

# Ghost-Free Theory with Third-Order Time Derivatives

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1804.07990, to appear in JHEP.

$$c = \hbar = M_G^2 = 1/(8\pi G) = 1$$

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- **Introduction**

  - What is the most general scalar-tensor theory ?

- **Healthy degenerate theory for point particles**

  - Review of second order (time) derivative system

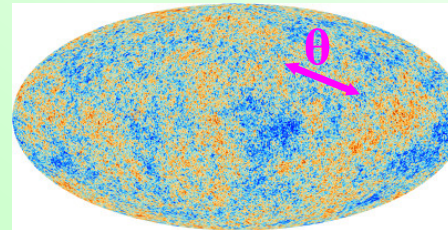
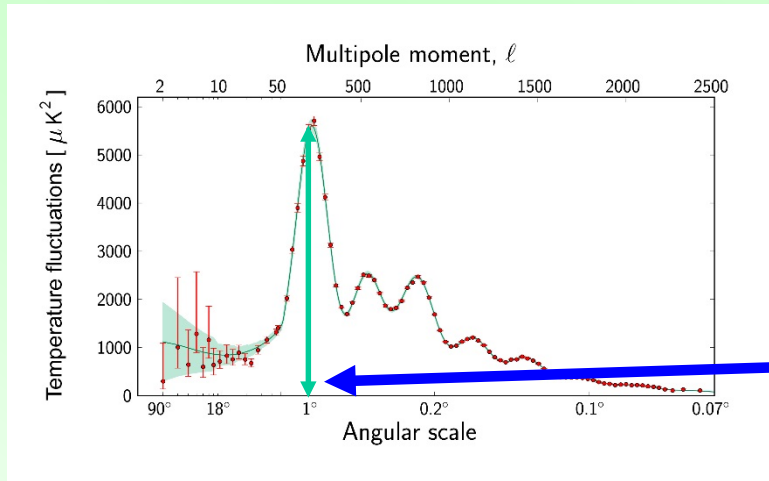
  - Third order (time) derivative system

- **Discussion and conclusions**

# Introduction

# Inflation is strongly supported by CMB observations

## Planck TT correlation :

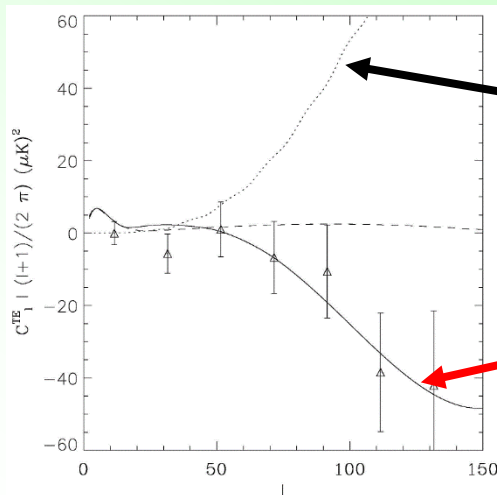


Green line : prediction by inflation  
Red points : observation by PLANCK

Angle  $\theta \sim 180^\circ / \ell$

Total energy density  $\leftrightarrow$  Geometry of our Universe

## WMAP TE correlation :



Our Universe is spatially flat as predicted by inflation !!

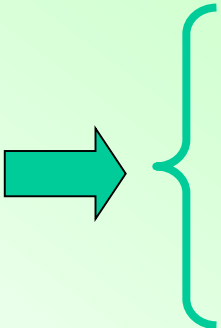
Causal seed models

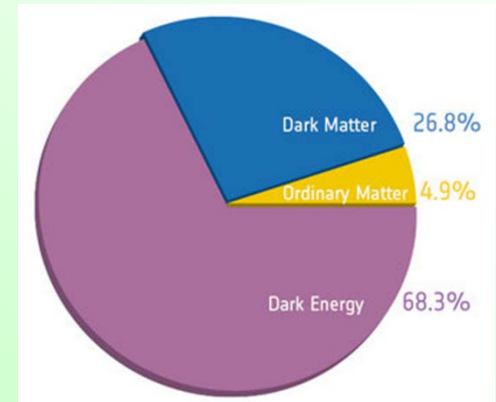
Superhorizon models  
(adiabatic perturbations)

Unfortunately, **primordial tensor perturbations** have not yet been observed.

# The presence of dark energy

The Universe is now **accelerating** !!

- 
- **Dark Energy is introduced**
  - or
  - **GR may be modified in the IR limit**



PLANCK

We have almost confirmed the presence of inflation and dark energy, but, unfortunately, we know neither **the identification of an inflaton** nor **that of dark energy**.

**Next task is to identify the inflaton and the origin of the dark energy.**

# Identification methods

- **Top down approach :**

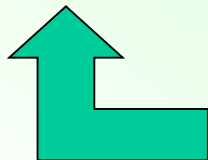
To **construct** the unique model from the **ultimate** theory like string theory.

