Kerr-Schild DFT and Classical Double Copy

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Strings, Branes and Gauge Theories

 Double copy structure states that the scattering amplitudes of the Yang-Mills theory and gravity are related by exchanging the color and kinematic factors [Bern, Carrasco Johansson 2008,2010]

 $c_i \leftrightarrows n_i$

color factor c_i : a polynomial of structure constants f^{abc} , kinematic factor n_i : a polynomial of Lorentz-invariant contractions of polarization vectors ϵ_i and momenta p_i .

 Gravity amplitudes can be obtained by just replacing the color factor to the kinematic factor without any knowledge of the gravity action or Feynman rules. Spectrum

$$\left.\begin{array}{l} \operatorname{\mathsf{graviton}}^{\pm 2}\left(p_{i}\right) = \operatorname{\mathsf{gluon}}^{\pm 1}\left(p_{i}\right) \otimes \operatorname{\mathsf{gluon}}^{\pm 1}\left(p_{i}\right) \\ \\ \operatorname{\mathsf{dilaton}} \\ \operatorname{\mathsf{axion}} \end{array}\right\} = \operatorname{\mathsf{gluon}}^{\pm 1}\left(p_{i}\right) \otimes \operatorname{\mathsf{gluon}}^{\pm 1}\left(p_{i}\right) \\ \end{array}\right.$$

• The double copy has the potential to provide a new way of quantum gravity

(perturbative) gravity = $(Yang-Mills)^2$

String theory origin

- For tree level amplitude, it is equivalent to the field theory limit, $\alpha' \rightarrow 0$, of the KLT relation in closed string theory.
- Tree level closed string and open string scattering amplitudes are related via the KLT relation [Kawai, Lewellen, Tye 1986]

$$M_n^{\text{tree}} = A_n^{\text{tree}} \mathcal{K}_n \tilde{A}_n^{\text{tree}}$$

where \mathcal{K}_n is the KLT kernel.

• KLT relation provides the string theory origin of double copy structure.



- Tree level scattering amplitude —> on-shell, no quantum effects.
 It is possible to deduce its extension to the level of the classical equations of motion.
- Q: Can solutions of the Einstein field equations be represented by solutions of the Yang-Mills equations beyond perturbative level?

Solution of GR
$$\stackrel{\longrightarrow}{\longleftrightarrow}$$
 Solution of YM

• Graviton $h_{\mu\nu}$ is given by the linearized perturbation of the metric

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

Recall the spectrum relation. Is it possible to represent $h_{\mu\nu} \sim A_{\mu} \tilde{A}_{\nu}$?

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- One way is the so called classical double copy based on Kerr-Schild formalism in GR [Monteiro, O'Connell, White, 2014]
- The Kerr-Schild ansatz is an extension of linear perturbation around a background metric *g*.
- Einstein equation is nonlinear PDE \implies Hard to solve
- What is the condition

Einstein equation becomes linear?

- Kerr and Schild proposed a metric ansatz which makes Einstein equation a linear equation [Kerr 1963], [Kerr, Schild 1965].
- Meyers-Perry BH, (A)dS Kerr, (A)dS Kerr-Newman, Black string, branes, Waves in flat and (A)dS spaces (PP-wave, Kundt wave, Shock wave) etc.

Kerr-Schild ansatz

$$g_{\mu\nu} = \tilde{g}_{\mu\nu} + \kappa \varphi \ell_{\mu} \ell_{\nu}$$

 $\tilde{g}_{\mu\nu}$: a background metric satisfying Einstein equation ℓ_{μ} : null vector

$$\ell_{\mu}\tilde{g}^{\mu\nu}\ell_{\nu} = \ell_{\mu}g^{\mu\nu}\ell_{\nu} = 0$$

• The main advantage of the Kerr-Schild ansatz is that it preserves some features of the linearized perturbation

$$g^{\mu\nu} = \tilde{g}^{\mu\nu} - \kappa \varphi \ell^{\mu} \ell^{\nu}, \qquad \det(g) = \det(\tilde{g})$$

 Suppose a vacuum Einstein equation, R_{μν} = 0. We get a consistency condition by contracting the null vectors ℓ^μ with R_{μν}

$$R_{\mu\nu}\ell^{\mu}\ell^{\nu} = -\kappa\varphi g^{\nu\sigma} \left(\ell^{\mu}\tilde{\nabla}_{\mu}\ell_{\nu}\right) \left(\ell^{\rho}\tilde{\nabla}_{\rho}\ell_{\sigma}\right) = 0\,,$$

where $\tilde{\nabla}_{\mu}$ is the covariant derivative with respect to the background metric \tilde{g} .

• Choosing affine parameter, ℓ^{μ} is null and geodesic

$$\ell^{\mu}\tilde{\nabla}_{\mu}\ell_{\nu}=0\,.$$

The vacuum Einstein equation reduces to

$$R_{\mu\nu} = \kappa R^{(1)}{}_{\mu\nu} + \kappa^2 \varphi \ell_{\mu} \ell^{\rho} R^{(1)}{}_{\rho\nu} = 0 \,,$$

where $R^{(1)}$ is the linear terms with respect to κ ,

$$R^{(1)}{}_{\mu
u} = \kappa \tilde{
abla}_{
ho} \left(\tilde{
abla}_{(\mu} (arphi \ell_{
u}) \ell^{
ho} \right) - rac{1}{2} \tilde{
abla}^{
ho} (arphi \ell_{\mu} \ell_{
u})
ight),$$

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and the Einstein equation is the same as $R^{(1)}_{\mu\nu} = 0$.

