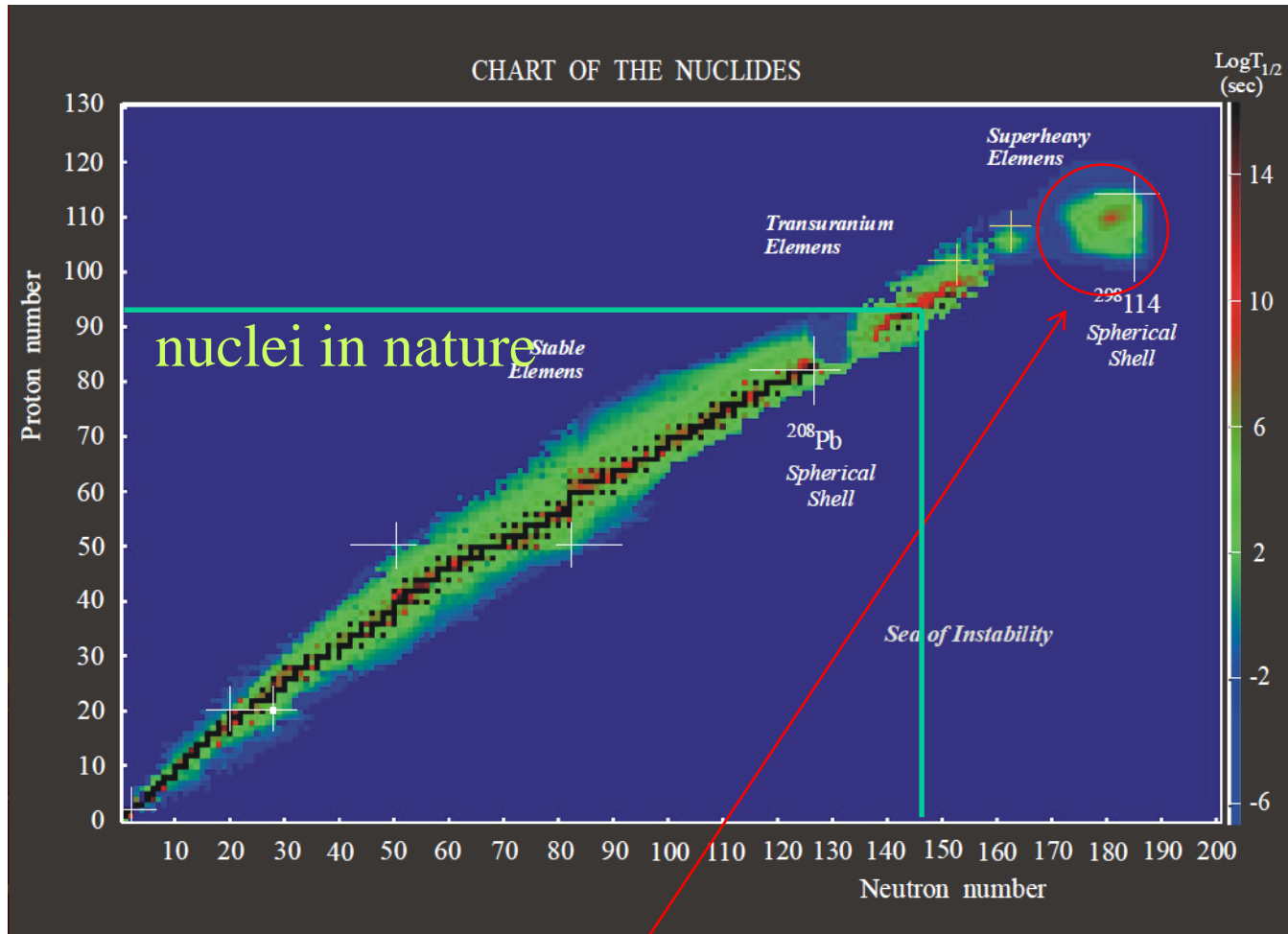
superheavy  
elements



# Neutron-rich nuclei (RIBF at RIKEN)



# Prediction of island of stability: an important motivation of SHE study



**island of stability around  $Z=114$ ,  $N=184$**

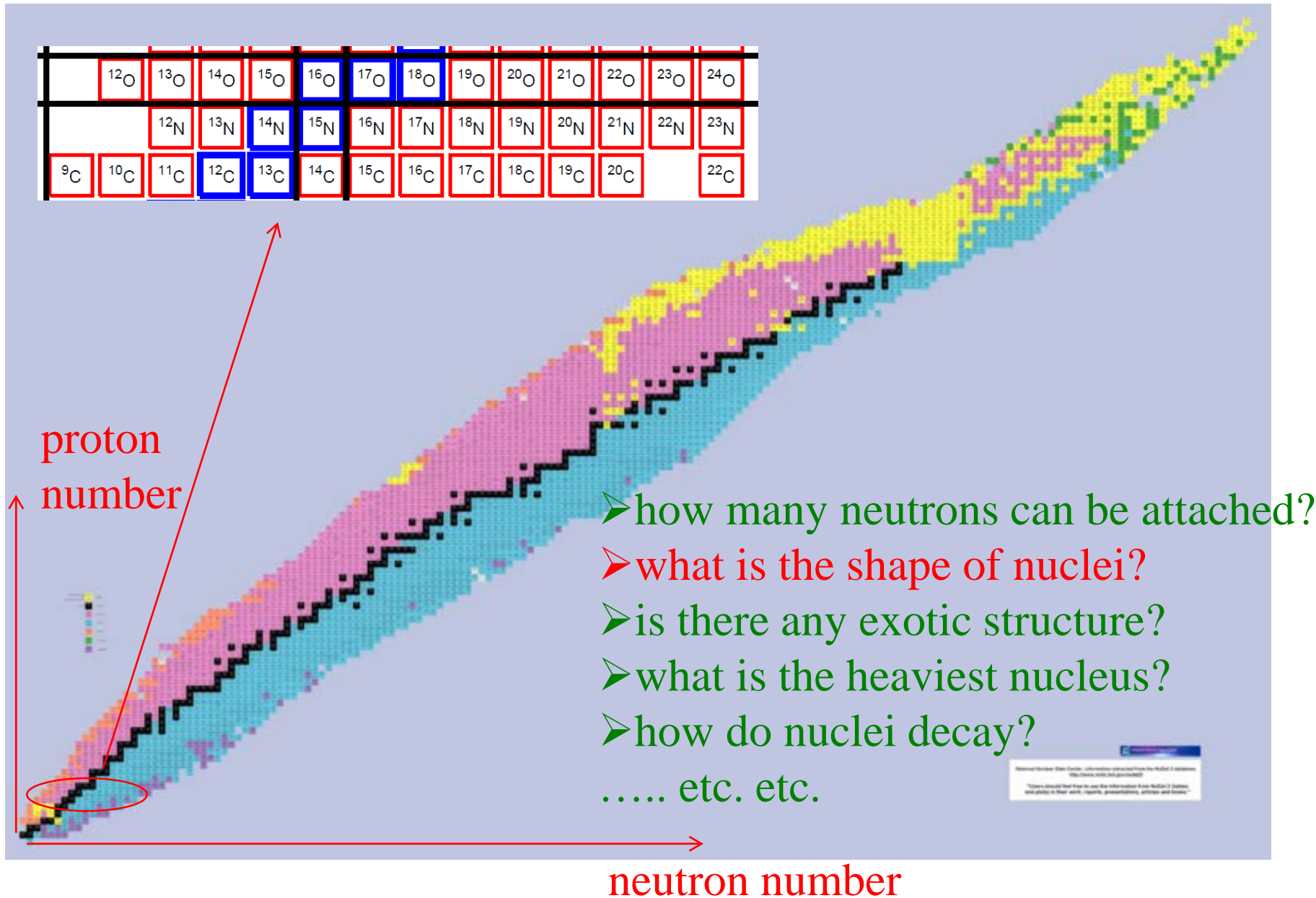
Yuri Oganessian

W.D. Myers and W.J. Swiatecki (1966), A. Sobiczewski et al. (1966)

