Complement on Thermodynamics of Taub-NUT-AdS spacetime

Miok Park

Korea Institute for Advanced Study (KIAS), Seoul, S. Korea

"String theory, gravity and cosmology (SGC2020)", @ APCTP, Pohang, S. Korea

November 20, 2020

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□

Pin point

1. Introduction to Taub-NUT-AdS spacetime

2. Non-charged TB-AdS (Analogy)

Sch-AdS black hole	non-charged TB-AdS
Small black hole/	Small TB-AdS/
Negative heat capacity	Negative heat capacity
Large black hole/	Large TB-AdS/
Positive heat capacity	Positive heat capacity

3. Dyonic TB-AdS (Differences)

dyonic RN black hole	dyonic TB-AdS
Regularity condition on A	Regularity condition on A
is NOT essential	is essential
for thermodynamics	for thermodynamics
electric potential $\Phi_E^{(1)} = \Phi_E^{(2)}$	electric potential $\Phi_E^{(1)} eq \Phi_E^{(2)}$
Q and P charges	Q and P charges
are independent	are NOT independent

4. Zero temperature limit of TB-AdS

Introduction to Taub-NUT(-AdS) spacetime

・ロト・<・<・<・・

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

AdS spacetime



$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

- AdS spacetime
- AdS black holes



$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

- AdS spacetime
- AdS black holes
- Taub-NUT-AdS spacetime

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

- AdS spacetime
- AdS black holes
- Taub-NUT-AdS spacetime
- Eguchi-Hanson-AdS spacetime

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

AdS spacetime : maximally symmetric spacetime

$$ds_{AdS}^{2} = -\left(1 + \frac{r^{2}}{l^{2}}\right)dt^{2} + \left(1 + \frac{r^{2}}{l^{2}}\right)^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$
$$R_{abcd} = -\frac{1}{l^{2}}\left(g_{ac}g_{bd} - g_{ad}g_{bc}\right)$$

- AdS black holes
- Taub-NUT-AdS spacetime
- Eguchi-Hanson-AdS spacetime

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

AdS spacetime : maximally symmetric spacetime

$$ds_{AdS}^{2} = -\left(1 + \frac{r^{2}}{l^{2}}\right)dt^{2} + \left(1 + \frac{r^{2}}{l^{2}}\right)^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$
$$R_{abcd} = -\frac{1}{l^{2}}\left(g_{ac}g_{bd} - g_{ad}g_{bc}\right)$$

AdS black holes : Asymptotically AdS spacetime

$$\lim_{r \to \infty} ds_{\text{AdS-BH}}^2 \sim ds_{\text{AdS}}^2, \qquad R_{abcd} \sim -\frac{1}{l^2} \left(g_{ac} g_{bd} - g_{ad} g_{bc} \right)$$

- Taub-NUT-AdS spacetime
- Eguchi-Hanson-AdS spacetime

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

AdS spacetime : maximally symmetric spacetime

$$ds_{AdS}^{2} = -\left(1 + \frac{r^{2}}{l^{2}}\right)dt^{2} + \left(1 + \frac{r^{2}}{l^{2}}\right)^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$
$$R_{abcd} = -\frac{1}{l^{2}}\left(g_{ac}g_{bd} - g_{ad}g_{bc}\right)$$

AdS black holes : Asymptotically AdS spacetime

$$\lim_{r \to \infty} ds_{\text{AdS-BH}}^2 \sim ds_{\text{AdS}}^2, \qquad R_{abcd} \sim -\frac{1}{l^2} \left(g_{ac} g_{bd} - g_{ad} g_{bc} \right)$$

Taub-NUT-AdS spacetime : Asymptotically locally AdS spacetime

$$\lim_{r \to \infty} ds_{\text{TN-AdS}}^2 \neq ds_{\text{AdS}}^2, \qquad R_{abcd} \sim -\frac{1}{l^2} \left(g_{ac} g_{bd} - g_{ad} g_{bc} \right)$$

Eguchi-Hanson-AdS spacetime

$$R_{ab} - \frac{1}{2}g_{ab}R + \Lambda g_{ab} = 0,$$

AdS spacetime : maximally symmetric spacetime

$$ds_{AdS}^{2} = -\left(1 + \frac{r^{2}}{l^{2}}\right)dt^{2} + \left(1 + \frac{r^{2}}{l^{2}}\right)^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$
$$R_{abcd} = -\frac{1}{l^{2}}\left(g_{ac}g_{bd} - g_{ad}g_{bc}\right)$$

AdS black holes : Asymptotically AdS spacetime

$$\lim_{r \to \infty} ds_{\text{AdS-BH}}^2 \sim ds_{\text{AdS}}^2, \qquad R_{abcd} \sim -\frac{1}{l^2} \left(g_{ac} g_{bd} - g_{ad} g_{bc} \right)$$