Examples

Schwarzschild BH in Eddington-Finkelstein coordinate

$$g_{\mu\nu} = \eta_{\mu\nu} + \frac{2GM}{r} \ell_{\mu} \ell_{\nu}$$

where

$$\ell^{\mu} = \left(1, \frac{x^i}{r}\right), \quad r^2 = x^i x_i, \quad i = 1 \dots 3$$

Kerr BH in KS coordinate

$$\begin{split} ds^2 &= -dt^2 + dx^2 + dy^2 + dz^2 \\ &+ \frac{2mr^3}{r^4 + a^2z^2} \Big[dt + \frac{z}{r} dz + \frac{r}{r^2 + a^2} (xdx + ydy) - \frac{a}{r^2 + a^2} (xdy - ydx) \Big]^2 \end{split}$$

and the null vector is given by

$$\varphi = \frac{2MGr^3}{r^4 + a^2z^2}, \qquad \ell_{\mu} = \left(1, \frac{rx + ay}{r^2 + a^2}, \frac{ry - ax}{r^2 + a^2}, \frac{z}{r}\right)$$

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Classical double copy in GR

• Consider KS ansatz on a flat background, $\tilde{g} = \eta$

$$g_{\mu\nu} = \eta_{\mu\nu} + \varphi \ell_{\mu} \ell_{\nu}$$

• Identify the null vector ℓ and φ with gauge field and the biadjoint scala field [Monteiro, O'Connell, White, 2014]

$$A_{\mu} = \varphi \ell_{\mu}$$

• Assume that spacetime is stationary (no time dependence) and choose ℓ^{μ} as $\ell^0 = 1$

$$R_{0i} = \frac{1}{2} \partial^{j} \left(\partial_{i} (\varphi \ell_{j}) - \partial_{j} (\varphi \ell_{i}) \right) = -\frac{1}{2} \partial^{j} F_{ij}$$

where $F_{ij} = \partial_i A_j - \partial_j A_i$

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• How can we include Kalb-Ramond field $B_{\mu\nu}$ and dilaton ϕ in Kerr-Schild formalism?



- Curved background generalization It is not clear how to define scattering amplitude in curved background in general. (time-dependent backgrounds, nonasymptotic flat spaces)
- Classical double copy for non Kerr-Schild type geometries?
- non-abelian structure?

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Generalized Kerr-Schild method in DFT

⇒ Linearization of the equations of motion for supergravities (string NSNS sector).

 \implies Arbitrary on-shell background \rightarrow More general spacetime

- Classical double copy for entire massless NSNS sector
- Classical double copy in Killing spinor equation

 \implies From the Killing spinor equation for gravitino, Yang-Mills BPS equation can be derived

Why DFT?

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- DFT is the best framework for describing the double copy structure.
- Double copy Left-right decomposition of closed string theory
- · Generalized metric is represented by the coset

$$\mathcal{H} \to \frac{O(d,d)}{O(d-1,1) \times O(1,1-d)}$$

and this implies there are two local Lorentz groups $\Longrightarrow \{e_{\mu}{}^{m}, \bar{e}_{\mu}{}^{\bar{m}}\}$

 These are related with local Lorentz groups for left-right sectors of closed string theory. [Arkani-Hamed,Kaplan, 2008], [Hohm, 2011]

$$\eta_{\mu\nu} + h_{\mu\nu} \to h_{m\bar{n}}$$

• Cheung and Remmen derived perturbative DFT action (without dilaton and $B_{\mu\nu}$) around an arbitrary curved background from Einstein-Hilbert action by assuming the two local Lorentz groups. [Cheung, Remmen, 2016]

Generalized Kerr-Schild ansatz

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- Double Field Theory: String low energy effective field theory which is manifest under O(d, d) T-duality. ⇒ Doubling the dimension
- Long History:
 - Sigma model side [Duff, 1990], [Tseytlin, 1990,1991], [Hull, 2004]
 - Low energy effective field theory side [Siegel, 1993], [Hull, Zwiebach, 2009]
- · Geometrization of all the form fields in supergravities.
- Recall the Maxwell equation:

$$\partial_{\mu}F^{\mu\nu} = j^{\nu}, \qquad dF = 0$$

Field contents

- Manifest under T-duality $\rightarrow O(d, d)$ tensors
- generalized metric \mathcal{H}_{MN} : rank 2 tensor wrt O(d, d), which is an O(d, d) element

$$\mathcal{H}_{MN}\mathcal{J}^{NP}\mathcal{H}_{PQ}=\mathcal{J}_{MQ}$$

where \mathcal{J}_{MN} is the O(d, d) metric parametrized

$$\mathcal{J}_{MN} = \begin{pmatrix} 0 & \delta^{\mu}{}_{\nu} \\ \delta_{\mu}{}^{\nu} & 0 \end{pmatrix}$$

• Parametrization in terms of supergravity fields $\{g, B, \phi\}$

$$\mathcal{H}_{MN} = \begin{pmatrix} g^{\mu\nu} & -g^{\mu\rho}B_{\rho\nu} \\ B_{\mu\rho}g^{\rho\nu} & g_{\mu\nu} - B_{\mu\rho}g^{\rho\sigma}B_{\sigma\nu} \end{pmatrix}$$

• DFT scalar d : scalar wrt O(d, d)

$$e^{-2d} = \sqrt{-g}e^{-2\phi}$$

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Linear perturbation

• First, we analyze the properties of linear perturbations of generalized metric around an on-shell background generalized metric \mathcal{H}_0 satisfying

$$\mathcal{H}_{0MN}\mathcal{J}^{NP}\mathcal{H}_{0PQ}=\mathcal{J}_{MQ}$$

• Split ${\mathcal H}$ into the background part and perturbation parts

$$\mathcal{H}_{MN} = \mathcal{H}_{0MN} + \kappa \hat{\gamma}_{MN} \,,$$

where $\hat{\gamma}$ describes perturbation and κ is a small expansion parameter.