(Recently, it may not be so actively studied.)

- **Bottom up approach**

To consider **the most general model**.

Then, we can **constrain models (or to single out the true model finally)** from the observational results.



**In this talk, we take the latter approach**

# Bottom up approach

- **Effective field theory approach :** (Weinberg 2008, Cheung et al. 2008)

The low-energy effective theory (after integrating out heavy mode with its mass  $M$ ).

**A ghost seems to appear** around the cut-off scale  $M$  ( $\gg E$ ).

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}m^2\phi^2 - \frac{1}{2}\frac{1}{M^2}(\square\phi)^2$$

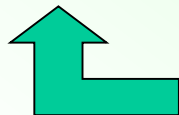
$E^2\phi^2$   $\frac{E^4}{M^2}\phi^2$

( $E$ : the energy scale we pay attention to)

$$M \gg E \quad \longleftrightarrow \quad \frac{1}{2}\partial_\mu\phi\partial^\mu\phi \gg \frac{1}{2}\frac{1}{M^2}(\square\phi)^2$$

- **Most general theory without ghost**

(if we are interested in the case in which **higher derivative terms play an important role in the dynamics.**)



In this talk, we take the latter approach



**The following question arises:**

**What is the most general  
scalar-tensor theory without ghost ?**

# How widely can we extend scalar tensor theory ?

- A kinetic term of an inflaton is not necessarily canonical.

$$\mathcal{L} = X - V(\phi), \quad X = -\frac{1}{2}g^{\mu\nu}\partial_\mu\phi\partial_\nu\phi \quad \longrightarrow \quad \mathcal{L} = K(\phi, X)$$

**(k-inflation)**  
(Armendariz-Picon et.al. 1999)

- An inflaton is not necessarily minimally coupled to gravity.

$$S = \int d^4x \sqrt{-g} \left( \frac{1}{2}M_G^2 R + \mathcal{L}_\phi \right) \quad \longrightarrow \quad \Delta S = \int d^4x \sqrt{-g} f(\phi) R$$

**(Higgs inflation)**

(Cervantes-Cota & Dehnen 1995, Bezrukov & M. Shaposhnikov 2008)

- Action may include higher derivatives.

(Nicolis et.al. 2009)

$$\mathcal{L} = K(\phi, X) \quad \longrightarrow \quad \Delta\mathcal{L} = G(\phi, X)\square\phi$$

**Theories with higher order derivatives  
are quite dangerous in general.**

# Example with higher order (time) derivatives

●  $L = \frac{1}{2}\ddot{q}^2(t)$   $\longrightarrow$   $q^{(4)} = 0$  requires **4** initial conditions.  
**EL eq.**



**2 (real) DOF**

●  $L_{\text{eq}}^{(1)} = \dot{q}u - \frac{1}{2}u^2$   $\longrightarrow$   $\begin{cases} \ddot{u} = 0, \\ \ddot{q} = u, \end{cases}$   $\longrightarrow$   $q^{(4)} = 0$   
**EL eq.**

$x \equiv \frac{q-u}{\sqrt{2}}, y \equiv \frac{q+u}{\sqrt{2}}$   $\longrightarrow$   $L_{\text{eq}}^{(1)} = -\dot{q}\dot{u} - \frac{1}{2}u^2 = \frac{1}{2}\dot{x}^2 - \frac{1}{2}\dot{y}^2 - \frac{1}{4}(x-y)^2.$

$\longrightarrow$   $H = \frac{1}{2}p_x^2 - \frac{1}{2}p_y^2 + \frac{1}{4}(x-y)^2.$   
 ( $p_x \equiv \dot{x}, p_y \equiv \dot{y}$ )

**2 (real) DOF = 1 healthy & 1 ghost**

●  $L_{\text{eq}}^{(2)} = \frac{1}{2}\dot{Q}^2 + \lambda(Q - \dot{q})$   $\longrightarrow$   $p \equiv \frac{\partial L_{\text{eq}}^{(2)}}{\partial \dot{q}} = -\lambda, P \equiv \frac{\partial L_{\text{eq}}^{(2)}}{\partial \dot{Q}} = \dot{Q}.$

$\longrightarrow$   $H = p\dot{q} + P\dot{Q} - L_{\text{eq}}^{(2)} = \frac{1}{2}P^2 + pQ.$

**Hamiltonian is unbounded through a linear momentum !!**

# Ostrogradski's theorem

(Ostrogradsky 1850)

Assume that  $L = L(\ddot{q}, \dot{q}, q)$  and  $\frac{\partial L}{\partial \ddot{q}}$  depends on  $\ddot{q}$  :

(Non-degeneracy)

→ 
$$\frac{\partial L}{\partial q} - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) + \frac{d^2}{dt^2} \left( \frac{\partial L}{\partial \ddot{q}} \right) = 0, \implies q^{(4)} = q^{(4)}(q^{(3)}, \ddot{q}, \dot{q}, q).$$

