## Scale Invariance in Newton-Cartan and Horava-Lifshitz Gravity

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# Outline

- Motivation, uses of Newton-Cartan (NC) structures
  - -holography
  - -field theory
  - -gravity
- Basic structures in NC and Torsional Newton-Cartan (TNC) from the
  - -Geometric view
  - -Gauging of the algebra
  - -Large c-expansion
- Making TNC dynamical
  - -Scale invariant Horava-Lifshitz gravity (HLG)
  - -Schrodinger invariant HLG
- Outlook

- NC introduced in problem of FQH [Son 13]
- TNC first observed as boundary geometry in z=2 Lifshitz holography [Christensen, Hartong, Rollier, Obers 13, Hartong, Kiritsis, Ober (1409)]
- TTNC introduced in FQH [Geracie,Son,Wu,Wu 14]
- TNC from gauging Schrödinger algebra [Bergshoeff,Hartong,Rosseel 14]
- TNC from gauging Bargmann (with torsion) [Hartong,Obers 15]
   coupling of non-relativistic field theories to TNC [Jensen 14]
   (independent of holography) [Hartong,Kiritsis, Obers 14]
- TNC related and 2D WCFT [Hofmann,Rollier 14]
- other approaches (c-> inf limit, affine spaces)

[Banerjee, Mitra, Mukherjee 14, Bekaert, Morand 14, Van den Bleeken 17]

## NC Geometry

 Newton-Cartan gravity, is originally developed as the generally covariant description of the Newtonian gravity [E.Cartan 1923,1924]

$$\nabla^2 \phi = 4\pi G\rho, \qquad \qquad \frac{d^2 x^a(t)}{dt^2} + \frac{\partial \phi(x)}{dx^a} = 0,$$

• Newton-Cartan geometry is described by a degenerate spatial metric  $h^{\mu\nu}$  of rank-d and a temporal vielbein  $\tau_{\mu}$  of rank-1, together with a  $\Gamma^{\mu}_{\nu\rho}$  connection on an orientable manifold M

$$h^{\mu\nu}\tau_{\nu}=0.$$

$$\begin{aligned} \nabla_{\mu}\tau_{\nu} &= \partial_{\mu}\tau_{\nu} - \Gamma^{\rho}_{\mu\nu}\tau_{\rho} &= 0, \\ \nabla_{\mu}h^{\nu\rho} &= \partial_{\mu}h^{\nu\rho} + \Gamma^{\nu}_{\sigma\mu}h^{\sigma\rho} + \Gamma^{\rho}_{\sigma\mu}h^{\sigma\nu} &= 0, \end{aligned} \xrightarrow{\qquad \qquad } \partial_{[\mu}\tau_{\nu]} = 0 \\ \tau_{\mu} &= \partial_{\mu}f \end{aligned}$$

Given any function t the temporal length of a curve  $\gamma: [s_1, s_2] \to \mathcal{M}$ 

$$\int_{s_1}^{s_2} t_{\mu} \xi^{\mu} ds = \int_{s_1}^{s_2} \nabla_{\mu} t \, \xi^{\mu} ds$$
$$= \int_{s_1}^{s_2} \frac{d}{ds} \left[ t(\gamma(s)) \right] ds = t(\gamma(s_2)) - t(\gamma(s_1)),$$

$$abla_{\mu}\tau_{\nu} = \partial_{\mu}\tau_{\nu} - \Gamma^{\rho}_{\mu\nu}\tau_{\rho} = 0, \longrightarrow \tau_{\rho}\Gamma^{\rho}_{\mu\nu} = \partial_{\mu}\tau_{\nu}.$$
 Temporal part