• From O(d, d) constraint for \mathcal{H}_0 , $(\mathcal{H}_0)^2 = 1$, one can define a background chirality and the corresponding projection operators

$$P_0 = rac{1}{2}ig(\mathcal{J}+\mathcal{H}_0ig)\,, \qquad ar{P}_0 = rac{1}{2}ig(\mathcal{J}-\mathcal{H}_0ig)\,,$$

One can show that γ̂ has mixed chirality

$$\hat{\gamma} = P_0 \hat{\gamma} \bar{P}_0 + \bar{P}_0 \hat{\gamma} P_0 \,, \qquad P_0 \hat{\gamma} P_0 = \bar{P}_0 \hat{\gamma} \bar{P}_0 = 0$$

generalized Kerr-Schild ansatz

- Following the conventional Kerr-Schild ansatz, we now assume that γ̂ is a finite perturbation and κ is a formal finite parameter. The chirality condition is no longer a linearized approximation, but an exact relation.
- We introduce an ansatz for the generalized metric

$$\mathcal{H}_{MN} = \mathcal{H}_{0MN} + \kappa \varphi \left(K_M \bar{K}_N + \bar{K}_M K_N \right),$$

where K and \bar{K} are null vectors

$$K_M K^M = 0, \qquad \bar{K}_M \bar{K}^M = 0,$$

and satisfy the chirality conditions as

$$P_{0MN}K^N = K_M, \qquad \bar{P}_{0MN}\bar{K}^N = \bar{K}_M, \qquad K_M\bar{K}^M = 0,$$

• We refer this form as generalized Kerr-Schild ansatz. This ansatz satisfies the O(d, d) constraint automatically without any approximation or truncation.

• Chirality condition \implies the K_M and \bar{K}_M are parametrized in terms of the *d*-dimensional vectors l^{μ} and \bar{l}^{μ}

$$K_{M} = \frac{1}{\sqrt{2}} \begin{pmatrix} l^{\mu} \\ (\tilde{g} + \tilde{B})_{\mu\nu} l^{\nu} \end{pmatrix}, \qquad \bar{K}_{M} = \frac{1}{\sqrt{2}} \begin{pmatrix} \bar{l}^{\mu} \\ (-\tilde{g} + \tilde{B})_{\mu\nu} \bar{l}^{\nu} \end{pmatrix}.$$

• Null condition $\implies l$ and \overline{l} are null vectors

$$l^{\mu}\tilde{g}_{\mu\nu}l^{\nu} = l^{\mu}l_{\mu} = 0, \qquad \bar{l}^{\mu}\tilde{g}_{\mu\nu}\bar{l}^{\nu} = \bar{l}^{\mu}\bar{l}_{\mu} = 0, \qquad l\cdot\bar{l}\neq 0$$

- More than one pair of null vectors?
- It is strictly forbidden in the Lorentzian signature metric! (Theory of quadratic form)

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Using the parametrization of generalized metric, we have

$$(g^{-1})^{\mu\nu} = (\tilde{g}^{-1})^{\mu\nu} + \kappa\varphi l^{(\mu}\bar{l}^{\nu)} ,$$

$$g_{\mu\nu} = \tilde{g}_{\mu\nu} - \frac{\kappa\varphi}{1 + \frac{1}{2}\kappa\varphi(l\cdot\bar{l})} l_{(\mu}\bar{l}_{\nu)} ,$$

$$B_{\mu\nu} = \tilde{B}_{\mu\nu} + \frac{\kappa\varphi}{1 + \frac{1}{2}\kappa\varphi(l\cdot\bar{l})} l_{[\mu}\bar{l}_{\nu]} ,$$

$$\det g = (\det \tilde{g}) \left(1 + \frac{1}{2}\kappa\varphi(l\cdot\bar{l})\right)^{-2}$$

- Though \mathcal{H} is linear in κ , g and B are nonlinear.
- If we identify *l^μ* and *l
 [¯]* and ignore the *B* field, then it reduces to the conventional Kerr-Schild ansatz,

$$g^{\mu\nu} = \tilde{g}^{\mu\nu} + \kappa \varphi l^{\mu} l^{\nu}, \qquad g_{\mu\nu} = \tilde{g}_{\mu\nu} - \kappa \varphi l_{\mu} l_{\nu}.$$

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Field equations and linear structure

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• In GR, equations of motion is written in terms of curvature tensor

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 0$$

In Riemannian geometry, Riemann tensor is given by commutator of covariant derivative

$$[\nabla_{\mu}, \nabla_{\nu}]V_{\rho} = R_{\mu\nu\rho\sigma}V^{\sigma}$$

DFT covariant derivative and curvature?

Field equations in DFT

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- Generalized Lie derivative: Recast the diffeomorphism and one-form gauge transform of $B_{\mu\nu}$ in an O(D, D) covariant way.
- "Semi" covariant derivative with respect to the gen. diffeomorphism [Jeon,KL,Park, 2011]

$$\nabla_M V_N = \partial_M V_N + \Gamma_{MNP} V^P$$

Generalized curvature tensor and scalar

$$\mathcal{S}_{MN} = P_M{}^P \bar{P}_N{}^Q P^{RS} S_{RPSQ}, \qquad \mathcal{S} := 2P^{MN} P^{PQ} S_{MPNQ}$$

where

$$S_{MNPQ} = \frac{1}{2} \left(R_{MNPQ} + R_{PQMN} - \Gamma^{R}{}_{MN} \Gamma_{RPQ} \right)$$
$$R_{MNPQ} = \partial_{M} \Gamma_{NPQ} - \partial_{N} \Gamma_{MPQ} + \Gamma_{MP}{}^{R} \Gamma_{NRQ} - \Gamma_{NP}{}^{R} \Gamma_{MRQ}$$

DFT field equation in terms of supergravity fields

$$S_{\text{eff.}} = \int dx^D \sqrt{-g} e^{-2\phi} \left(R + 4\partial_\mu \phi \partial^\mu \phi - \frac{1}{12} H_{\lambda\mu\nu} H^{\lambda\mu\nu} \right)$$

EOM

$$\mathcal{G}_{\mu\nu} = R_{\mu\nu} + 2\nabla_{(\mu}\partial_{\nu)}\phi - \frac{1}{4}H_{\mu}^{\ \rho\sigma}H_{\nu\rho\sigma} = 0,$$

$$\mathcal{B}_{\mu\nu} = -\frac{1}{2}\nabla^{\rho}H_{\rho\mu\nu} + \partial^{\rho}\phi H_{\rho\mu\nu} = 0.$$

· For simplicity consider a flat background,

$$\mathcal{H}_{0MN} = egin{pmatrix} \eta^{\mu
u} & 0 \ 0 & \eta_{\mu
u} \end{pmatrix}, \qquad d_0 = ext{const.}$$