→ modern calculations:  $Z=114, 120$ , or  $126$ ,  $N=184$

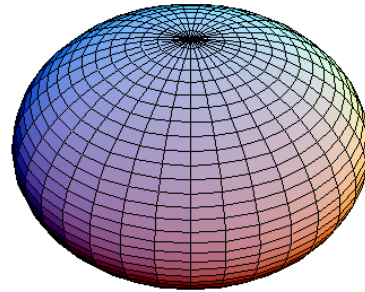
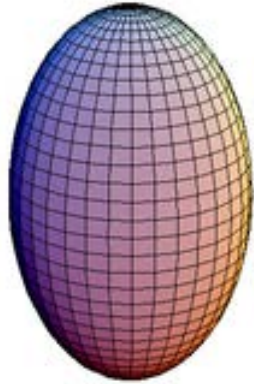
e.g., H. Koura et al. (2005)

# Nuclear Chart: 2D map of atomic nuclei



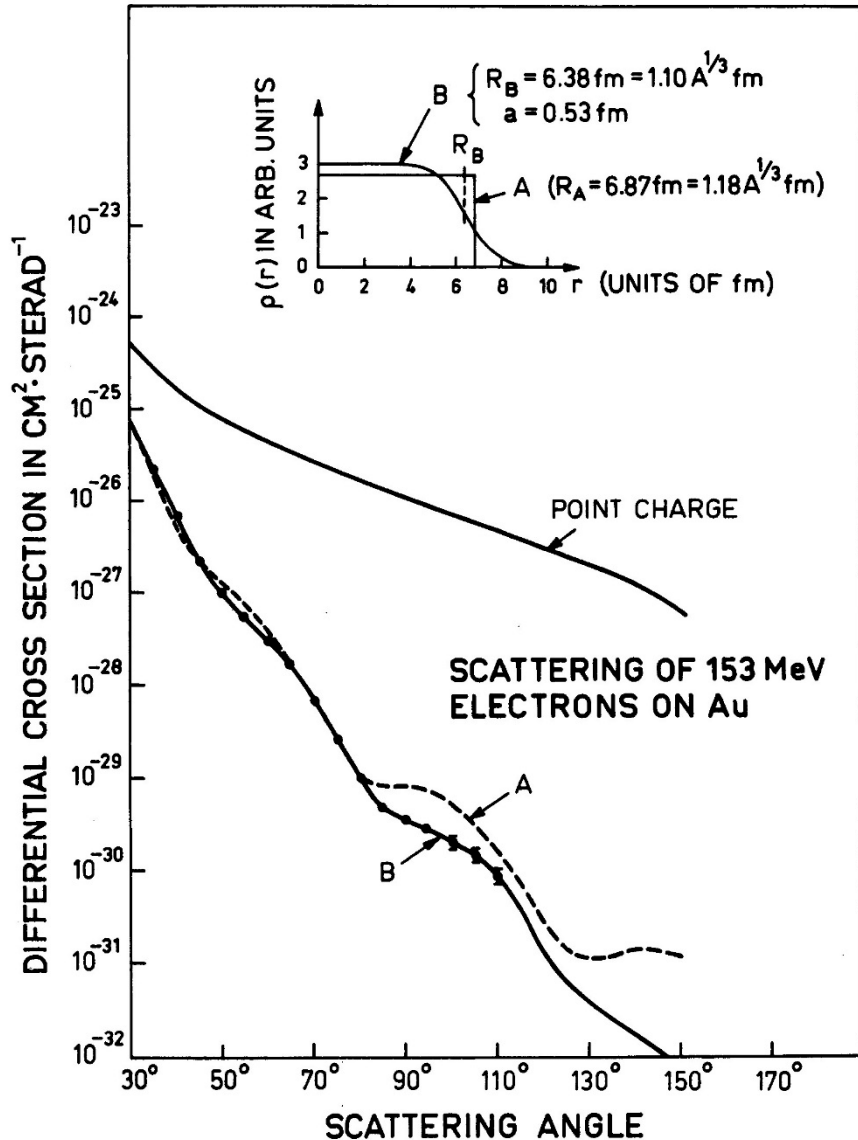


a nucleus is not always spherical



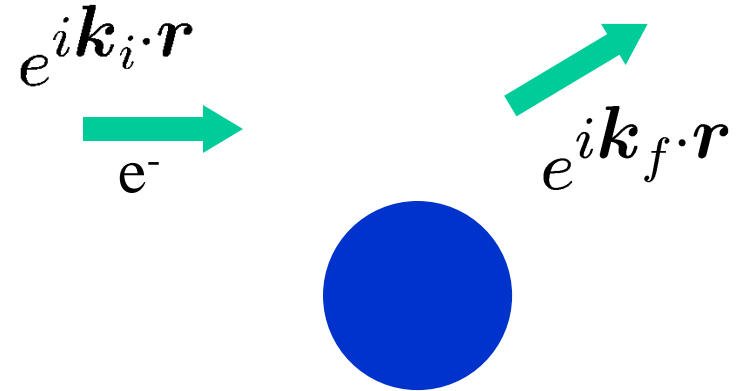
Quantum shape  
dynamics

# Density Distribution



## High energy electron scattering

Born approximation:



$$\frac{d\sigma}{d\Omega} = \frac{Z_P^2 e^4}{(4E \sin^2 \theta/2)^2} |F(\mathbf{q})|^2$$

Form factor

$$F(\mathbf{q}) = \int e^{-i\mathbf{q} \cdot \mathbf{r}} \rho(\mathbf{r}) d\mathbf{r}$$

(Fourier transform of the density)

# Born approximation

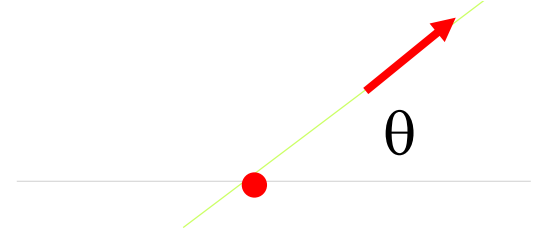
$$\psi_i(\mathbf{r}) = e^{i\mathbf{p}_i \cdot \mathbf{r} / \hbar}$$



$V(r)$



$$\psi_f(\mathbf{r}) = e^{i\mathbf{p}_f \cdot \mathbf{r} / \hbar}$$



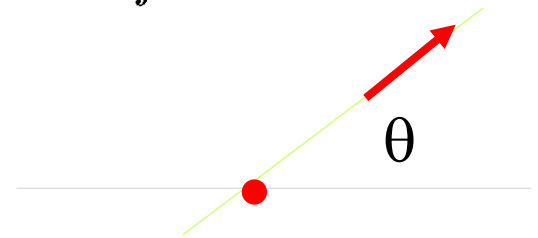
# Born approximation

$$\psi_f(\mathbf{r}) = e^{i\mathbf{p}_f \cdot \mathbf{r} / \hbar}$$

$$\psi_i(\mathbf{r}) = e^{i\mathbf{p}_i \cdot \mathbf{r} / \hbar}$$



$V(r)$



$$W_{fi} = \frac{\mu p_i}{4\pi^2 \hbar^4} \int d\Omega |\tilde{V}(\mathbf{q})|^2$$

momentum transfer



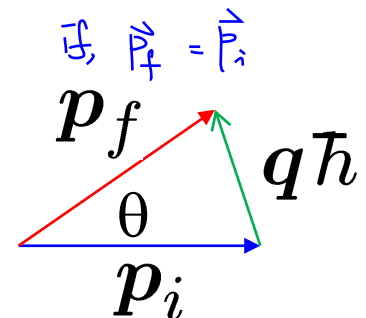
$$\tilde{V}(\mathbf{q}) = \int d\mathbf{r} e^{i(\mathbf{p}_i - \mathbf{p}_f) \cdot \mathbf{r} / \hbar} V(r) \equiv \int d\mathbf{r} e^{-i\mathbf{q} \cdot \mathbf{r}} V(r)$$

incident flux:  $j_{inc} = \rho_i v = p_i / \mu$



$$\sigma = \frac{W_{fi}}{j_{inc}} = \int d\Omega \frac{\mu^2}{4\pi^2 \hbar^4} |\tilde{V}(\mathbf{q})|^2$$

$$= \frac{d\sigma}{d\Omega}$$



$$q\hbar = 2p_i \sin \frac{\theta}{2}$$

# Electron scattering

$$V(r) = -e^2 \int d\mathbf{r}' \frac{\rho(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}$$