Canonical variables :

$$\begin{cases} q, & p := \frac{\partial L}{\partial \dot{q}} - \frac{d}{dt} \frac{\partial L}{\partial \ddot{q}} \left( = \frac{\partial L_{\text{eq}}}{\partial \dot{q}} \right), \\ Q := \dot{q}, & P := \frac{\partial L}{\partial \ddot{q}} \left( = \frac{\partial L_{\text{eq}}}{\partial \dot{Q}} \right). \end{cases}$$

$$L_{\text{eq}} = L(\dot{Q}, \dot{q}, q) + \lambda(Q - \dot{q})$$

Non-degeneracy  $\Leftrightarrow \ddot{q} = \ddot{q}(q, \dot{q}, \frac{\partial L}{\partial \ddot{q}}) \Leftrightarrow \dot{Q} = \ddot{q} = \ddot{q}(q, Q, P)$

Hamiltonian:  $H(q, Q, p, P) := p\dot{q} + P\dot{Q} - L$   
 $= pQ + P\ddot{q}(q, Q, P) - L(q, Q, \ddot{q}(q, Q, P)).$

→ **p depends linearly on H so that no system of this form can be stable !!**

**N.B.** 
$$\frac{\partial L}{\partial \phi} - \partial_\mu \left( \frac{\partial L}{\partial (\partial_\mu \phi)} \right) + \partial_\mu \partial_\nu \left( \frac{\partial L}{\partial (\partial_\mu \partial_\nu \phi)} \right) = 0. \implies \frac{i}{(p^2 + m_1^2)(p^2 + m_2^2)} = \frac{1}{m_2^2 - m_1^2} \left( \frac{i}{p^2 + m_1^2} - \frac{i}{p^2 + m_2^2} \right).$$
  
 (propagators)

**How to circumvent Ostrogradsky's arguments to obtain healthy higher order derivative theories ?**

# Loophole of Ostrogradski's theorem

We can **break the non-degeneracy condition** which requires that  $\frac{\partial L}{\partial \ddot{q}}$  depends on  $\ddot{q}$ .

(NB: another interesting possibility is infinite derivative theory)

In case Lagrangian depends on only **a position  $q$  and its velocity  $\dot{q}$** , **degeneracy** implies that **EOM is first order**, which represents not the dynamics but **the constraint**.



In case Lagrangian depends on  **$q$ ,  $\dot{q}$ ,  $\ddot{q}$** , degeneracy implies that **EOM can be (more than) second order**, which can represent the **dynamics**.

# Generalized Galileon = Horndeski

Deffayet et al. 2009, 2011

equivalence

Horndeski 1974

Kobayashi, MY, Yokoyama 2011

$$\mathcal{L}_2 = K(\phi, X)$$

$$\mathcal{L}_3 = -G_3(\phi, X) \square \phi,$$

$$\mathcal{L}_4 = G_4(\phi, X) R + G_{4X} \left[ (\square \phi)^2 - (\nabla_\mu \nabla_\nu \phi)^2 \right],$$

$$\mathcal{L}_5 = G_5(\phi, X) G_{\mu\nu} \nabla^\mu \nabla^\nu \phi$$

$$-\frac{1}{6} G_{5X} \left[ (\square \phi)^3 - 3 (\square \phi) (\nabla_\mu \nabla_\nu \phi)^2 + 2 (\nabla_\mu \nabla_\nu \phi)^3 \right].$$

$$X = -\frac{1}{2} (\nabla \phi)^2, \quad G_{iX} \equiv \partial G_i / \partial X.$$

This is **the most general scalar tensor theory whose Euler-Lagrange EOMs are up to second order** though the action includes second derivatives.

Many of inflation and dark energy models can be understood in a unified manner.

- NB :**
- $G_4 = M_G^2 / 2$  yields the Einstein-Hilbert action
  - $G_4 = f(\phi)$  yields a non-minimal coupling of the form  $f(\phi)R$
  - The new Higgs inflation with  $G^{\mu\nu} \partial_\mu \phi \partial_\nu \phi$  comes from  $G_5 \propto \phi$  after integration by parts.



# Horndeski theory

Horndeski 1974

In 1974, Horndeski presented the most general action (in four dimensions) constructed from the metric  $g$ , the scalar field  $\phi$ , and their derivatives,  $\partial g_{\mu\nu}, \partial^2 g_{\mu\nu}, \partial^3 g_{\mu\nu}, \dots, \partial\phi, \partial^2\phi, \partial^3\phi, \dots$  still having second-order equations.