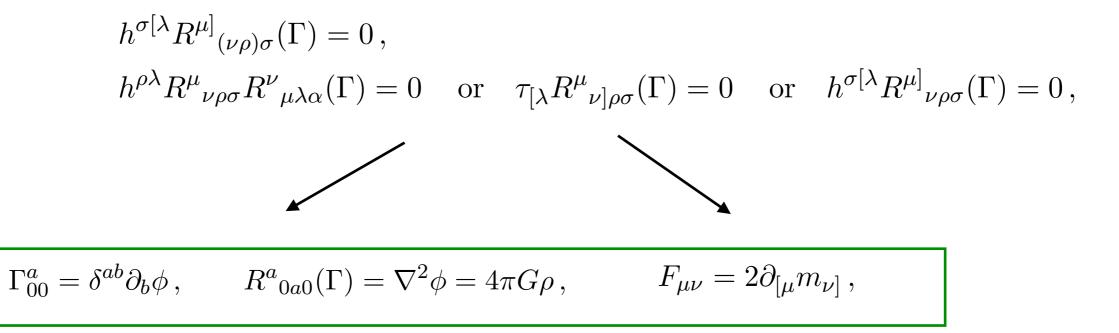
For the spatial part introduce the following spatial inverse metric  $h_{\mu\nu}$  and temporal inverse vielbein  $\tau^{\mu}$ 

$$h^{\mu\sigma}h_{\nu\sigma} = P^{\mu}_{\nu} = \delta^{\mu}_{\nu} - \tau^{\mu}\tau_{\nu}, \qquad \tau^{\mu}\tau_{\mu} = 1, \qquad h^{\mu\nu}\tau_{\nu} = 0, \qquad h_{\mu\nu}\tau^{\nu} = 0.$$

Solution to the compatibility conditions then

$$\Gamma^{\rho}_{\mu\nu} = \tau^{\rho}\partial_{\mu}\tau_{\nu} + \frac{1}{2}h^{\rho\sigma} \left(\partial_{\nu}h_{\sigma\mu} + \partial_{\mu}h_{\sigma\nu} - \partial_{\sigma}h_{\mu\nu}\right) - h^{\rho\sigma}\tau_{(\mu}F_{\nu)\sigma} ,$$

Impose Ehlers and Trautman conditions for Newtonian gravity



[J. Ehlers 81, A. Trautman 63, G. Dautcourt 93, R. Andringa, E. Bergshoeff, S. Panda and M. de Roo 11]

There is a caveat here ! The inverse metrics we introduce  $\tau^{\mu}$ ,  $h_{\mu\nu}$ are not unique i.e. given 1-form  $\psi_{\mu}$  [K. Jensen, 14]  $\tau'^{\mu} = \tau^{\mu} + h^{\mu\nu}\psi_{\nu}$ ,  $h'_{\mu\nu} = h_{\mu\nu} - (\tau_{\mu}P^{\rho}_{\nu} + \tau_{\nu}P^{\rho}_{\mu})\psi_{\rho} + \tau_{\mu}\tau_{\nu}h^{\rho\sigma}\psi_{\rho}\psi_{\sigma}$ , Moreover,  $m'_{\mu} = m_{\mu} - P^{\nu}_{\mu}\psi_{\nu} + \frac{1}{2}\tau_{\mu}h^{\nu\rho}\psi_{\nu}\psi_{\rho}$ ,

#### **Twistless torsional NC (TTNC) geometry**

 $(h^{\mu\nu}, \tau_{\mu}, b_{\mu}, M_{\mu}),$ 

From the compatibility condition temporal part is fixed

$$\tau_{\rho}\Gamma^{\rho}_{[\mu\nu]} = \partial_{[\mu}\tau_{\nu]} \,.$$

Impose twistless condition

$$\tau_{\lambda}\tau_{[\rho}\Gamma^{\lambda}_{\mu\nu]} = \tau_{[\rho}\partial_{\mu}\tau_{\nu]} = 0\,,$$