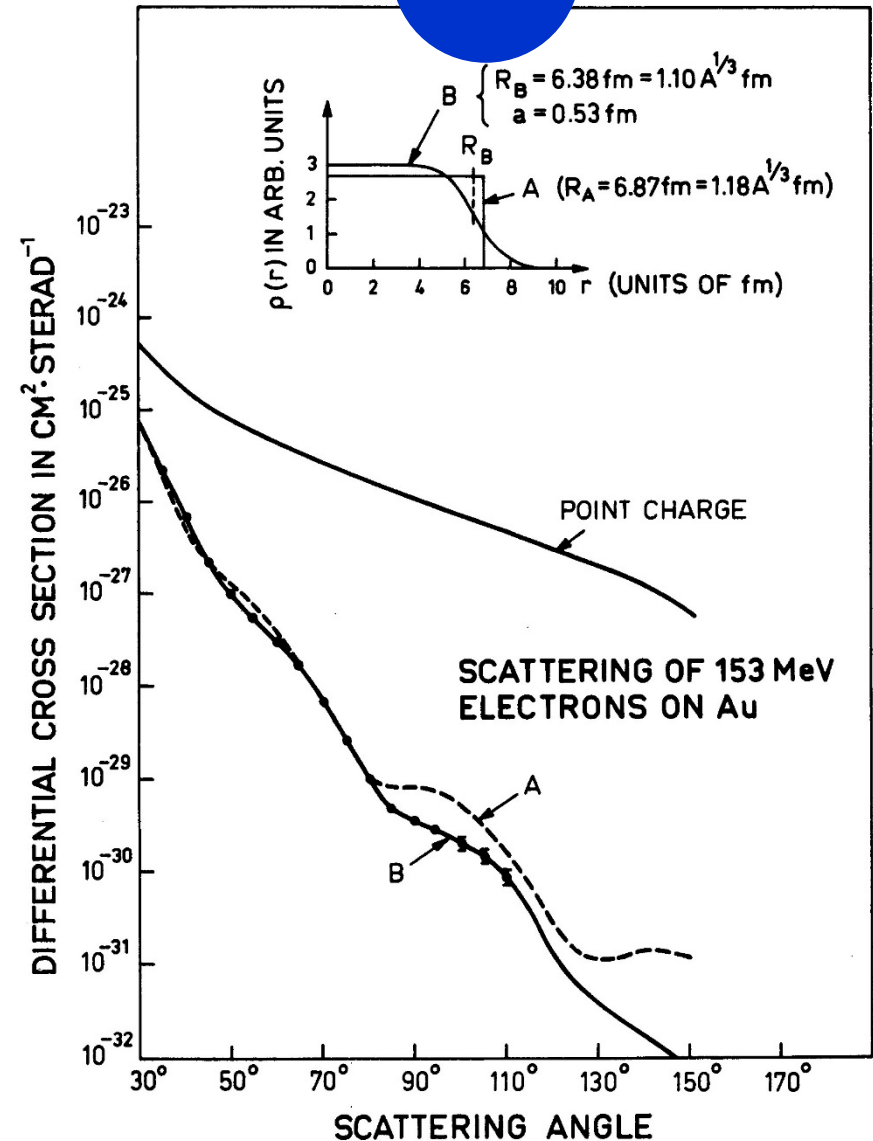
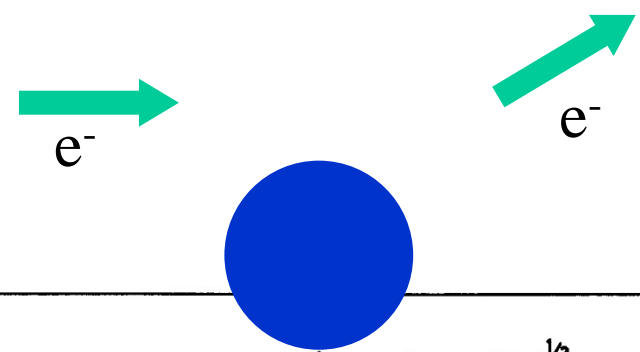
$$\begin{aligned} \frac{d\sigma}{d\Omega} &= \frac{Z_P^2 e^4}{(4E \sin^2 \theta/2)^2} |F(\mathbf{q})|^2 \\ &= \left( \frac{d\sigma_{\text{Ruth}}}{d\Omega} \right) |F(\mathbf{q})|^2 \end{aligned}$$

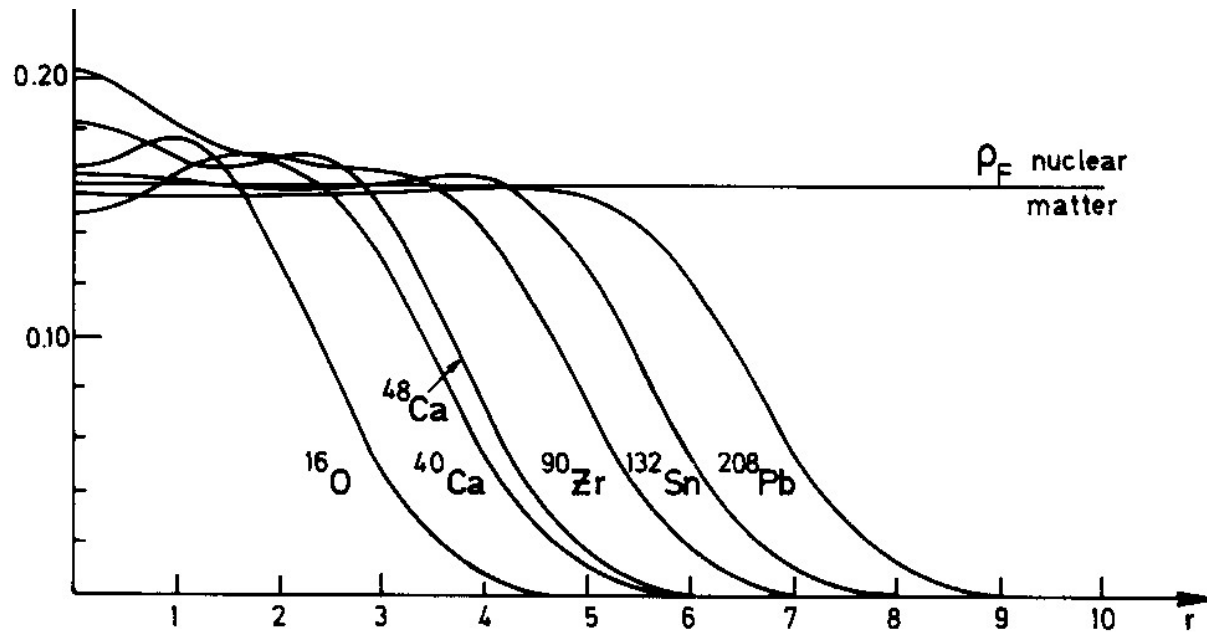
## Form factor

$$F(\mathbf{q}) = \int e^{-i\mathbf{q}\cdot\mathbf{r}} \rho(\mathbf{r}) d\mathbf{r}$$

\* relativistic correction:

$$\begin{aligned} \frac{d\sigma_{\text{Ruth}}}{d\Omega} &\rightarrow \frac{d\sigma_{\text{Mott}}}{d\Omega} \\ &= \frac{d\sigma_{\text{Ruth}}}{d\Omega} \cdot \left( 1 - \frac{v^2}{c^2} \sin^2 \frac{\theta}{2} \right) \\ &\sim \frac{d\sigma_{\text{Ruth}}}{d\Omega} \cdot \cos^2 \frac{\theta}{2} \quad (v \rightarrow c) \end{aligned}$$