$$\mathcal{L}_H = \delta_{\mu\nu\sigma}^{\alpha\beta\gamma} \left[ \kappa_1 \nabla^\mu \nabla_\alpha \phi R_{\beta\gamma}{}^{\nu\sigma} + \frac{2}{3} \kappa_{1X} \nabla^\mu \nabla_\alpha \phi \nabla^\nu \nabla_\beta \phi \nabla^\sigma \nabla_\gamma \phi + \kappa_3 \nabla_\alpha \phi \nabla^\mu \phi R_{\beta\gamma}{}^{\nu\sigma} + 2\kappa_{3X} \nabla_\alpha \phi \nabla^\mu \phi \nabla^\nu \nabla_\beta \phi \nabla^\sigma \nabla_\gamma \phi \right] + \delta_{\mu\nu}^{\alpha\beta} \left[ (F + 2W) R_{\alpha\beta}{}^{\mu\nu} + 2F_X \nabla^\mu \nabla_\alpha \phi \nabla^\nu \nabla_\beta \phi + 2\kappa_8 \nabla_\alpha \phi \nabla^\mu \phi \nabla^\nu \nabla_\beta \phi \right] - 6 (F_\phi + 2W_\phi - X\kappa_8) \square\phi + \kappa_9.$$

$$\left\{ \begin{array}{l} \kappa_1, \kappa_3, \kappa_8, \kappa_9, F : \text{functions of } \phi \text{ \& } X \text{ with} \\ W = W(\phi) \\ \delta_{\mu_1\mu_2\dots\mu_n}^{\alpha_1\alpha_2\dots\alpha_n} = n! \delta_{\mu_1}^{[\alpha_1} \delta_{\mu_2}^{\alpha_2} \dots \delta_{\mu_n}^{\alpha_n]} \end{array} \right. \quad F_X = 2(\kappa_3 + 2X\kappa_{3X} - \kappa_1\phi).$$

**What is the relation between Generalized Galileon and Horndeski's models ?**

$\Rightarrow$  **Both models are completely equivalent :** Kobayashi, MY, Yokoyama 2011

$$\left\{ \begin{array}{l} K = \kappa_9 + 4X \int^X dX' (\kappa_8\phi - 2\kappa_3\phi\phi), \\ G_3 = 6F_\phi - 2X\kappa_8 - 8X\kappa_3\phi + 2 \int^X dX' (\kappa_8 - 2\kappa_3\phi), \\ G_4 = 2F - 4X\kappa_3, \\ G_5 = -4\kappa_1, \end{array} \right. \left\{ \begin{array}{l} \mathcal{L}_2 = K(\phi, X), \\ \mathcal{L}_3 = -G_3(\phi, X) \square\phi, \\ \mathcal{L}_4 = G_4(\phi, X) R + G_{4X} [(\square\phi)^2 - (\nabla_\mu \nabla_\nu \phi)^2], \\ \mathcal{L}_5 = G_5(\phi, X) G_{\mu\nu} \nabla^\mu \nabla^\nu \phi \\ \quad - \frac{1}{6} G_{5X} [(\square\phi)^3 - 3(\square\phi) (\nabla_\mu \nabla_\nu \phi)^2 + 2(\nabla_\mu \nabla_\nu \phi)^3]. \end{array} \right.$$

# Beyond Horndeski theory

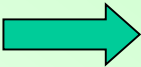
- Gleyzes, Langlois, Piazza, and Vernizzi (GLPV) pointed out that there is extended theory with the number of propagating degrees of freedom unchanged, even though apparent EOMs of the theory are higher (third) order.

(Gleyzes et al. 2014, Zumalacarregui & Garcia-Bellido 2014)

- Noui and Langlois pointed out the importance of degeneracy of kinetic matrix of terms with different order derivatives and proposed degenerate higher order scalar-tensor (DHOST) theory.

(Noui & Langlois 2016, etc)

- Another direction is to consider infinitely many higher derivatives (nonlocal theory). (Barnaby & Kamran 2008, References therein, also Leonardo's talk, Yun Soo's talk etc)

e.g.  $e^{-\square/\Lambda^2} (\square - m^2)\phi = 0.$    $\frac{i}{e^{p^2/\Lambda^2}(p^2 + m^2)}$  Only a pole of propagator appears with  $p^2 = m^2$ .

(Biswas, A. Mazumdar, W. Siegel 2006)

We are interested in an issue whether further extension is possible.

**As far as I know, all of the (finite) higher order derivative theories without ghosts include up to second time derivatives.**

**(Xian Gao proposed arbitrary higher spatial derivative theory in 2014.)**

**What happens if we consider  
“third” order time derivative theories ?  
(Is it straightforward extension of second  
order time derivative theories ???)**

**Healthy degenerate theories with  
“second” order derivatives for point particles**

**(as a first step with keeping in mind  
future extension to scalar-tensor theory)**

# Lagrangian up to second order derivatives

(Motohashi, Noui, Suyama, MY, Langlois. 2016)

$$L(\ddot{\phi}^a, \dot{\phi}^a, \phi^a; \dot{q}^i, q^i) \quad (a = 1, \dots, n; \quad i = 1, \dots, m)$$

{ Variables with **second** order derivatives:  $(\phi^a)$   
Variables with **first** order derivatives:  $(q^i)$