By Frobenius theorem

$$\partial_{[\mu}\tau_{\nu]} = zb_{[\mu}\tau_{\nu]}\,,$$

Solution to the compatibility conditions is spatial contortion  $\Gamma^{\rho}_{\mu\nu} = \tau^{\rho}\partial_{\mu}\tau_{\nu} + \frac{1}{2}h^{\rho\sigma}\left(\partial_{\nu}h_{\sigma\mu} + \partial_{\mu}h_{\sigma\nu} - \partial_{\sigma}h_{\mu\nu}\right) - h^{\rho\sigma}\tau_{(\mu}F_{\nu)\sigma} - K_{\mu\nu}{}^{\rho},$ 

Transformation under Milne boost

$$\delta_{M}\Gamma^{\rho}_{\mu\nu} = h^{\rho\sigma} \left\{ (P^{\alpha}_{\sigma}\partial_{[\mu}\tau_{\nu]} + P^{\alpha}_{\mu}\partial_{[\sigma}\tau_{\nu]} + P^{\alpha}_{\nu}\partial_{[\sigma}\tau_{\mu]})\psi_{\alpha} + \frac{1}{2}h^{\alpha\beta}\psi_{\alpha}\psi_{\beta}(\tau_{\nu}\partial_{[\mu}\tau_{\sigma]} + \tau_{\mu}\partial_{[\nu}\tau_{\sigma]}) \right\}$$

$$(2)$$

• If  $\delta_M K_{\mu\nu}{}^{\rho} = 0$ , then we have to achieve Milne invariance by considering combinations of  $\tau^{\mu}$ ,  $h_{\mu\nu}$  and  $m_{\mu}$ . However  $m_{\mu}$  is a U(1) connection therefore such combinations will fail U(1) invariance. So we will also introduce a scalar field  $\chi$  such that,

$$\delta_{\mathrm{U}(1)}\chi = \sigma \qquad \qquad M_{\mu} = m_{\mu} - \partial_{\mu}\chi,$$

And the following definitions will still satisfy orthogonality conditions and they are Milne+U(1) invariant

$$\hat{\tau}^{\mu} = \tau^{\mu} + h^{\mu\nu} M_{\nu}, \qquad \hat{h}_{\mu\nu} = h_{\mu\nu} - \tau_{\mu} M_{\nu} - \tau_{\nu} M_{\mu} + 2\tau_{\mu} \tau_{\nu} \Phi,$$

$$\Phi = \tau^{\sigma} M_{\sigma} + \frac{1}{2} h^{\rho \sigma} M_{\rho} M_{\sigma} \,.$$

Therefore, the connection for TTNC can be written as

$$\widehat{\Gamma}^{\rho}_{\mu\nu} = \widehat{\tau}^{\rho}\partial_{\mu}\tau_{\nu} + \frac{1}{2}h^{\rho\sigma} \Big(\partial_{\nu}\widehat{h}_{\sigma\mu} + \partial_{\mu}\widehat{h}_{\sigma\nu} - \partial_{\sigma}\widehat{h}_{\mu\nu}\Big) + h^{\rho\sigma}\tau_{\mu}\tau_{\nu}\partial_{\sigma}\Phi - K_{\mu\nu}{}^{\rho},$$

#### Non-Relativistic scale symmetry and NC

The defining property of the scale symmetry is via breaking of the compatibility condition
 by a particular non-metricity tensor

$$\nabla_{\mu}\tau_{\nu} = zb_{\mu}\tau_{\nu}, \qquad \nabla_{\mu}h^{\nu\rho} = -2b_{\mu}h^{\nu\rho},$$

which is preserved by the following transformations

 $\tau_{\mu} \to e^{z\Lambda_D(x)}\tau_{\mu} , \qquad h^{\mu\nu} \to e^{-2\Lambda_D(x)}h^{\mu\nu} , \qquad b_{\mu} \to b_{\mu} + \partial_{\mu}\Lambda_D(x) , \qquad \Gamma^{\rho}_{\mu\nu} \to \Gamma^{\rho}_{\mu\nu} , \qquad \delta_M b_{\mu} = 0 .$ 