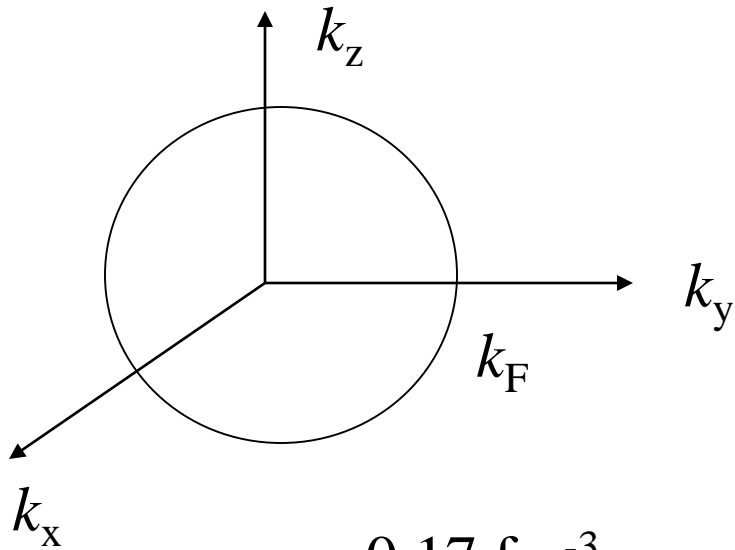
Fermi distribution

$$\rho(r) = \frac{\rho_0}{1 + \exp[(r - R_0)/a]}$$

$$\begin{aligned} \rho_0 &\sim 0.17 \text{ (fm}^{-3}\text{)} && \leftarrow \text{Saturation property} \\ R_0 &\sim 1.1 \times A^{1/3} \text{ (fm)} \\ a &\sim 0.57 \text{ (fm)} \end{aligned}$$

# Momentum Distribution

Fermi gas approximation



$$\rho = \underbrace{2 \times 2 \times 4\pi}_{\substack{\uparrow\downarrow \\ \text{spin-isospin degeneracy}}} \int_0^{k_F} \frac{k^2 dk}{(2\pi)^3}$$

$$= \frac{2}{3\pi^2} k_F^3$$

(note: spin-isospin degeneracy)

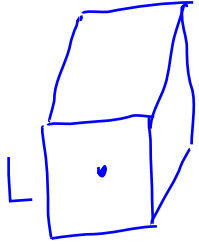
$$\rho = 0.17 \text{ fm}^{-3} \longrightarrow k_F \sim 1.36 \text{ fm}^{-1}$$

$$\longleftrightarrow \frac{v_F}{c} = \frac{k_F \cdot \hbar c}{mc^2} = 0.285$$

*p = \hbar k = mv*

$$\text{Fermi energy: } \epsilon_F = \frac{k_F^2 \hbar^2}{2m} \sim 37 \text{ (MeV)}$$

*= p^2*



free particle  $\psi(\vec{r}) = \frac{1}{\sqrt{L^3}} e^{i\vec{k} \cdot \vec{r}}$

by B.C  $k_x = \frac{2\pi}{L} n_x, k_y = \frac{2\pi}{L} n_y, k_z = \frac{2\pi}{L} n_z$

$$n_x = \frac{L}{2\pi} k_x$$

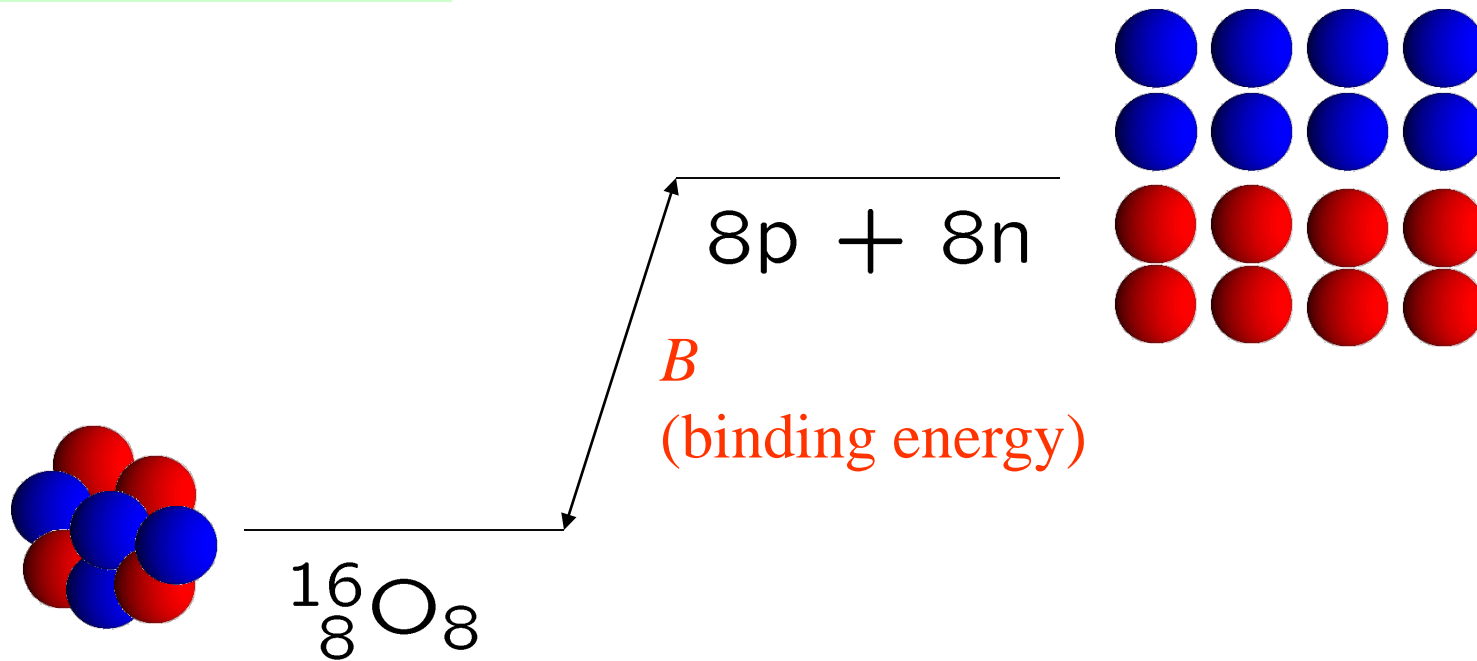
$$dn = \left(\frac{L}{2\pi}\right)^3 d^3 k$$

$$A = \int_0^{\epsilon_F} dn$$

$$\# \text{ of state: } \rho = \frac{A}{L^3} = \int_0^{k_F} \frac{d^3 k}{(2\pi)^3} = 4\pi \int_0^{k_F} \frac{k^2 dk}{(2\pi)^3}$$