$\phi^a$  and  $q^i$  generically obey **fourth** order and **second** order equations of motion, respectively  $\rightarrow$  **Ostrogradsky instability**

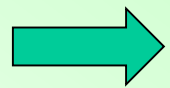
Let us derive the conditions to escape such an instability by Hamiltonian analysis, starting from the **equivalent Lagrangian**:

$$L_{eq}^{(1)}(\dot{Q}^a, Q^a; \dot{\phi}^a, \phi^a; \dot{q}^i, q^i, \lambda_a) \equiv L(\overset{\uparrow}{\dot{Q}^a}, \overset{\uparrow}{Q^a}, \phi^a; \dot{q}^i, q^i) + \lambda_a(\dot{\phi}^a - Q^a).$$

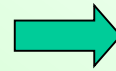
# Hamiltonian analysis

$$L_{eq}^{(1)}(\dot{Q}^a, Q^a; \dot{\phi}^a, \phi^a; \dot{q}^i, q^i, \lambda_a) \equiv L(\dot{Q}^a, Q^a, \phi^a; \dot{q}^i, q^i) + \lambda_a(\dot{\phi}^a - Q^a).$$

(3n+m) canonical variables :  $Q^a, \phi^a, q^i, \lambda^a$

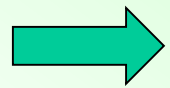


(3n+m) canonical momenta :  $P_a, \pi_a, p_i, \rho_a$



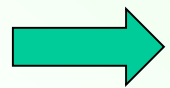
(4n+2m)  
phase DOF  
n ghosts!!

$$P_a = \frac{\partial L}{\partial \dot{Q}^a} \equiv L_{\dot{Q}^a}, \quad \pi_a = \frac{\partial L_{eq}^{(1)}}{\partial \dot{\phi}^a} = \lambda_a, \quad p_i = \frac{\partial L}{\partial \dot{q}^i} \equiv L_{\dot{q}^i}, \quad \rho_a = \frac{\partial L_{eq}^{(1)}}{\partial \dot{\lambda}^a} = 0.$$



Two sets of n primary constraints:

$$\Phi_a = \pi_a - \lambda_a \approx 0, \quad \Psi_a = \rho_a \approx 0,$$



$$H = H_0 + \pi_a Q^a \quad \text{with} \quad H_0 = P_a \dot{Q}^a + p_i \dot{q}^i - L(\dot{Q}^a, Q^a, \phi^a; \dot{q}^i, q^i)$$


$\pi^a$  appears linearly in Hamiltonian and hence it is unbounded if the system is nondegenerate without further primary constraints.

# Degenerate Lagrangian

For **healthy** theories, we have to eliminate **n** DOF from constraints

$$\begin{pmatrix} \delta P_a \\ \delta p_i \end{pmatrix} = K \begin{pmatrix} \delta \dot{Q}^b \\ \delta \dot{q}^j \end{pmatrix}, \quad K \equiv \begin{pmatrix} L_{\dot{Q}^a \dot{Q}^b} & L_{\dot{Q}^a \dot{q}^j} \\ L_{\dot{q}^i \dot{Q}^b} & L_{\dot{q}^i \dot{q}^j} \end{pmatrix} = \begin{pmatrix} L_{ab} & L_{aj} \\ L_{ib} & L_{ij} \end{pmatrix}$$


Kinetic matrix must be **degenerate** !!

  $L_{ab} - L_{ai} L^{ij} L_{jb} = 0$  (First DC)

Assume  $\det L_{ij} \neq 0$  ( $q^i$  is a normal variable)

$$\therefore \left( \det K = \det(L_{ab} - L_{ai} L^{ij} L_{jb}) \det L_{ij} \right)$$

↑
↑  
 (all **n** eigenvalues are zero) (m **non-zero** eigenvalues)

  $\Xi_a \equiv P_a - F_a(p_i, Q^b, \phi^b, q^i) \approx 0.$

Assume  $\det L_{ij} \neq 0$

(Additional primary constraints)

# Additional primary constraints

(Motohashi & Suyama 2014)

$$\Xi_a \equiv P_a - F_a(p_i, Q^b, \phi^b, q^i) \approx 0.$$

$$\Rightarrow H_T = H(P_a, p_i, \pi_a, Q^a, q^i, \phi^a) + \mu^a \Phi_a + \nu^a \Psi_a + \xi^a \Xi_a$$

( $\Phi_a = \pi_a - \lambda_a \approx 0$ ,  $\Psi_a = \rho_a \approx 0$ )

$$\Rightarrow \begin{cases} \dot{\Phi}_a = \{\Phi_a, H\} - \nu^a + \xi^b \partial F_b / \partial \phi^a \approx 0, \\ \dot{\Psi}_a = \mu^a \approx 0, & (\{\Phi_a, \Xi_b\} = \partial F_b / \partial \phi^a) \Rightarrow \text{Fix } \mu^a \text{ \& } \nu^a \\ \dot{\Xi}_a = \{\Xi_a, H\} + \xi^b \underbrace{\{\Xi_a, \Xi_b\}}_{\text{Mab}} \approx 0. \end{cases}$$

If  $\det \text{Mab} \neq 0$ , all  $\xi^a$  are fixed and **no secondary constraints**.