 Unlike the relativistic scenarios, the inclusion of the non-metricity modifies the antisymmetric part of the connection

#### impose tt condition

$$\tau_{\rho}\Gamma^{\rho}_{[\mu\nu]} = \partial_{[\mu}\tau_{\nu]} - zb_{[\mu}\tau_{\nu]} \cdot \longrightarrow \partial_{[\mu}\tau_{\nu]} = zb_{[\mu}\tau_{\nu]} ,$$

•Again, one can introduce Milne invariant hatted quantities

$$\hat{\tau}^{\mu} \to e^{-z\Lambda_D(x)}\hat{\tau}^{\mu}, \quad \hat{h}_{\mu\nu} \to e^{2\Lambda_D(x)}\hat{h}_{\mu\nu}, \quad M_{\mu} \to e^{-(z-2)\Lambda_D(x)}M_{\mu}.$$
  
$$M_{\mu} = m_{\mu} - \partial_{\mu}\chi - (z-2)b_{\mu}\chi,$$

$$\hat{\Gamma}^{\rho}_{\mu\nu} = \hat{\tau}^{\rho} \mathcal{D}_{\mu} \tau_{\nu} + \frac{1}{2} h^{\rho\sigma} \Big( \mathcal{D}_{\nu} \hat{h}_{\sigma\mu} + \mathcal{D}_{\mu} \hat{h}_{\sigma\nu} - \mathcal{D}_{\sigma} \hat{h}_{\mu\nu} \Big) + h^{\rho\sigma} \tau_{\mu} \tau_{\nu} \mathcal{D}_{\sigma} \Phi - K_{\mu\nu}{}^{\rho}$$

 $\begin{array}{c} [H,G_a] = P_a & [P_a,G_b] = 0 \\ [II,G_a] = -I_a & [II_a,G_b] = 0 \\ [II,G_a] = -I_a & [II_a,G_b] = 0 \\ \\ Barg[Hafna] = P_a & [H,G_a] = P_a & [P_a,G_b] = 0 \\ H,P_a,J_{ab},G_a & H,P_a,J_{ab},G_{ab} & N \end{array}$ Non-relativistic algebras (not all)  $[P_a, G_b] = N\delta_{ab}$  $\begin{bmatrix} I & a, \Box_b \end{bmatrix} = N \quad \begin{bmatrix} P_a, G_b \end{bmatrix} = N\delta_{ab}$  $\begin{bmatrix} D, H \end{bmatrix} = zH \quad \begin{bmatrix} D, P_a \end{bmatrix} = P_a \\ H, P_a, J_{ab}, D, \quad G_a, N, K(z = 2) \begin{bmatrix} D & H & \dots & \dots \\ D, H \end{bmatrix} = zE \quad \begin{bmatrix} D, H \end{bmatrix} = zH \quad \begin{bmatrix} D, P_a \end{bmatrix} = zH \quad \begin{bmatrix} D, P_a \end{bmatrix} = zH \quad \begin{bmatrix} D, P_a \end{bmatrix} = a \\ H, P_a, J_{ab}, \quad H, P_a, J_{ab}, D, \quad G_a, N, K(z = b \\ - b & D, N \end{bmatrix} = (2 - z)N$  Lifshitz  $[D, N] = (2 \quad [D, N] = (2 - z)N$ Schrodinger [D, IV] = (2 - 2)IV