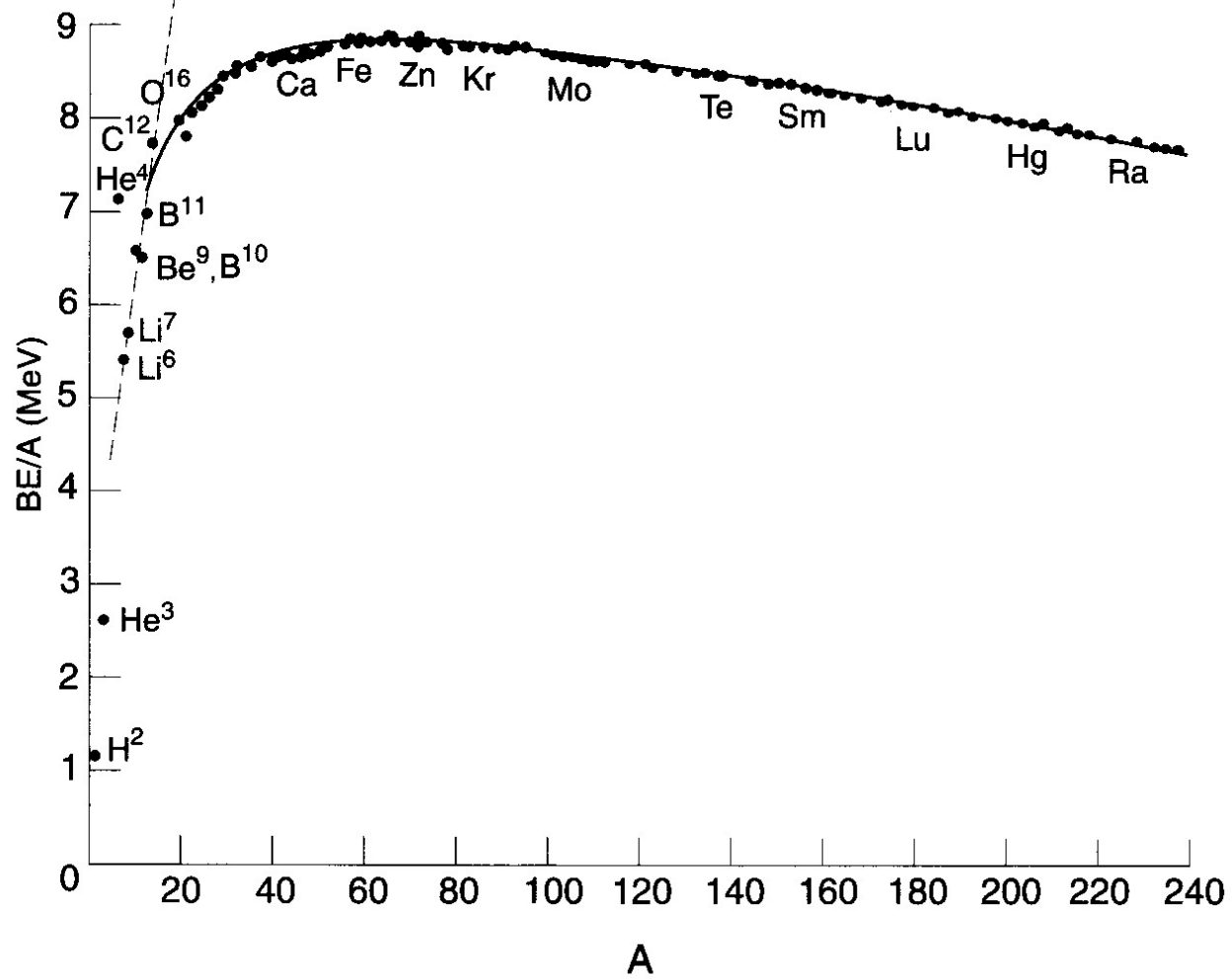


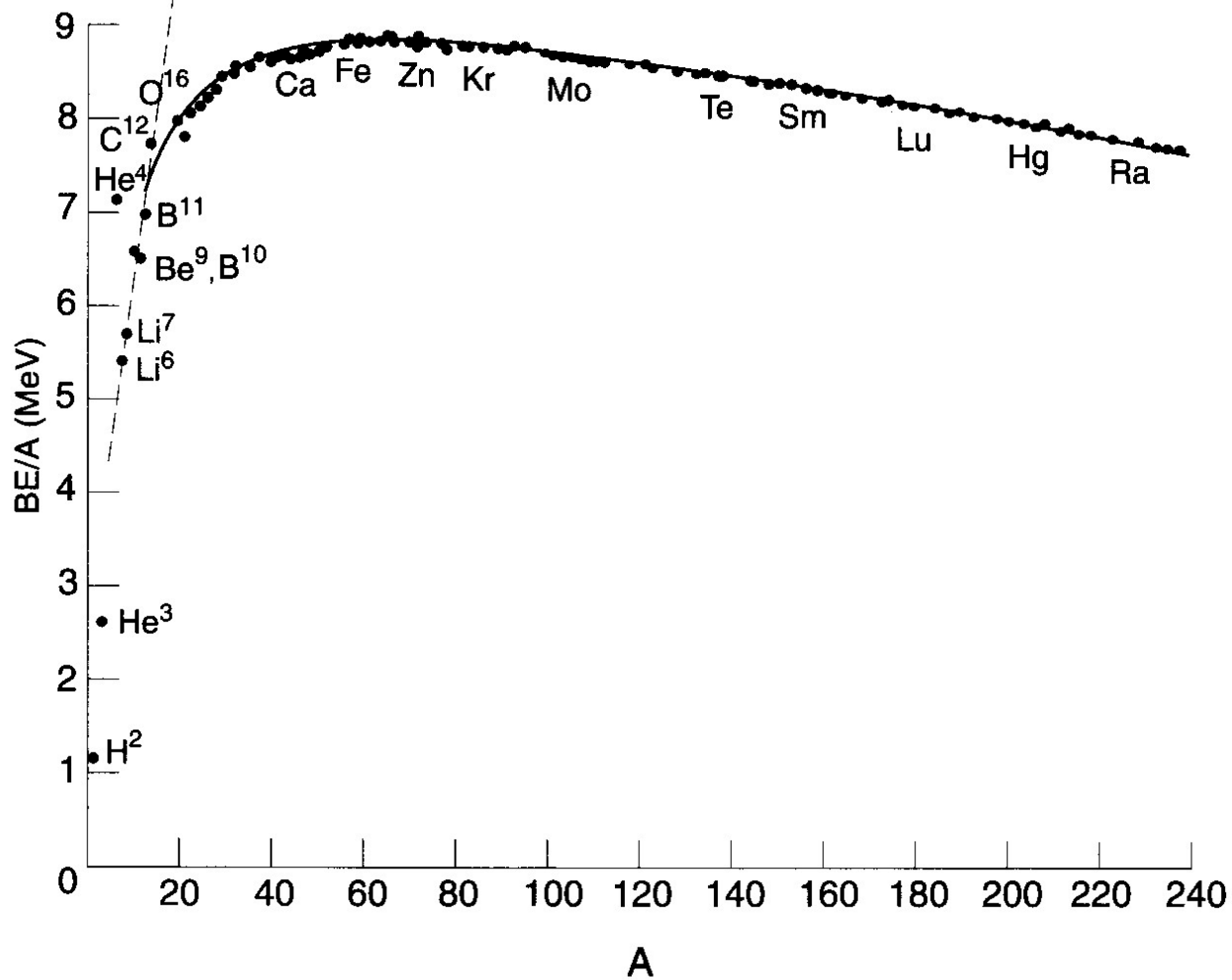
# Nuclear Mass



$$m(N, Z)c^2 = Zm_p c^2 + Nm_n c^2 - B$$

↑  
binding E  
 $B > 0$





1.  $B(N,Z)/A \sim 8.5 \text{ MeV} (A > 12)$