$\Rightarrow$  **Not sufficient number of constraints** to eliminate all ghosts.

$\Rightarrow$  Simplest case for **healthy theory** : **Mab = 0 (second DC)**

$$\Theta_a \equiv \{\Xi_a, H\} = -\pi_a + \{\Xi_a, H_0\} - \partial F_a / \partial \phi_b Q^b \approx 0$$

(**New constraint**, which **fixes all  $\pi_a$**  in terms of the other phase space variables)



# Summary of second order system

$$L(\ddot{\phi}^a, \dot{\phi}^a, \phi^a; \dot{q}^i, q^i)$$

$$\longleftrightarrow L_{eq}^{(1)}(\dot{Q}^a, Q^a; \dot{\phi}^a, \phi^a; \dot{q}^i, q^i, \lambda_a) \equiv L(\dot{Q}^a, Q^a, \phi^a; \dot{q}^i, q^i) + \lambda_a(\dot{\phi}^a - Q^a).$$

(3n+m) canonical variables :  $\overset{\phi^a}{\mathbf{Q}^a}, \phi^a, \mathbf{q}^i, \lambda^a$   
 (3n+m) canonical momenta :  ~~$\mathbf{P}^a, \pi_a, \mathbf{p}_i, \rho_a$~~

(2n+2m) phase DOF  
 no ghosts!!

● **First DC** :  $L_{ab} - L_{ai}L^{ij}L_{jb} = 0$ 

$$\left[ \begin{pmatrix} \delta P_a \\ \delta p_i \end{pmatrix} = K \begin{pmatrix} \delta Q^b \\ \delta q^j \end{pmatrix}, \quad K \equiv \begin{pmatrix} L_{\dot{Q}^a \dot{Q}^b} & L_{\dot{Q}^a \dot{q}^j} \\ L_{\dot{q}^i \dot{Q}^b} & L_{\dot{q}^i \dot{q}^j} \end{pmatrix} = \begin{pmatrix} L_{ab} & L_{aj} \\ L_{ib} & L_{ij} \end{pmatrix} \right]$$

$$\longrightarrow \Xi_a \equiv P_a - F_a(p_i, Q^b, \phi^b, q^i) \approx 0. \quad (\text{primary constraints})$$
 Assume  $\det L_{ij} \neq 0$

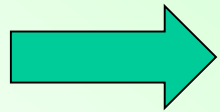
● **Second DC** :  $M_{ab} = \{\Xi_a, \Xi_b\} = -\frac{\partial F_a}{\partial Q^b} + \frac{\partial F_b}{\partial Q^a} + \frac{\partial F_a}{\partial q_i} \frac{\partial F_b}{\partial p_i} - \frac{\partial F_a}{\partial p_i} \frac{\partial F_b}{\partial q_i} = 0.$

$$\longrightarrow \Theta_a \equiv \{\Xi_a, H\} = -\pi_a + \{\Xi_a, H_0\} - \partial F_a / \partial \phi_b Q^b \approx 0 \quad (\text{secondary CS})$$

$$H = H_0 + \pi_a Q^a \quad \text{with} \quad H_0 = P_a \dot{Q}^a + p_a \dot{q}^a - L(\dot{Q}^a, Q^a, \phi^a; \dot{q}^i, q^i)$$

**EOMs** :  $(L_{ab} - L_{ai}L^{ij}L_{jb})\phi^{b(4)} + M_{ab}\phi^{b(3)} + \dots = 0 \longrightarrow \text{second order}$

**Is the extension to third order (time)  
derivative system straightforward ???  
No essential difference ???**



**There is big difference !!**

# Quadratic model with third order derivatives

(Motohashi, Suyama, MY 2018)

$$\begin{aligned}
 L = & \frac{a_{nm}}{2} \ddot{\psi}^n \ddot{\psi}^m + \frac{b_{nm}}{2} \ddot{\psi}^n \ddot{\psi}^m + \frac{c_{nm}}{2} \dot{\psi}^n \dot{\psi}^m \\
 & + \frac{d_{nm}}{2} \psi^n \psi^m + e_{nm} \ddot{\psi}^n \ddot{\psi}^m + f_{nm} \dot{\psi}^n \dot{\psi}^m \\
 & + \frac{A_{ij}}{2} \dot{q}^i \dot{q}^j + \frac{B_{ij}}{2} q^i q^j + C_{ij} \dot{q}^i q^j + \alpha_{ni} \ddot{\psi}^n \dot{q}^i.
 \end{aligned}$$