Generalized Kerr-Schild ansatz

$$\mathcal{H}_{MN} = \mathcal{H}_{0MN} + \kappa \varphi \big(K_M \bar{K}_N + \bar{K}_M K_N \big)$$
$$d = d_0 + \kappa f.$$

• An on-shell condition from the DFT equations of motion, $S_{KL} = 0$,

$$K^{K}\bar{K}^{L}S_{KL} = 2K^{K}\bar{K}^{L}\partial_{K}\partial_{L}f - \frac{1}{2}\varphi(K^{K}\partial_{K}\bar{K}_{M})(K^{L}\partial_{L}\bar{K}^{M}) + \frac{1}{2}\varphi(\bar{K}^{K}\partial_{K}K_{M})(\bar{K}^{L}\partial_{L}K^{M}) = 0.$$

Recall that in GR, $R_{\mu\nu}\ell^{\mu}\ell^{\nu} = -\kappa\varphi g^{\nu\sigma} \left(\ell^{\mu}\tilde{\nabla}_{\mu}\ell_{\nu}\right) \left(\ell^{\rho}\tilde{\nabla}_{\rho}\ell_{\sigma}\right) = 0$.

On-shell Condition

• We shall impose stronger conditions

$$\bar{K}^M \partial_M K_P = 0$$
, $K^M \partial_M \bar{K}_P = 0$, $K^P \partial_P f = 0$, $\bar{K}^P \partial_P f = 0$.

DFT connection satisfies

 $K^P \Gamma_{PMN} \bar{K}^N = 0 \,, \qquad \bar{K}^P \Gamma_{PMN} K^N = 0 \,, \qquad \Gamma^P{}_{PM} K^M = \Gamma^P{}_{PM} \bar{K}^M = 0 \,$

and this implies

$$K^M \nabla_M \bar{K}_N = K^M \partial_M \bar{K}_N, \qquad \bar{K}^M \nabla_M K_N = \bar{K}^M \partial_M K_N$$

• Using the parametrization of K and \overline{K} on a flat background,

$$K_M = \frac{1}{\sqrt{2}} \begin{pmatrix} l^{\mu} \\ l_{\mu} \end{pmatrix}, \qquad \bar{K}_M = \frac{1}{\sqrt{2}} \begin{pmatrix} \bar{l}^{\mu} \\ -\bar{l}_{\mu} \end{pmatrix}$$

The on-shell constraint on l and \overline{l} is written as

$$\begin{split} l^{\mu}\partial_{\mu}\bar{l}_{\nu} &= 0 \,, \qquad \bar{l}^{\mu}\partial_{\mu}l_{\nu} = 0 \,, \\ l^{\mu}\partial_{\mu}f &= 0 \,, \qquad \bar{l}^{\mu}\partial_{\mu}f = 0 \,, \end{split}$$

 Interestingly, these can be interpreted as the parallel transport equations along the *l* and *l* with the torsionful connections.

$$\bar{l}^{\mu} \nabla^{+}_{\mu} l_{\nu} = 0, \qquad l^{\mu} \nabla^{-}_{\mu} \bar{l}_{\nu} = 0,$$

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where $\nabla^{\pm}_{\mu} = \nabla_{\mu} \pm \frac{1}{2} H_{\mu}$ and $H_{\mu\nu\rho} = 3\partial_{[\mu}B_{\nu\rho]}$.

Equations of motion

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Substituting the KS ansatz into the equations of motion in the flat backgrounds

$$-2\kappa\partial_{K}\partial_{L}(\varphi K^{K}\bar{K}^{L})+4\kappa\mathcal{H}_{0}^{KL}\partial_{K}\partial_{L}f-4\kappa^{2}\mathcal{H}_{0}^{KL}\partial_{K}f\partial_{L}f=0.$$

and

$$\kappa \Big[-\frac{1}{2} \mathcal{H}_{0}^{MN} \partial_{M} \partial_{N} (\varphi K_{(K} \bar{K}_{L)}) + \partial_{M} \partial_{N} (\varphi K^{N} \bar{K}_{(K)}) P_{0L})^{M} - \partial_{M} \partial_{N} (\varphi K_{(K} \bar{K}^{N}) \bar{P}_{0L})^{M} + 4 P_{0(K}{}^{M} \bar{P}_{0L})^{N} \partial_{M} \partial_{N} f \Big] + \kappa^{2} \mathcal{H}_{0}^{MN} \partial_{M} f \partial_{N} (\varphi K_{(K} \bar{K}_{L})) = 0.$$

- Unlike the conventional KS formalism in GR, the equations are quadratic in κ due to the presence of f. If we set f = 0, field equations reduce to linear equations
- If we consider the power series expansion of f, then the linear terms are enough to determine φ , l and \overline{l} or $g_{\mu\nu}$ and $B_{\mu\nu}$ completely.

In terms of d-dimensional vector indices, the field equations reduces to

 $\Box \left(\varphi l_{\mu} \bar{l}_{\nu}\right) - \partial^{\rho} \partial_{\mu} \left(\varphi l_{\rho} \bar{l}_{\nu}\right) - \partial^{\rho} \partial_{\nu} \left(\varphi l_{\mu} \bar{l}_{\rho}\right) + \partial_{\mu} \partial_{\nu} \left(\varphi l \cdot \bar{l}\right) + \partial_{\mu} \partial_{\nu} H = 0.$

- Note that *R_{µν}* is not symmetric tensor:
 - symmetric part \rightarrow eom of g
 - antisymmetric part \rightarrow eom of B
- It is interesting that the generalized KS ansatz for $g_{\mu\nu}$ and $B_{\mu\nu}$ is not linear in κ , l^{μ} and \bar{l}^{μ} , but the field equations are linear in these fields.

• Curved background generalization is straightforward.