$\iff$  Short range nuclear force

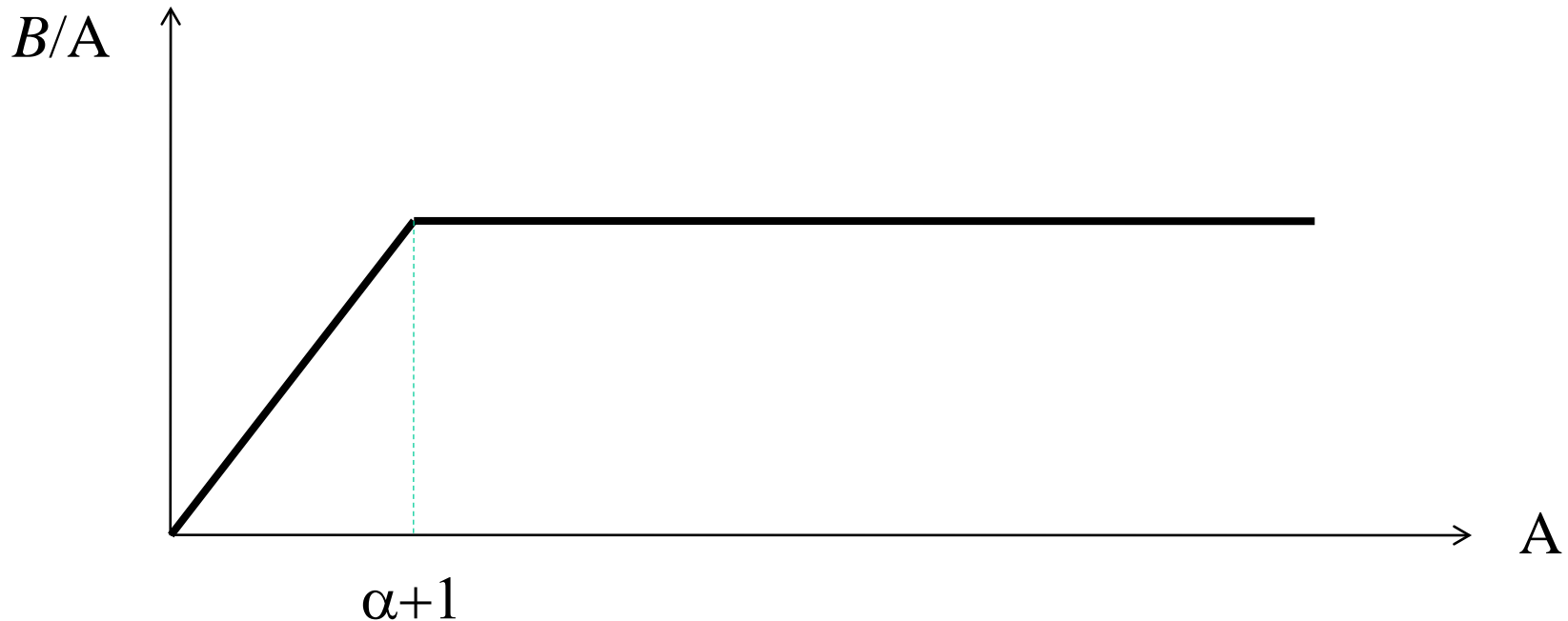
핵자는 거위의 뼚뼚 이웃하지만 상호작용함

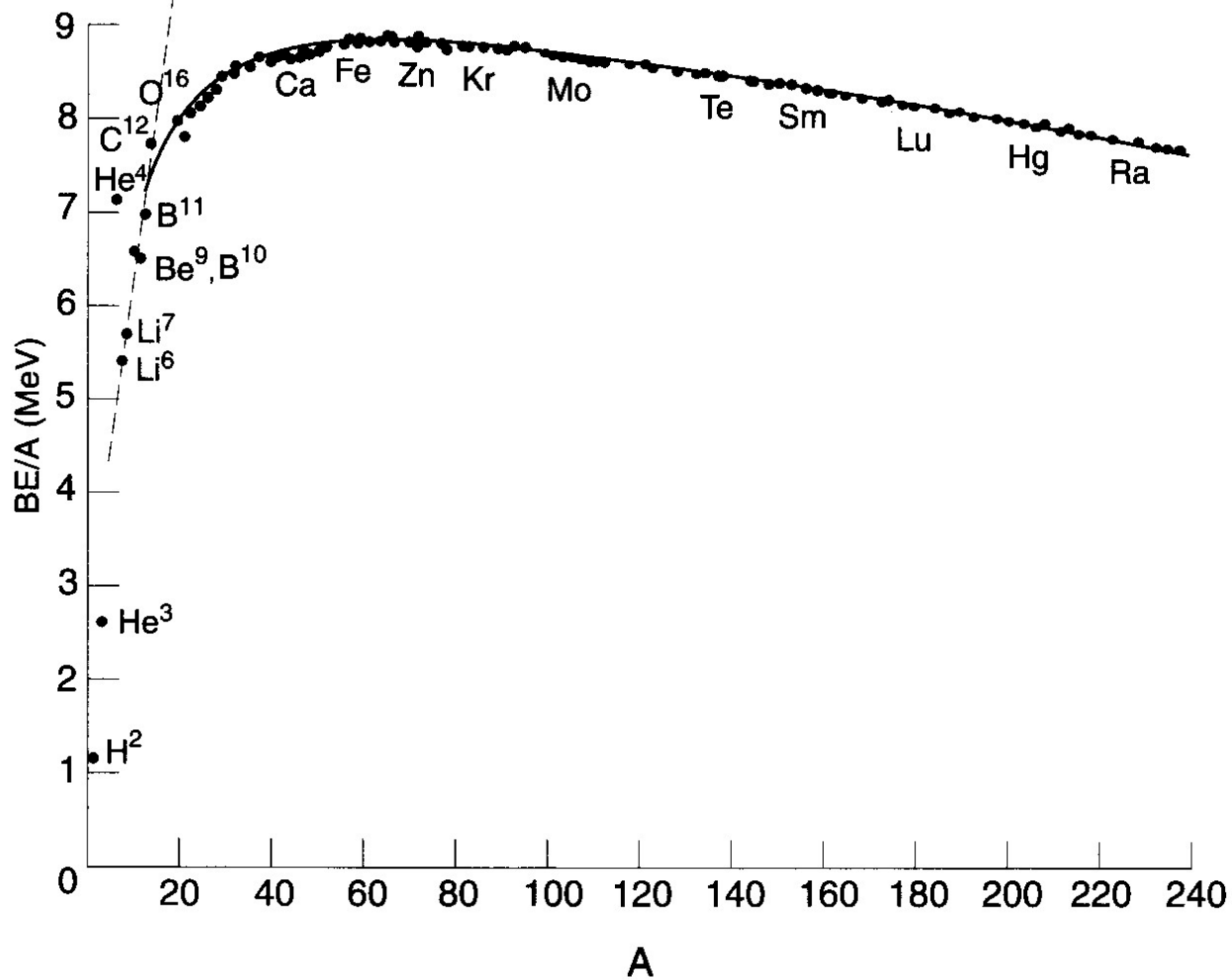
If one nucleon interacts only with surrounding  $\alpha$  nucleons