$(n = 1, \dots, N)$   
 $(i = 1, \dots, I)$

$$(\psi^n = \psi^n(t), \quad q^i = q^i(t), \quad a_{nm}, b_{nm}, \dots : \text{const}, \quad \det A_{ij} \neq 0)$$

$$\longleftrightarrow L_{\text{eq}} = L(\overset{\uparrow}{\dot{\psi}^n}, \overset{\uparrow}{\psi^n}, \overset{\uparrow}{\psi^n}, \dot{Q}^n, Q^n, R^n, \psi^n, \dot{q}^i, q^i) + \xi_n(\psi^n - R^n) + \lambda_n(\dot{R}^n - Q^n).$$

**Canonical momenta for  $(Q^n, R^n, \psi^n, q^i, \lambda_n, \xi_n)$  :**

$$\begin{aligned}
 P_{Q^n} &= a_{nm} \dot{Q}^m + \alpha_{ni} \dot{q}^i + e_{nm} Q^m, & P_{R^n} &= \lambda_n, \quad \pi_{\psi^n} = \xi_n, \\
 p_i &= \alpha_{ni} \dot{Q}^n + A_{ij} \dot{q}^j + C_{ij} q^j, & \rho_{\lambda_n} &= 0, \quad \rho_{\xi_n} = 0.
 \end{aligned}$$

**Primary constraints**

# Linear dependence of momenta on Hamiltonian

$$\left\{ \begin{array}{l} \text{Canonical variables : } (Q^n, R^n, \psi^n, q^i, \lambda_n, \xi_n) \\ \text{Canonical momenta : } (P_{Q^n}, P_{R^n}, \pi_{\psi^n}, p^i, \rho_{\lambda_n}, \rho_{\xi_n}) \end{array} \right. \quad \begin{array}{l} \text{primary constraints} \\ \swarrow \end{array}$$

$\uparrow \dot{\psi}^n \quad \uparrow \dot{\psi}^n$

→  $H = H_0 + P_{R^n} Q^n + \pi_{\psi^n} R^n.$

$$\left( H_0 = \frac{1}{2} A^{ij} \tilde{p}_i \tilde{p}_j - \frac{1}{2} b_{nm} Q^n Q^m - \frac{1}{2} c_{nm} R^n R^m - \frac{1}{2} d_{nm} \psi^n \psi^m - f_{nm} Q^n R^m - \frac{1}{2} B_{ij} q^i q^j : \text{no dependence on } P_{R^n}, \pi_{\psi^n} \right)$$

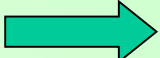
Due to the **linear** dependence of **momenta**,  
Hamiltonian is **unbounded !!**

According to Ostrogradsky arguments and the lesson from second order derivative system, we expect that we have only to **remove this linear momentum dependence** for a healthy theory.


# Conditions for a healthy theory

$$H = H_0 + P_{R^n} Q^n + \pi_{\psi^n} R^n.$$


- **First DC :**  $a_{nm} - \alpha_{ni} A^{ij} \alpha_{mj} = 0.$  (kinetic matrix is degenerate)

Assume  $\det A_{ij} \neq 0$    $\Psi_n \equiv P_{Q^n} - e_{nm} Q^m - A^{ij} \alpha_{nj} \tilde{p}_i \approx 0.$  (primary constraint)

- **Second DC :**  $\{\Psi_n, \Psi_m\} = -2[e_{nm} + \alpha_{ni} (A^{-1} C A^{-1})^{ij} \alpha_{mj}] = 0.$

  $\Upsilon_n \equiv -\{\Psi_n, H\} = P_{R^n} - b_{nm} Q^m + \dots \approx 0,$  (secondary CS)

- **Third DC :**  $\{\Upsilon_n, \Psi_m\} = -b_{nm} - \alpha_{ni} [(4\bar{C}^2 + \bar{B}) A^{-1}]^{ij} \alpha_{mj} = 0.$

  $\Lambda_n \equiv -\{\Upsilon_n, H\} = \pi_{\psi^n} + 2f_{nm} Q^m + \dots \approx 0,$  (tertiary CS)

Now, the linear dependences are completely eliminated !!

No ghost (no instability) ???

# The ghosts still lurk in the Hamiltonian

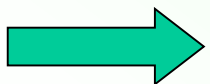
After erasing  $P_{R^n}, b_{nm}, \pi_{\psi^n}$  using DCs and constraints, the Hamiltonian  $H$  reduces to

$$H = \frac{1}{2}A^{ij}\bar{p}_i\bar{p}_j - \frac{1}{2}B_{ij}\bar{q}^i\bar{q}^j + \frac{1}{2}c_{nm}R^n R^m - \frac{1}{2}d_{nm}\psi^n\psi^m \\ + \alpha_{ni}[(4\bar{C}^2 + \bar{B})A^{-1}]^{ij}\bar{p}_j R^n - 2\alpha_{ni}(\bar{C}\bar{B})^i_j\bar{q}^j R^n \\ - 2[f_{nm} + 4\alpha_{ni}(\bar{C}^3 A^{-1})^{ij}\alpha_{mj}]Q^m R^n,$$

$$(\bar{p}_i \equiv p_i - C_{ik}q^k + 2\alpha_{nk}A^{kl}C_{li}Q^n, \quad \bar{q}^i \equiv q^i + \alpha_{nk}A^{ki}Q^n)$$

$Q^n (= \dot{\psi}^n)$  appears only **linearly** in  $H$ ,  
making the **Hamiltonian  $H$  unbounded !!**

Eliminating the **linear momentum terms** is always **necessary** to kill the Ostrogradsky ghosts, but is **not sufficient** for **higher-than-second-order derivative system**.