Comments on dilaton

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 So far we have considered a flat background with the Cartesian coordinates only. In a coordinate independent form in terms of the covariant derivative

$$\begin{aligned} \mathcal{R} &= \kappa \Big[\,\tilde{\nabla}_{0\mu} \tilde{\nabla}_{0\nu} \big(\varphi l^{\mu} \bar{l}^{\nu} \big) - 4 \tilde{\nabla}_{0}^{\mu} \partial_{\mu} f \,\Big] + 4\kappa^{2} \partial^{\mu} f \partial_{\mu} f = 0 \,, \\ \mathcal{R}_{\mu\nu} &= \frac{\kappa}{4} \Big[\,\tilde{\nabla}_{0\rho} \tilde{\nabla}_{0}^{\rho} \big(\varphi l_{\mu} \bar{l}_{\nu} \big) - \tilde{\nabla}_{0}^{\rho} \tilde{\nabla}_{0\mu} \big(\varphi l_{\rho} \bar{l}_{\nu} \big) - \tilde{\nabla}_{0}^{\rho} \tilde{\nabla}_{0\nu} \big(\varphi l_{\mu} \bar{l}_{\rho} \big) + 4 \tilde{\nabla}_{0\mu} \partial_{\nu} f \,\Big] \\ &- \frac{\kappa^{2}}{2} \partial^{\rho} f \tilde{\nabla}_{\rho} \big(\varphi l_{\mu} \bar{l}_{\nu} \big) \,. \end{aligned}$$

where $\tilde{\nabla}_{0\mu}$ is a covariant derivative for a flat background in an arbitrary coordinate system.

• Note that the DFT dilaton is not a scalar field, but a density that transform under a coordinate transform $x^{\mu} \rightarrow x'^{\mu}(x)$ as

$$e^{-2d} \to e^{-2d'} = \left| \frac{\partial x'}{\partial x} \right| e^{-2d}$$

 We can find a new coordinate x^{'µ} that makes the new DFT dilaton d' vanish by requiring that the Jacobian is e^{2d}.

$$\left. \frac{\partial x'}{\partial x} \right| = e^{2d}$$

Thus, for a given d, we can make the DFT dilaton vanishes.

 As discussed, all the higher order terms in κ in the field equations include *f*. Using this fact, if we perform a coordinate transformation, the equations of motion reduces to linear

$$\begin{split} \tilde{\mathcal{R}} &= \kappa \check{\nabla}_{0\mu} \check{\nabla}_{0\nu} \left(\varphi l^{\mu} \bar{l}^{\nu} \right) = 0 \,, \\ \tilde{\mathcal{R}}_{\mu\nu} &= \frac{\kappa}{4} \Big[\, \check{\nabla}_{0\rho} \check{\nabla}^{\rho}_{0} \left(\varphi l_{\mu} \bar{l}_{\nu} \right) - \check{\nabla}^{\rho}_{0} \check{\nabla}_{0\mu} \left(\varphi l_{\rho} \bar{l}_{\nu} \right) - \check{\nabla}^{\rho}_{0} \check{\nabla}_{0\nu} \left(\varphi l_{\mu} \bar{l}_{\rho} \right) \, \Big] \,, \end{split}$$

where $\check{\nabla}_{0\mu}$ is a covariant derivative for a flat space with the particular coordinate where f' = 0.

However, it is not practical in solving eom, but useful for classical double copy

Killing Spinor equation

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- The Killing spinor equation reduce the supergravity field equations to first order in derivatives. Combined with the generalized KS ansatz, it will lead to linear equations.
- The SUSY variation of fermions provides the Killing spinor equations, which are

$$\delta\rho = -\gamma^{p}\mathcal{D}_{p}\varepsilon = -\gamma^{p}V_{p}{}^{M}\partial_{M}\varepsilon - \frac{1}{4}V^{M}{}_{p}\Phi_{Mmn}\gamma^{pmn}\varepsilon - \frac{1}{2}V^{Mm}\Phi_{Mmn}\gamma^{n}\varepsilon = 0,$$

$$\delta\psi_{\bar{p}} = \bar{V}^{M}{}_{\bar{p}}\mathcal{D}_{M}\varepsilon = \bar{V}^{M}{}_{\bar{p}}\partial_{M}\varepsilon + \frac{1}{4}\bar{V}^{M}{}_{\bar{p}}\Phi_{Mmn}\gamma^{mn}\varepsilon = 0,$$

• For simplicity, let us choose ε as a Killing spinor for the background geometry satisfying

$$\partial_p \phi \gamma^p \varepsilon_0 + \frac{1}{12} \tilde{H}_{mnp} \gamma^{mnp} \varepsilon_0 = 0,$$

 $\tilde{D}_{\bar{p}}^+ \varepsilon_0 = 0,$

where ε_0 is the background Killing spinor.

Then the Killing spinor equations are greatly simplified as

$$\left(\partial_{\mu}\Psi + \frac{1}{2}\tilde{D}^{+}_{\nu}(\varphi' l_{\mu}\bar{l}^{\nu})\right)\gamma^{\mu}\varepsilon_{0} = 0\,,$$

and

$$\left(\tilde{D}_{\mu}\left(\varphi l_{\nu}\bar{l}_{\rho}\right) - \frac{1}{2}\tilde{H}_{\mu\rho\sigma}\left(\varphi l_{\nu}\bar{l}^{\sigma}\right)\right)\gamma^{\mu\nu}\varepsilon_{0} = 0.$$

where $\Psi=e^{-2\kappa f},\,\varphi'=e^{-2\kappa f}\varphi$ and ε_0 is the background Killing spinor.

For the flat background case

$$\begin{split} & \left(\partial_{\mu}\Psi + \frac{1}{2}\partial_{\nu}\left(\varphi'l_{\mu}l^{\bar{\nu}}\right)\right)\gamma^{\mu}\varepsilon_{0} = 0\,,\\ & \partial_{\mu}\left(\varphi l_{\nu}\bar{l}_{\rho}\right)\gamma^{\mu\nu}\varepsilon_{0} = 0\,, \end{split}$$

where ε_0 is a constant spinor.

 These equations are remarkably simple, and much easier to solve than the full Killing spinor equations.