#### $z \neq 2$ Scale Invariant Hořava-Lifshitz Gravity

Elements of the scale extended Bargmann algebra are

$$[P_a, G_b] = \delta_{ab}N, \qquad [D, G_a] = (z - 1)G_a, \qquad [D, N] = (z - 2)N,$$
$$[J_{ab}, P_c] = 2\delta_{c[a}P_{b]}, \qquad [J_{ab}, G_c] = 2\delta_{c[a}G_{b]}, \qquad [J_{ab}, J_{cd}] = 4\delta_{[a[c}J_{b]d]}.$$

$$[D, P_a] = -P_a$$
,  $[D, H] = -zH$ ,  $[H, G_a] = P_a$ ,

Following the algebra, transformation rules are given as

$$\begin{split} \delta\tau_{\mu} &= \partial_{\mu}\xi - z\xi b_{\mu} + z\Lambda_{D}\tau_{\mu} \,, \\ \delta e_{\mu}{}^{a} &= \partial_{\mu}\xi^{a} - \omega_{\mu}{}^{ab}\xi_{b} - b_{\mu}\xi^{a} + \lambda^{a}{}_{b}e_{\mu}{}^{b} + \lambda^{a}\tau_{\mu} - \omega_{\mu}{}^{a}\xi + \Lambda_{D}e_{\mu}{}^{a} \,, \\ \delta\omega_{\mu}{}^{ab} &= \partial_{\mu}\lambda^{ab} + 2\lambda^{c[a}\omega_{\mu}{}^{b]}{}_{c} \,, \\ \delta\omega_{\mu}{}^{a} &= \partial_{\mu}\lambda^{a} - \omega_{\mu}{}^{ab}\lambda_{b} + \lambda^{a}{}_{b}\omega_{\mu}{}^{b} + (z-1)\lambda^{a}b_{\mu} - (z-1)\Lambda_{D}\omega_{\mu}{}^{a} \,, \\ \delta m_{\mu} &= \partial_{\mu}\sigma - \xi^{a}\omega_{\mu a} + \lambda^{a}e_{\mu a} + (z-2)\sigma b_{\mu} - (z-2)\Lambda_{D}m_{\mu} \,, \\ \delta b_{\mu} &= \partial_{\mu}\Lambda_{D} \,, \end{split}$$

Also the curvatures follows easily

$$\begin{aligned} R_{\mu\nu}(H) &= 2\partial_{[\mu}\tau_{\nu]} - 2zb_{[\mu}\tau_{\nu]} \,, \\ R_{\mu\nu}{}^{a}(P) &= 2\partial_{[\mu}e_{\nu]}{}^{a} - 2\omega_{[\mu}{}^{ab}e_{\nu]b} - 2\omega_{[\mu}{}^{a}\tau_{\nu]} - 2b_{[\mu}e_{\nu]}{}^{a} \,, \\ R_{\mu\nu}{}^{ab}(J) &= 2\partial_{[\mu}\omega_{\nu]}{}^{ab} - 2\omega_{[\mu}{}^{c[a}\omega_{\nu]}{}^{b]}{}_{c} \,, \\ R_{\mu\nu}{}^{a}(G) &= 2\partial_{[\mu}\omega_{\nu]}{}^{a} + 2\omega_{[\mu}{}^{b}\omega_{\nu]}{}^{a}{}_{b} - 2(z-1)\omega_{[\mu}{}^{a}b_{\nu]} \,, \\ R_{\mu\nu}(D) &= 2\partial_{[\mu}b_{\nu]} \,, \\ R_{\mu\nu}(N) &= 2\partial_{[\mu}m_{\nu]} - 2\omega_{[\mu}{}^{a}e_{\nu]a} + 2(z-2)b_{[\mu}m_{\nu]} \,. \end{aligned}$$

In order to make contact with the NC geometry we have discussed, impose the following set of constraints

$$R_{\mu\nu}(H) = 0$$
,  $R_{\mu\nu}{}^{a}(P) = 0$ ,  $R_{\mu\nu}(N) = 0$ ,  $R_{\mu\nu}(D) = 0$ ,

$$e_{[\mu}{}^{b}R_{\nu\rho]}{}^{a}{}_{b}(J) + \tau_{[\mu}R_{\nu\rho]}{}^{a}(G) = 0, \qquad e_{[\mu}{}^{a}R_{\nu\rho]a}(G) = 0.$$