**Needs to fix  $Q^n$  as well.**

# Healthy theory

**Canonical variables :**  $(\cancel{Q}^n, R^n, \psi^n, q^i, \cancel{\lambda}_n, \cancel{\xi}_n)$  primary constraints  
 $\uparrow \dot{\psi}^n$     $\uparrow \dot{\psi}^n$

**Canonical momenta :**  $(\cancel{P}_Q^n, \cancel{P}_R^n, \cancel{\pi}_\psi^n, p^i, \cancel{\rho}_\lambda^n, \cancel{\rho}_\xi^n)$

$$H = H_0 + P_{R^n} Q^n + \pi_{\psi^n} R^n. \quad \text{secondary \& tertiary CS}$$

● **Fourth DC :**  $\{\Lambda_n, \Psi_m\} = 2(f_{nm} - \alpha_{ni} M^{ij} \alpha_{mj}) = 0.$

➡  $\Omega_n \equiv -\{\Lambda_n, H\} = c_{nm} Q^m - d_{nm} \psi^m + \dots \approx 0, \quad \text{(quaternary CS)}$

**$Q^n$  is now fixed (and expressed in terms of other variables).**

➡ **Only healthy DOFs  $(\psi^n, R^n (= \dot{\psi}^n), q^i, p^i)$  remain.**

**Hamiltonian is bounded and no ghosts.**

➡ **EOMs can be reduced to the **second order** system**

# Dirac matrix

$$D = \begin{array}{c|cccccc} & \overbrace{\Phi_\beta \quad \bar{\Phi}_\beta}^{\text{C1}} & \Psi_m & \Omega_m & \Upsilon_m & \Lambda_m \\ \hline \Phi_\alpha & 0 & -\mathbf{1} & * & * & * & * \\ \bar{\Phi}_\alpha & \mathbf{1} & 0 & 0 & 0 & 0 & 0 \\ \Psi_n & * & 0 & 0 & -Z_{mn} & 0 & 0 \\ \Omega_n & * & 0 & Z_{nm} & * & * & * \\ \Upsilon_n & * & 0 & 0 & * & 0 & Z_{mn} \\ \Lambda_n & * & 0 & 0 & * & -Z_{nm} & * \end{array}$$

Condition to complete Dirac procedure :

$$\det Z_{nm} = \det\{\Omega_n, \Psi_m\} \neq 0$$

➡  $\det D \neq 0$  ➡ All constraints are **second class**

➡ **Healthy (N+I) DOFs.**



# Reduction

(Motohashi, Suyama, MY 2018)

$$L(q^i, \dot{q}^i, \ddot{q}^i, \dots) \quad \longrightarrow \quad \text{Healthy theory}$$

A chain of constraints :  $\Phi_A \approx 0$  ( $A = 1, 2, \dots$ )

If all of them are **second class**, according to Maskawa & Nakajima (1976), a **canonical transformation** exists such that new variables are classified to

$$(Q^a, P_a), \quad (\tilde{Q}^m, \tilde{P}_m) \quad \text{with} \quad \tilde{Q}^m \approx 0 \quad \& \quad \tilde{P}_m \approx 0$$

$\swarrow$   
**unconstrained** and governed by  $H(Q^a, P_a)$ , that is,  $\dot{Q}^a = \frac{\partial H}{\partial P_a}$ ,  $\dot{P}_a = -\frac{\partial H}{\partial Q^a}$

$$\downarrow$$
$$P_a = P_a(Q^b, \dot{Q}^b)$$

$$\longrightarrow \quad L'(Q^a, \dot{Q}^a) = P_a \dot{Q}^a - H(Q^a, P_a)$$

**This Lagrangian is nondegenerate and contains up to first-order time derivatives of  $Q_a$ .**

(If a theory has first-class constraints, only gauge fixing is necessary.)

# Summary

- We have investigated how to obtain **healthy degenerate theory with higher-than-second derivative**.
- Eliminating **linear momentum** terms in the Hamiltonian is **necessary and sufficient** to kill the ghosts for **second** order derivative system.
- On the other hand, this is necessary but **not sufficient** for the Lagrangian with **higher than second-order** derivatives. The **(lurked) linear** dependence of **canonical variables corresponding to second (even higher) derivatives** must be removed as well.