Cassical double copy

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- The KLT and BCJ relations indicate that not only the pure Einstein equation, but also the field equations of entire massless NS-NS sector should be related to the gauge theory.
- Suppose that the full geometry admits at least one Killing vector ξ^μ.
- We can locally choose a coordinate system $x^{\mu} = \{x^i, y\}$ such that the Killing vector is a constant, $\xi^{\mu} = \partial x^{\mu} / \partial y = \delta_y^{\mu}$. The Killing vector ensures the following identities from the torsion free condition

$$\tilde{\nabla}_{\mu}\xi_{\nu} = \tilde{\nabla}_{[\mu}\xi_{\nu]} = \partial_{[\mu}\xi_{\nu]} = 0\,.$$

 Consider the Lie derivative of an arbitrary rank-*n* tensor F_{μ1μ2}...μ_n with respect to a constant Killing vector ξ^μ

$$\mathcal{L}_{\xi}F_{\mu_{1}\mu_{2}\cdots\mu_{n}} = \xi^{\rho}\partial_{\rho}F_{\mu_{1}\mu_{2}\cdots\mu_{n}} + \sum_{i=1}^{n}\partial_{\mu_{i}}\xi^{\rho}F_{\mu_{1}\cdots\mu_{i-1}\rho\mu_{i+1}\cdots\mu_{n}}$$
$$= \xi^{\rho}\tilde{\nabla}_{\rho}F_{\mu_{1}\mu_{2}\cdots\mu_{n}} + \sum_{i=1}^{n}\tilde{\nabla}_{\mu_{i}}\xi^{\rho}F_{\mu_{1}\cdots\mu_{i-1}\rho\mu_{i+1}\cdots\mu_{n}} = 0,$$

 Since we are assuming that the Killing vector is covariantly constant, this shows that

$$\xi^{\rho}\tilde{\nabla}_{\rho}F_{\mu_{1}\mu_{2}\cdots\mu_{n}}=0\,.$$

• We also normalize l_{μ} and \bar{l}_{μ} as

$$\xi \cdot l = \xi \cdot \bar{l} = 1$$

Single Copy

 Classical double copy is achieved by contracting the constant Killing vector ξ^μ with the generalized Ricci tensor in the specific coordinate

$$\check{\mathcal{R}}_{\mu\nu} = \frac{\kappa}{4} \Big[\check{\nabla}_{0\rho} \check{\nabla}^{\rho}_{0} \big(\varphi l_{\mu} \bar{l}_{\nu} \big) - \check{\nabla}^{\rho}_{0} \check{\nabla}_{0\mu} \big(\varphi l_{\rho} \bar{l}_{\nu} \big) - \check{\nabla}^{\rho}_{0} \check{\nabla}_{0\nu} \big(\varphi l_{\mu} \bar{l}_{\rho} \big) \Big] = 0$$

Since *R_{μν}* is not symmetric tensor, we get three independent equations as follows:

$$\begin{split} \xi^{\nu} \tilde{\mathcal{R}}_{\mu\nu} &= \frac{\kappa}{4} \Big[\, \tilde{\nabla}_{0}{}^{\rho} \tilde{\nabla}_{0\rho} \big(\varphi l_{\mu} \big) - \tilde{\nabla}_{0}{}^{\rho} \tilde{\nabla}_{0\mu} \big(\varphi l_{\rho} \big) \, \Big] \,, \\ \xi^{\mu} \tilde{\mathcal{R}}_{\mu\nu} &= \frac{\kappa}{4} \Big[\, \tilde{\nabla}_{0}{}^{\rho} \tilde{\nabla}_{0\rho} \big(\varphi \bar{l}_{\nu} \big) - \tilde{\nabla}_{0}^{\rho} \tilde{\nabla}_{0\nu} \big(\varphi \bar{l}_{\rho} \big) \, \Big] \,, \end{split}$$

• we identify φl_{μ} and $\varphi \bar{l}_{\mu}$ with gauge fields

$$A_{\mu} = \varphi l_{\mu} \,, \qquad \bar{A}_{\mu} = \varphi \bar{l}_{\mu}$$

• Then $\xi^{\nu} \mathcal{R}_{\mu\nu}$ and $\xi^{\mu} \mathcal{R}_{\mu\nu}$ reduce to a pair of Maxwell equations

$$\partial^{\mu}F_{\mu\nu} = 0, \qquad \partial^{\mu}\bar{F}_{\mu\nu} = 0,$$

where $F_{\mu\nu}$ and $\bar{F}_{\mu\nu}$ are the field strength of the Maxwell fields of A_{μ} and \bar{A}_{μ} respectively,

Zeroth copy

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• Contracting ξ^{μ} with all the free indices of $\mathcal{R}^{(1)}_{\mu\nu}$, we make a scalar equation

$$\xi^{\mu}\xi^{\nu}\check{\mathcal{R}}_{\mu\nu} = \Box\varphi = 0\,,$$

 Monteiro, O'Connell and White identified φ as the biadjoint scalar field [Cachazo,He,Yuan 2013]

$$\Phi^{aa'} = \varphi c^a \bar{c}^{a'}$$

where c^a and $\bar{c}^{\bar{a}}$ are color index vectors for Lie group G_1 and G_2 .

• It can be understood as a linearized equation of motion for $\Phi^{aa'}$

$$\partial^2 \Phi^{aa'} - g f^{abc} f^{a'b'c'} \Phi^{bb'} \Phi^{cc'} = 0$$

• This shows that the generalized KS type solution can be written in terms of the solutions of the two independent Maxwell equation and free scalar field equations

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• On a flat background, Killing spinor equation for gravitino is given by

$$\kappa \partial_{[m} \left(\phi l_{n]} \bar{l}_{\mu} \right) \gamma^{mn} \varepsilon = 0$$

• contraction with a Killing vector ξ^{μ}

$$F_{\mu\nu}\gamma^{\mu\nu}\varepsilon=0\,.$$

• This is the typical BPS equation of N = 1 SYM. This shows the classical double copy is still valid for supersymmetric backgrounds

Curved Background generalization

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- Recently Adamo, Casali, Mason, Nekova showed that BCJ color-kinematic duality can be extended pp-wave background.
- The classical double copy is well studied in a flat background, but curved background generalization was not obvious. (only for some simple cases, (A)dS background etc [Gonalez, Penco, Trodden, 2017], [Bahjat-Abbas, Luna, White, 2017])
- The KLT relation in a curved background It is not clear how to define scattering amplitude in curved background in general. (time-dependent backgrounds, nonasymptotic flat spaces)
- Classical double copy in general background may give a clue, however, curved background generalization is an open problem.