The twistlessness condition can be seen from the curvature constraint

$$R_{\mu\nu}(H) = 0 \qquad \Rightarrow \qquad \partial_{[\mu}\tau_{\nu]} = zb_{[\mu}\tau_{\nu]},$$

#### Also the following composite gauge fields are solved

$$\begin{split} \omega_{\mu}{}^{ab} &= -2e^{\nu[a}\partial_{[\mu}e_{\nu]}{}^{b]} + e^{\nu[a}e^{-b]\rho}\partial_{\nu}e_{\rho}{}^{c}e_{\mu c} + 2e_{\mu}{}^{[a}e_{\nu}{}^{b]}b^{\nu} \\ &- e^{\nu a}e^{\rho b}\tau_{\mu}(\partial_{[\nu}m_{\rho]} + (z-2)b_{[\nu}m_{\rho]}) \,, \\ \omega_{\mu}{}^{a} &= \tau^{\nu}\partial_{[\mu}e_{\nu]}{}^{a} + e^{\nu a}\tau^{\rho}e_{\mu b}\partial_{[\nu}e_{\rho]}{}^{b} + e_{\mu}{}^{a}\tau^{\nu}b_{\nu} + e^{\nu a}(\partial_{[\mu}m_{\nu]} + (z-2)b_{[\mu}m_{\nu]}) \\ &+ \tau_{\mu}\tau^{\rho}e^{\nu a}(\partial_{[\rho}m_{\nu]} + (z-2)b_{[\rho}m_{\nu]}) \,. \end{split}$$

Scale-extended Bargmann algebra does not allow the existence of a scalar field with a homogeneous dilatation and U(1) transformation due to the commutation relation.

Therefore we will work with the following set of compensating multiplets. Note that the derivative of  $\chi$  is the very definition of  $M_{\mu}$ 

$$\delta \phi = \omega \Lambda_D \phi$$
,  $\delta \chi = \sigma + (2 - z) \Lambda_D \chi$ .

$$\mathcal{D}_{\mu}\chi = \partial_{\mu}\chi - (2-z)b_{\mu}\chi - m_{\mu}.$$

In order to discuss the construction of kinetic term in HLG let us consider the following

$$\mathcal{D}_{\mu}M_{a} = \partial_{\mu}M_{a} + (z-1)b_{\mu}M_{a} - \omega_{\mu}{}^{ab}M_{b} - \omega_{\mu a}.$$

 $\delta \mathcal{D}_a M_b = -z \Lambda_D \mathcal{D}_a M_b \,. \qquad \delta_G M_a = \lambda_a \,.$ 

$$\begin{aligned} S_{z\neq2}^{(2)} &= \int dt \, d^d x \, e \, \phi \, (\mathcal{D}_a M^a)^2, \\ S_{z\neq2}^{(3)} &= \int dt \, d^d x \, e \, \phi \, (\mathcal{D}_a M_b)^2, \end{aligned} \qquad \delta_D e = (d+z) \Lambda_D \, e \, . \\ S_{z\neq2}^{(4)} &= \int dt \, d^d x \, e \, \phi \, (\mathcal{D}_a M^a), \end{aligned}$$

Here, for  $S_{z\neq2}^{(2)}$  and  $S_{z\neq2}^{(3)}$ , the scaling dimension of  $\phi$  is given by  $\omega = z - d$  while for the  $S_{z\neq2}^{(4)}$ , the scaling dimension of  $\phi$  is  $\omega = 2 - d$ .