$$B \sim \alpha A/2 \longrightarrow B/A \sim \alpha/2 \text{ (const.)}$$

For  $A < \alpha+1$ , one nucleon interacts with all the other nucleons

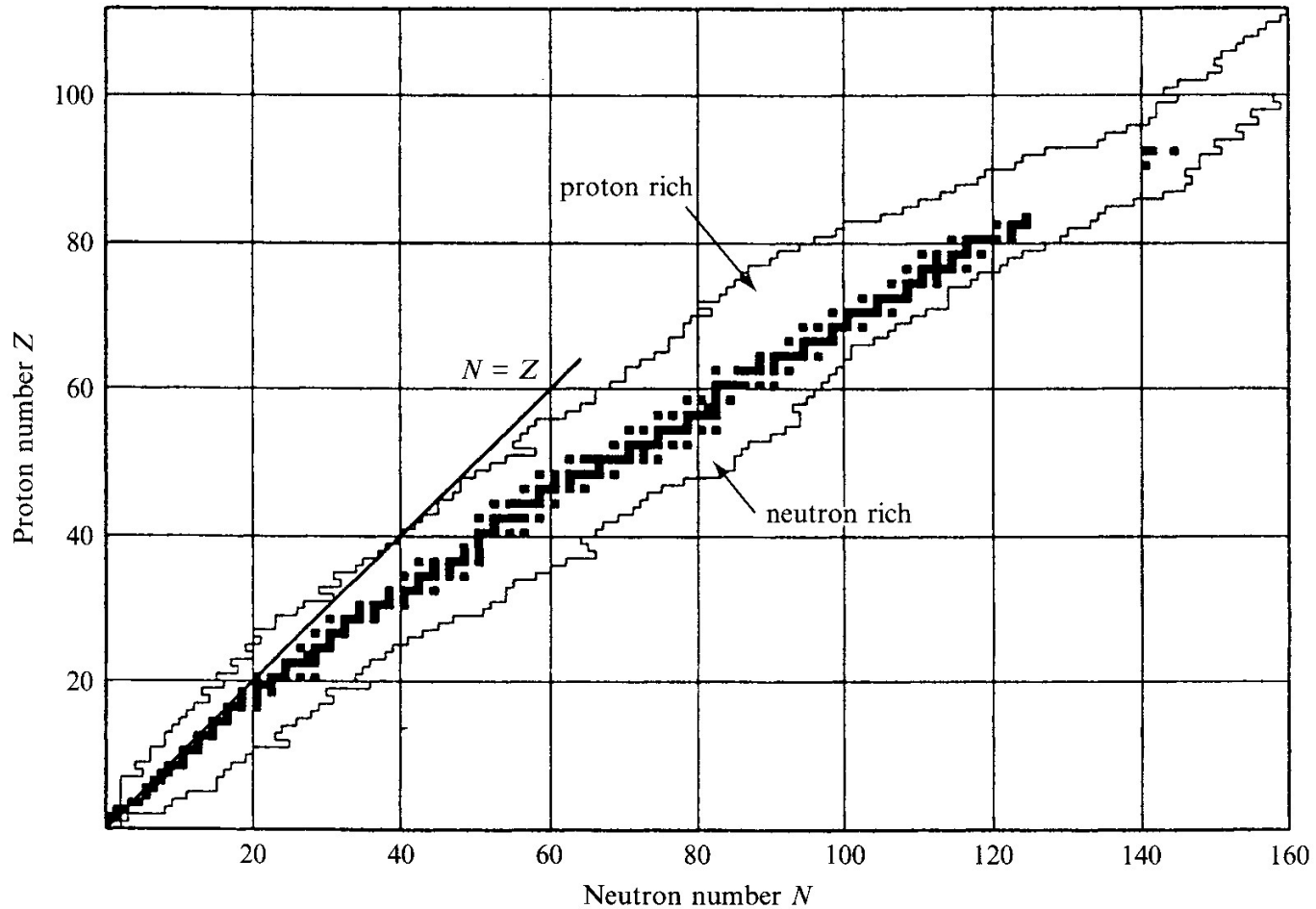
$$\longrightarrow B/A \propto A$$



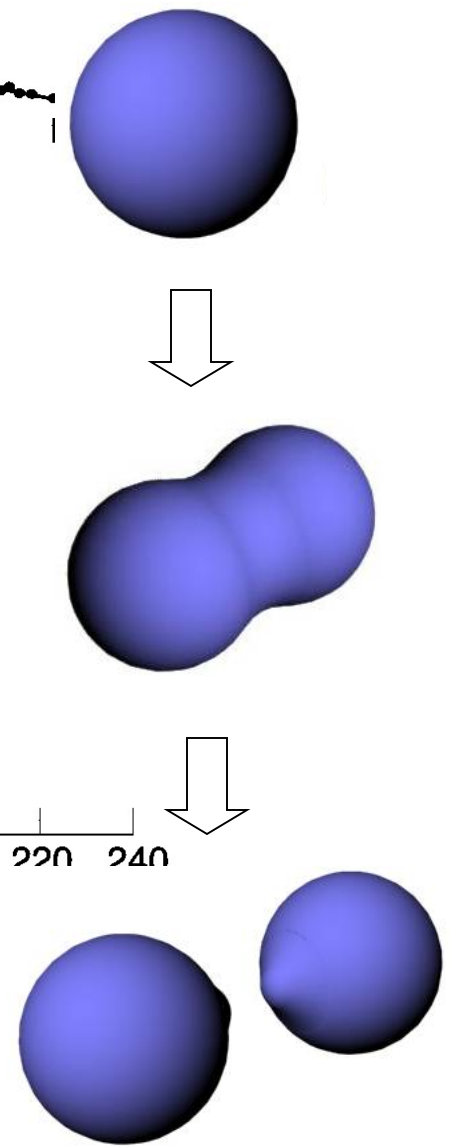
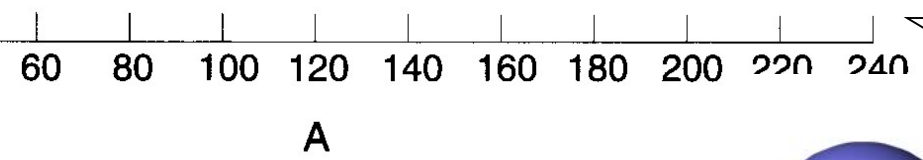
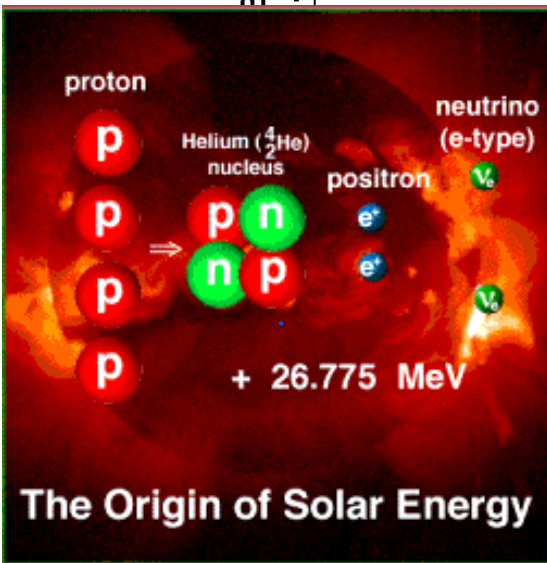
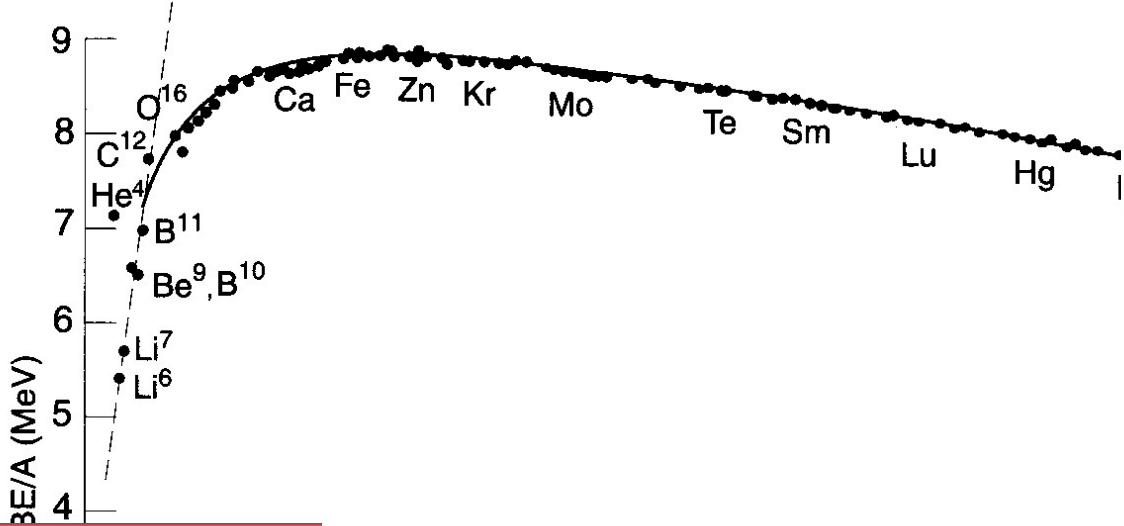


1.  $B(N,Z)/A \sim 8.5 \text{ MeV} (A > 12) \iff$  Short range nuclear force
2. Effect of Coulomb force for heavy nuclei

# Nuclear Chart



Stable nuclei:  $N \geq Z$



1.  $B(A, Z)/A \approx 0.8 \text{ MeV}$  ( $A > 12$ )  $\iff$  Short range
2. Effect of Coulomb force for heavy nuclei
3. Fusion for light nuclei
4. Fission for heavy nuclei