Generalization to heterotic supergravity

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Heterotic supergravity: relaxed null condition

$$\mathcal{H}_{\hat{M}\hat{N}} = \mathcal{H}_{0\hat{M}\hat{N}} + \kappa\varphi \left(K_{\hat{M}}\bar{K}_{\hat{N}} + K_{\hat{N}}\bar{K}_{\hat{M}} \right),$$

In terms of the heterotic supergravity fields

$$\begin{split} g^{\mu\nu} &= \tilde{g}^{\mu\nu} + \kappa \varphi l^{(\mu} \bar{l}^{\nu)} ,\\ g_{\mu\nu} &= \tilde{g}_{\mu\nu} - \frac{\kappa \varphi}{1 + \frac{\kappa \varphi}{2} (l \cdot \bar{l})} l_{(\mu} \bar{l}_{\nu)} + \frac{1}{4} \Big(\frac{\kappa \varphi}{1 + \frac{\kappa \varphi}{2} (l \cdot \bar{l})} \Big)^2 (\bar{l} \cdot \bar{l}) l_{\mu} l_{\nu} ,\\ B_{\mu\nu} &= \tilde{B}_{\mu\nu} + \frac{\kappa \varphi}{1 + \frac{\kappa \varphi}{2} (l \cdot \bar{l})} \Big(l_{[\mu} \bar{l}_{\nu]} - \sqrt{\frac{\alpha'}{2}} \tilde{A}_{[\mu}{}^{\alpha} l_{\nu]} j_{\alpha} \Big) ,\\ A_{\mu\alpha} &= \tilde{A}_{\mu\alpha} + \frac{1}{\sqrt{2\alpha'}} \frac{\kappa \varphi}{1 + \frac{\kappa \varphi}{2} (l \cdot \bar{l})} l_{\mu} j_{\alpha} , \end{split}$$

where l is a null vector, but \overline{l} is not.

• It is possible to couple U(1) gauge fields.



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Chiral null model

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- A class of string backgrounds which have one conserved chiral null current on the world sheet. [Horowitz, Tseytlin, 1994]
- It is a generalization of the gravitational wave and fundamental string background and is exact in the α' expansion.
- In the target space they have a null Killing vector and unbroken supersymmetries.
- Special cases are the Taub-NUT geometry and rotating black holes.
- The explicit geometry is given by

$$ds^{2} = F(x^{i})du \left(dv + K(u, x^{i})du + 2V_{i}(u, x^{i})dx^{i} \right) + dx^{i}dx^{i},$$

$$B_{uv} = F(x^{i}), \qquad B_{ui} = 2F(x^{i})V_{i}(u, x^{i}),$$

$$\phi = \phi(u) + \frac{1}{2}\log F(x^{i}),$$

• This fits into the generalized Kerr-Schild ansatz in a flat background.

$$\mathrm{d}s^{2} = \mathrm{d}u\mathrm{d}\tilde{v} + \mathrm{d}x^{i}\mathrm{d}x^{i} + (F-1)\mathrm{d}u\left(\mathrm{d}\tilde{v} - \tilde{V}_{i}\tilde{V}^{i}\mathrm{d}u + \tilde{V}_{i}\mathrm{d}x^{i}\right),$$

where

$$V_i = \tilde{V}_i + \frac{1}{2} \partial_i X , \qquad v = \tilde{v} - X(x, u) ,$$
$$X(x, u) = \int^u \left(K + \frac{4F}{(F-1)} \tilde{V}_i \tilde{V}^i \right) (\vec{x}, u') du' ,$$

• The associated φ and null vectors l and \overline{l} can be easily read off

$$\begin{aligned} \kappa \varphi &= F^{-1} - 1 \,, \\ l_u &= 1 \,, \\ \bar{l}_u &= -\left(\frac{2F}{F-1}\right)^2 \tilde{V}_i \tilde{V}^i \,, \qquad \bar{l}_{\bar{v}} = 1 \,, \qquad \bar{l}_i = \frac{2F}{F-1} \tilde{V}_i \,, \end{aligned}$$

and one can easily show that l and \overline{l} are orthogonal with respect to the flat background metric.

• l and \overline{l} satisfy the generalized geodesic constraint

$$l^{\mu}\partial_{\mu}\bar{l}_{\nu}=0\,,\qquad \bar{l}^{\mu}\partial_{\mu}l_{\nu}=0$$

• Equations of motion imply, $\kappa f = \phi(u)$

$$\partial_i \partial^i F^{-1} = 0 ,$$

$$-\partial^i \partial_i K + 2\partial^i \partial_u V_i + 4F^{-1} \partial_u^2 \phi = 0 ,$$

$$-4\partial^j \mathcal{F}_{ji} + 4\partial_u \phi \partial_i F^{-1} = 0 .$$

where $\mathcal{F}_{ij} = \partial_i V_j - \partial_j V_i$.

• This is the same exactly with the equation derived by Callan, Maldacena and Peet.

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F1-NS5 system

- To examine our formalism in the case of a curved background, let us consider superposition of a number Q₁ of fundamental strings and Q₅ of NS5 brane system.
- Wrap a number Q₅ of NS5-branes on T⁵ along x₅, ... x₉. The fundamental strings wrap one of the directions of the torus along x⁵ direction.

$$ds^{2} = F_{1}^{-1} \left(-dt^{2} + dx_{5}^{2} \right) + F_{5} \left(dx_{1}^{2} + \dots + dx_{4}^{2} \right) + dx_{6}^{2} + \dots dx_{9}^{2}$$

$$e^{-2\phi} = g_{s}F_{1}F_{5}^{-1}$$

$$B_{05} = F_{1}^{-1} - 1,$$

$$H_{ijk} = \epsilon_{ijkl}\partial^{l}F_{5}, \qquad i, j, k, l = 1, 2, 3, 4$$

where ϵ_{ijkl} is the flat space epsilon tensor and

$$F_1 = 1 + rac{16\pi^4 {lpha'}^3 Q_1}{g_s^2 V_4 r^2} \,, \qquad F_5 = 1 + rac{lpha' Q_5}{r^2} \,,$$

here V_4 is the volume of the T^4 .