•  $n_t = 1$ : When we have a single time derivative acting on  $\phi$  or  $\chi$ , we first consider the models with no spatial derivatives

$$S_{z\neq2}^{(5)} = \int dt \, d^d x \, e \, \phi(\mathcal{D}_0 \phi + M^a \mathcal{D}_a \phi) \,,$$
  
$$S_{z\neq2}^{(6)} = \int dt \, d^d x \, e \, \phi^2(M_0 + \frac{1}{2}M_a M^a) \,.$$

•  $n_t = 2$ : When we have two time derivative acting on  $\phi$  or  $\chi$ , the possible  $z \neq 2$  scaleinvariant actions are

$$S_{z\neq2}^{(13)} = \int dt \, d^d x \, e \, \phi \left( \mathcal{D}_0^2 \phi + 2M^a \mathcal{D}_0 \mathcal{D}_a \phi + M^a M^b \mathcal{D}_a \mathcal{D}_b \phi \right),$$
  

$$S_{z\neq2}^{(14)} = \int dt \, d^d x \, e \, \phi^2 \left( \mathcal{D}_0 M_0 + 2M^a \mathcal{D}_0 M_a + M^a M^b \mathcal{D}_a M_b \right),$$
  

$$S_{z\neq2}^{(15)} = \int dt \, d^d x \, e \, \phi^2 \left( M_0 + \frac{1}{2} M_a M^a \right)^2.$$

### NC and HLG connection [Hartong, Obers 2015]

Milne invariant Riemannian tensor can be defined as

$$g_{\mu\nu} = \hat{h}_{\mu\nu} + \tau_{\mu}\tau_{\nu}, \qquad g^{\mu\nu} = h^{\mu\nu} + \hat{\tau}^{\mu}\hat{\tau}^{\nu},$$

From this definition, ADM decomposition of the metric and the twistless torsion condition one can show the following

$$\tau_{\mu} = \psi \partial_{\mu} \tau \,. \qquad \tau = t$$

$$\begin{split} \tau_t &= N \,, \quad \tau_i = 0 \,, \\ h^{tt} &= h^{ti} = h^{it} = 0 \,, \quad h^{ij} = \gamma^{ij} \,, \\ \hat{\tau}^t &= N^{-1} \,, \quad \hat{\tau}^i = -N^{-1}N^i \,, \\ \hat{h}_{tt} &= \gamma_{ij}N^iN^j \,, \quad \hat{h}_{ti} = \hat{h}_{it} = \gamma_{ij}N^j \,, \quad \hat{h}_{ij} = \gamma_{ij} \,, \end{split}$$

Using these relations U(1) invariant vector field reads

$$M_t = -\frac{1}{2N}\gamma_{ij}N^iN^j + \Phi N, \qquad M_i = -\frac{1}{N}\gamma_{ij}N^j$$

The dilatation gauge field can also be written in terms of HLG variables through the TT condition

$$\partial_{[\mu}\tau_{\nu]} = zb_{[\mu}\tau_{\nu]} = zb_a e_{[\mu}{}^a\tau_{\nu]}$$

It is useful to identify a vector

$$a_{\mu} = \mathcal{L}_{\hat{\tau}} \tau_{\mu} = \hat{\tau}^{\nu} (\partial_{\nu} \tau_{\mu} - \partial_{\mu} \tau_{\nu}) = -z e_{\mu}{}^{a} b_{a} ,$$
$$a_{t} = N^{i} a_{i} , \qquad a_{i} = -N^{-1} \partial_{i} N .$$

Finally the extrinsic curvature (kinetic term of HLG)

$$K_{ij} = \frac{1}{2N} \left( \partial_t \gamma_{ij} - \bar{\nabla}_i N_j - \bar{\nabla}_j N_i \right),$$

Can be identified from the following tensor

$$K'_{ab} = \mathcal{D}_a M_b = \widetilde{\nabla}_{(a} M_{b)} + z b_{(a} M_{b)} - \delta_{ab} b^c M_c - \delta_{ab} b_0.$$

Gathering up all these pieces the scale extended HLG is

$$S_{z\neq 2}^{\text{HL}} = S_{z\neq 2}^{(3)} - \lambda S_{z\neq 2}^{(2)} + S_{\mathcal{V}},$$