# Semi-empirical mass formula

$$R = r_0 A^{1/3}$$

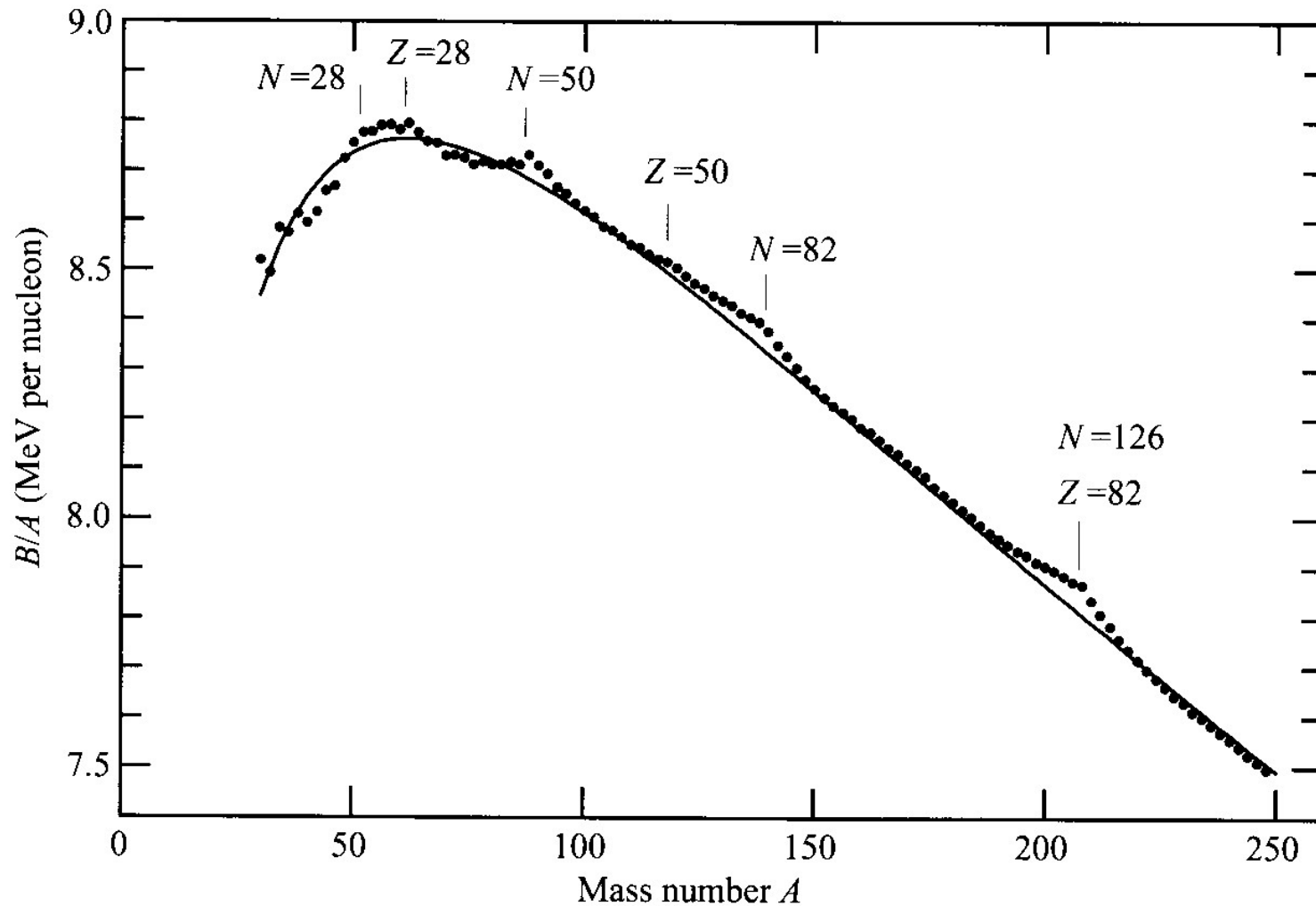
(Bethe-Weizacker formula: Liquid-drop model)

$$B(N, Z) = a_v A - a_s A^{2/3} - a_C \frac{Z^2}{A^{1/3}} - a_{\text{sym}} \frac{(N - Z)^2}{A}$$

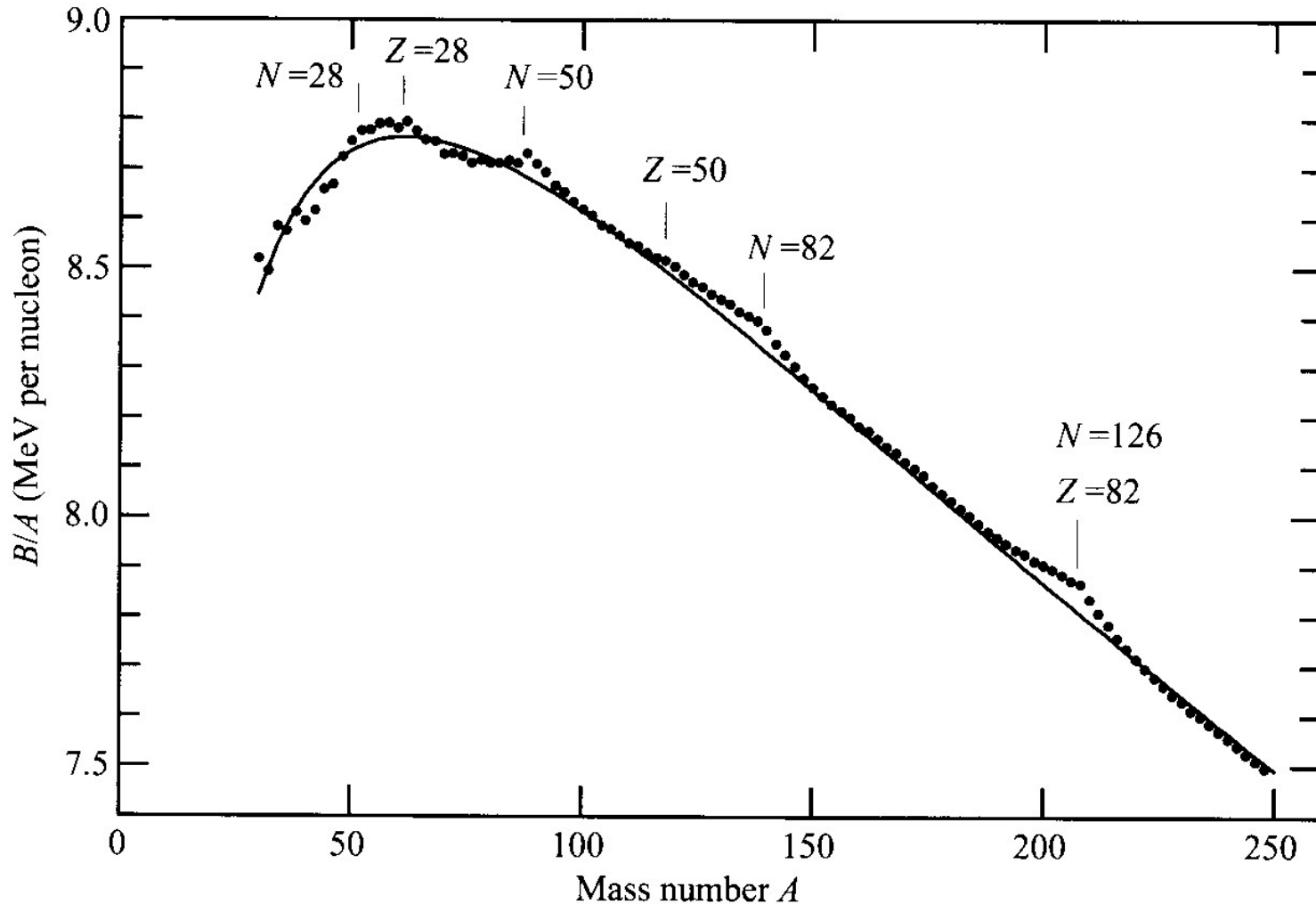
- Volume energy:  $a_v A$
- Surface energy:  $-a_s A^{2/3}$
- Coulomb energy:  $-a_C Z^2 / A^{1/3}$
- Symmetry energy:  $-a_{\text{sym}} (N - Z)^2 / A$



# How well does the Bethe-Weizacker formula reproduce the data?



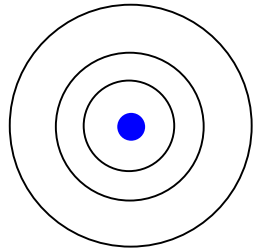
# How well does the Bethe-Weizacker formula reproduce the data?



cf.  $N, Z = 2, 8, 20, 28, 50, 82, 126$ : large binding energy  
“magic numbers”

(note) Atomic magic numbers (Noble gas)

He (Z=2), Ne (Z=10), Ar (Z=18), Kr (Z=36), Xe (Z=54), Rn (Z=86)



shell structure

元素の周期表

	1A	2A	3A	4A	5A	6A	7A	8	1B	2B	3B	4B	5B	6B	7B	0		
1	H															He		
2	Li	Be									B	C	N	O	F	Ne		
3	Na	Mg									Al	Si	P	S	Cl	Ar		
4	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
5	Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
6	Cs	Ba	L	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
7	Fr	Ra	A															
	L	La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu		
	A	Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr		

Legend:

- 典型金属元素 (Orange)
- 半金属元素 (Light Green)
- 非金属元素 (Cyan)
- 遷移金属元素 (Yellow)
- 希ガス (Pink)

Copyright © 2002 RSCS