The NS5-brane background is treated as a background

$$d\tilde{s}^{2} = -dt^{2} + F_{5}(dx_{1}^{2} + \dots + dx_{4}^{2}) + dx_{5}^{2} + \dots dx_{9}^{2}$$

$$e^{-2\tilde{\phi}} = g_{s}F_{5}^{-1}$$

$$\tilde{H}_{ijk} = \epsilon_{ijkl}\partial^{l}F_{5}, \qquad i, j, k, l = 1, 2, 3, 4$$

where we put tilde for all the background quantities.

We can read off the corresponding generalized KS ansatz

$$\varphi = F_1 - 1$$
, $l = dt + dx_5$, $\bar{l} = -dt + dx_5$.

Note that l_{μ} and \bar{l}_{μ} are orthogonal to the background 3-form flux, $\tilde{H}_{\mu\nu\rho}$

$$l^{\mu}\tilde{H}_{\mu\nu\rho} = \bar{l}^{\mu}\tilde{H}_{\mu\nu\rho} = 0\,,$$

and $\phi = \frac{1}{\sqrt{-g}}$ of

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- Internal momenta and axion charges are interchangeable via T-duality [Horne, Horowitz, Steif, 1991]
- Consider a 5-dimensional uncharged black string, Schwarzschild $\times S^1$.
- 1. Boost along the circle direction, y ⇒ Generates Off diagonal terms in metric
 2. Take a *T*-duality along the y direction
- The explicit geometry is given by

$$ds^{2} = \left(1 + \frac{2mS^{2}}{r}\right)^{-1} \left[-\left(1 - \frac{2m}{r}\right) dt^{2} + dy^{2} \right] + \left(1 - \frac{2m}{r}\right)^{-1} dr^{2} + r^{2} d\Omega,$$

$$B_{yt} = \frac{C}{S} \left(1 + \frac{2mS^{2}}{r}\right)^{-1}, \qquad e^{-2\phi} = 1 + \frac{2mS^{2}}{r},$$

where $C = \cosh \alpha$ and $S = \sinh \alpha$, and α is a boost parameter.

Uncharged black string solution:

$$ds^{2} = -\left(1 - \frac{2M}{r}\right)d\tilde{t}^{2} + \left(1 - \frac{2M}{r}\right)^{-1}dr^{2} + r^{2}d\Omega^{2} + dy^{2}.$$
 (1)

 Using the conventional Kerr-Schild ansatz for Schwarzschild BH in Eddington-Finkelstein coordinate, we have

$$\mathrm{d}s^2 = -\mathrm{d}\hat{t}^2 + \mathrm{d}r^2 + r^2\mathrm{d}\Omega^2 + \mathrm{d}y^2 - \kappa\varphi(\mathrm{d}\hat{t} + \mathrm{d}r)^2, \qquad \kappa\varphi = -\frac{2M}{r}, \qquad (2)$$

where $\hat{t} = \tilde{t} + (r_* - r)$ and r_* is the tortoise coordinate defined

$$r_* = r + 2M \log \left| \frac{r}{2M} - 1 \right|, \qquad \mathrm{d}r_* = \mathrm{d}r \left(1 - \frac{2M}{r} \right)^{-1}.$$
 (3)

• Applying a boost along the *y*-direction, $\hat{t} \rightarrow \hat{t} \cosh \alpha + y \sinh \alpha$,

 $y \rightarrow \hat{t} \sinh \alpha + y \cosh \alpha$, we get

$$ds^{2} = -d\hat{t}^{2} + dr^{2} + r^{2}d\Omega^{2} + dy^{2} - \kappa\varphi (Cd\hat{t} + Sdy + dr)^{2}, \qquad (4)$$

and the null vectors are identical, $l = \overline{l} = C d\hat{t} + S dy + dr$.

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 Let us take a T-duality along the *y*-direction. According to the Buscher's rule for the null vectors, the null vectors split into

$$l \to l' = C d\hat{t} + S dy + dr$$
, $\bar{l} \to \bar{l}' = C d\hat{t} - S dy + dr$,

and $l \cdot \overline{l} = -2S^2$. The corresponding metric and Kalb-Ramond field are

$$ds'^{2} = \left(1 + \frac{2MS^{2}}{r}\right)^{-1} \left[-\left(1 - \frac{2M}{r}\right) d\hat{t}^{2} + \frac{4MC}{r} d\hat{t} dr + \left(1 + \frac{2MC^{2}}{r}\right) dr^{2} + dy^{2} \right]$$
$$B' = \left(-\frac{C}{S} + \frac{2MCS}{r} \left(1 + \frac{2MS^{2}}{r}\right)^{-1}\right) d\hat{t} \wedge dy - \frac{2MS}{r} \left(1 + \frac{2MS^{2}}{r}\right)^{-1} dr \wedge dy.$$

• Finally, we make a further coordinate transform $\hat{t} = t + C(r_* - r)$

$$ds^{2} = -d\hat{t}^{2} + dr^{2} + r^{2}d\Omega^{2} + dy^{2} + \frac{2M}{r + 2MS^{2}} (Cd\hat{t} + Sdy + dr) (Cd\hat{t} - Sdy + dr)$$
$$B = \frac{2M}{r + 2MS^{2}} (Cd\hat{t} + Sdy + dr) \wedge (Cd\hat{t} - Sdy + dr) .$$

 In this example, the DFT dilaton vanishes, thus the generalized KS field equations become linear.

- A novel solution generating technique in supergravities via generalized Kerr-Schild method in DFT
- Classical double copy including $B_{\mu\nu}$ and dilaton
- Classical double copy in Killing spinor equation
- Including RR sector, Introducing U(1) gauge fields using Kaluza-Klein reduction, Gauged supergravity extension via Scherk-Schwarz reduction.
- M-theory extension: Exceptional field theories $(SL(5), SO(5, 5), E_6, E_7 \text{ and } E_8)$
- Finding the most general solutions in a flat or curved backgrounds and their physical interpretations. Applications to AdS/CFT?
- Scattering amplitude computation in the DFT language. Extension double copy structure to curved backgrounds.

Thank you

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