# Outlook

- The scale or Schrodinger symmetry corresponds to a special choice of nonmetricity in the compatibility equation, and it is possible to have a more general classification of non-relativistic geometries by imposing a more general nonmetricity. This classification has been done for the relativistic scenarios.
- The true non-relativistic analogue of the relativistic conformal symmetry, which leaves the action of a massless non-relativistic particle is invariant is called the Galilean conformal algebra.
- Is it possible to get HLG from the large c expansion by [vDB17]
- Generating solutions for HLG from the known relativistic ones.

## Large c expansion [VdB17]

We assume an expansion of the metric (D = d + 1, Lorentzian) in even powers<sup>2</sup> of a variable c (thought of physically as the speed of light):

$$g_{\mu\nu} = \sum_{i=-1}^{\infty} {}^{(2i)}_{g\mu\nu} c^{-2i} \qquad g^{\mu\nu} = \sum_{i=0}^{\infty} {}^{(2i)}_{g\mu\nu} c^{-2i}$$
(1)

We furthermore assume that  $g_{\mu\nu}^{(-2)}$  is of rank 1 and negative, so we can write

$$g_{\mu\nu}^{(-2)} = -\tau_{\mu}\tau_{\nu}$$
 (2)

The two expansions (1) are of course related by the condition that one series provides the inverse of the other. We can expand this condition  $g_{\mu\rho}g^{\rho\nu} = \delta^{\nu}_{\mu}$  order by order and solve the resulting equations explicitly. As a first step one obtains from the leading equation (order  $c^2$ ) the result that

$$g^{(0)}\mu\nu = h^{\mu\nu}$$
 with  $h^{\mu\nu}\tau_{\nu} = 0$  (6)

$$\delta_{\chi}\tau^{\mu} = -h^{\mu\rho}\chi_{\rho} \qquad \delta_{\chi}h_{\mu\nu} = \tau_{\mu}\chi_{\nu} + \tau_{\nu}\chi_{\mu}$$

$$\begin{aligned} \hat{\tau}^{\mu} &= \tau^{\mu} - h^{\mu\nu}C_{\nu} \\ \hat{h}_{\mu\nu} &= h_{\mu\nu} + 2\tau_{(\mu}C_{\nu)} + 2\hat{\Phi}\tau_{\mu}\tau_{\nu} \\ \hat{\Phi} &= -\tau^{\rho}C_{\rho} + \frac{1}{2}h^{\rho\sigma}C_{\rho}C_{\sigma} \\ \hat{\beta}^{\mu\nu} &= \beta^{\mu\nu} + h^{\mu\rho}h^{\mu\sigma}C_{\rho}C_{\sigma} \\ \hat{B}_{\mu} &= B_{\mu} + h_{\mu\rho}\beta^{\rho\sigma}C_{\sigma} + \frac{1}{2}\tau_{\mu}\left(\beta^{\rho\sigma}C_{\rho}C_{\sigma} + (\tau^{\rho}C_{\rho} - h^{\rho\sigma}C_{\rho}C_{\sigma})^{2}\right) \\ &-C_{\mu}(\tau^{\rho}C_{\rho} - h^{\rho\sigma}C_{\rho}C_{\sigma}) \\ \hat{\gamma}^{\mu\nu} &= \gamma^{\mu\nu} + 2h^{\rho(\mu}h^{\nu)\sigma}C_{\rho}\left(\hat{B}_{\sigma} - 2\hat{\Phi}C_{\sigma}\right) \end{aligned}$$

For the compatibility condition assume the connection expansion

$$\Gamma^{\lambda}_{\mu\nu} = \sum_{i=-1}^{\infty} \Gamma^{(2i)}_{\mu\nu} c^{-2i}$$

Then the same structure for TTNC is obtained.