Taub-NUT-AdS spacetime : Asymptotically locally AdS spacetime

$$\lim_{r \to \infty} ds_{\text{TN-AdS}}^2 \neq ds_{\text{AdS}}^2, \qquad R_{abcd} \sim -\frac{1}{l^2} \left(g_{ac} g_{bd} - g_{ad} g_{bc} \right)$$

Eguchi-Hanson-AdS spacetime

$$\lim_{r \to \infty} ds_{\text{EH-AdS}}^2 \neq ds_{\text{AdS}}^2, \qquad R_{abcd} \neq -\frac{1}{l^2} \left(g_{ac} g_{bd} - g_{ad} g_{bc} \right)$$

▲□▶ ▲□▶ ▲三▶ ▲三▶ 三三 のへで

Misner string for $\kappa = 1$

$$e^{\hat{t}} = F(r)[dt + 4s\sin^2\left(\frac{\theta}{2}\right)d\phi], \qquad e^{\hat{r}} = F(r)^{-1}dr,$$
$$e^{\hat{\theta}} = (r^2 + s^2)^{1/2}d\theta, \qquad e^{\hat{\phi}} = (r^2 + s^2)^{1/2}\sin(\theta)d\phi$$

string singularity : $(\nabla t)^2$ is regular as $\theta \to 0$, it diverges as $\theta \to \pi$

$$-(\nabla t)^{2} = \frac{1}{F(r)^{2}} - \frac{(2s)^{2}}{r^{2} + s^{2}} \tan^{2}\left(\frac{\theta}{2}\right)$$

On constant r hypersurface

$$e^{\hat{t}} = dt_N + 4s\sin^2\left(\frac{\theta}{2}\right)d\phi, \qquad (0 \le \theta < \pi)$$
$$e^{\hat{t}} = dt_S - 4s\cos^2\left(\frac{\theta}{2}\right)d\phi, \qquad (0 < \theta \le \pi)$$

► combine two patches → time coordinate becomes periodic

$$t_N \equiv t_N + 8\pi s$$

Euclidean Taub-NUT-AdS space

Euclidean Taub-NUT-AdS is written

$$ds_E^2 = F_E(r)^2 \left[d\tau + 4\chi \sin^2\left(\frac{\theta}{2}\right) d\phi \right]^2 + \frac{dr^2}{F_E(r)^2} + (r^2 - \chi^2) d\Omega_2^2,$$

$$F_E(r)^2 = \frac{l^2(r^2 - \chi^2)^2 + (\kappa - 4l^{-2}\chi^2)(r^2 + \chi^2) - 2Mr}{r^2 - \chi^2}$$

• Horizon is located at $F_E(r_+) = 0$

- NUT solution occurs when the fixed point set of ∂_{τ} is zero dimensional, e.g. $r_{+} = \chi$ (denoted as TN-AdS)
- ► Bolt solution occurs when the fixed point set of ∂_{τ} is two dimensional, e.g. $r_+ \neq \chi$ (denoted as TB-AdS)

Thermodynamics of non-charged TB/TN-AdS space

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ □ のへぐ

non-charged TB/TN-AdS

The Euclidean action and metric

$$I_E = -\frac{1}{16\pi G_4} \int d^4x \sqrt{g} \left(R + \frac{6}{l^2} \right) - \frac{1}{8\pi G_4} \int d^3x \sqrt{h} K$$
(1)

$$I_{\rm ren.} = I_E + \frac{1}{8\pi G_4} \int d^3x \sqrt{h} \left(\frac{2}{l} + \frac{l}{2}R_3\right),$$
 (2)

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R - \frac{3}{l^2}g_{\mu\nu} = 0$$
(3)

$$ds_{\rm E}^2 = f_E(r)(d\tau + 2\chi\lambda(\theta)d\phi)^2 + \frac{dr^2}{f_E(r)} + (r^2 - \chi^2)(d\theta^2 + Y(\theta)^2 d\phi^2), \quad (4)$$
$$f_E = \frac{l^{-2}(r^2 - \chi^2)^2 + (\kappa - 4l^{-2}\chi^2)(r^2 + \chi^2) - 2Mr}{r^2 - \chi^2}.$$
$$\left(\cos\theta \qquad \int \sin\theta \qquad \text{for } \kappa = 1 \right)$$

$$\lambda(\theta) = \begin{cases} -\theta & , \quad Y(\theta) = \begin{cases} 1 & \text{for } \kappa = 0\\ -\cosh \theta & & \end{cases} \text{ for } \kappa = -1, \end{cases}$$

< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > <

Thermodynamics of non-charged TN-AdS

- ► Taub-"NUT"-AdS solution when $r_+ = \chi$ where $f_E(r_+) = 0$
- metric function

$$f_n(r) = \frac{l^{-2}(r-\chi)^2(r+3\chi)\chi + \kappa(r-\chi)\chi}{\chi(r+\chi)}, \qquad M_n = \chi\kappa - \frac{4\chi^3}{l^2}$$

Hawking temperature

$$T_n = \frac{1}{4\pi} f'_n(\chi) = \frac{\kappa}{8\pi\chi}$$

Entropy and Energy

$$S = \left(\beta \frac{\partial}{\partial \beta} - 1\right) I_{\text{ren.}} = \frac{2\pi\omega\chi^2}{G} \left(\kappa - \frac{6\chi^2}{l^2}\right),$$
$$E = \partial_\beta I_{\text{ren.}} = \frac{\chi\omega}{2G} \left(\kappa - \frac{4\chi^2}{l^2}\right)$$

First law and free energy

$$dE = TdS, \qquad F = E - TS$$

▲□▶▲□▶▲□▶▲□▶ ▲□▶

- ► Taub-Bolt-AdS solution when $r_+ \neq \chi$ where $f_E(r_+) = 0$
- mass parameter

$$M_b = \frac{-6\chi^2 r_b^2 + r_b^4 - 3\chi^4 + \kappa \left(r_b^2 + \chi^2\right) l^2}{2l^2 r_b}.$$

Hawking temperature

$$T = \frac{1}{4\pi} f'_{\rm E}(\chi) = \frac{M_b \left(r_b^2 + \chi^2\right) + \left(9\chi^4 r_b - 2\chi^2 r_b^3 + r_b^5\right) l^{-2} - 2\kappa r_b \chi^2}{2\pi \left(\chi^2 - r_b^2\right)^2}.$$

Entropy and Energy

$$S_{(\kappa \neq 1)} = \left(\beta \frac{\partial}{\partial \beta} - 1\right) I_{\text{ren.}} = \frac{\pi \omega}{2G} (r_b^2 - \chi^2)$$
$$E_{(\kappa \neq 1)} = \partial_\beta I_{\text{ren.}} = \frac{\omega}{4G} \left(\kappa + \frac{r_b^3 - 3r_b \chi^2}{l^2}\right)$$

First law and free energy

$$dE = TdS, \qquad F = E - TS$$

▲□▶▲□▶▲□▶▲□▶ ▲□▶

Hawking temperature

$$T = \frac{1}{4\pi} f'_{\rm E}(\chi) = \frac{M_b \left(r_b^2 + \chi^2\right) + \left(9\chi^4 r_b - 2\chi^2 r_b^3 + r_b^5\right) l^{-2} - 2\kappa r_b \chi^2}{2\pi \left(\chi^2 - r_b^2\right)^2}$$
$$= \frac{1}{8\pi\chi}$$

Hawking temperature

$$T = \frac{1}{4\pi} f'_{\rm E}(\chi) = \frac{M_b \left(r_b^2 + \chi^2\right) + \left(9\chi^4 r_b - 2\chi^2 r_b^3 + r_b^5\right) l^{-2} - 2\kappa r_b \chi^2}{2\pi \left(\chi^2 - r_b^2\right)^2}$$
$$= \frac{1}{8\pi\chi}$$

$$\rightarrow r_b = \frac{1}{12\chi} \left(l^2 \pm \sqrt{l^4 + 144\chi^4 - 48l^2\chi^2\kappa} \right), \quad \chi = \chi_{\max} = \frac{l}{2}\sqrt{\frac{1}{3}(2-\sqrt{3})}$$

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

Hawking temperature





Entropy and Energy



▲□▶▲□▶▲□▶▲□▶ ▲□ → ④へ⊙

First law and free energy

 $dE = TdS, \qquad F = E - TS$



First law and free energy

$$dE = TdS, \qquad F = E - TS$$

Heat capacity

$$C = T \frac{\partial S}{\partial T} = T \left(\frac{\partial T}{\partial M} \frac{\partial M}{\partial r_b} \frac{\partial r_b}{\partial S} \right)^{-1} = \frac{\pi \omega}{G} \frac{r_b (12\chi^3 - l^2 r_b)}{(-12r_b\chi + l^2)}$$

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

First law and free energy

$$dE = TdS, \qquad F = E - TS$$

Heat capacity



▲□▶▲□▶▲□▶▲□▶ = めへで

Thermodynamics of dyonic RN black hole

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ □ のへぐ

$$I_{\rm ren} = \int_{\mathcal{M}} \sqrt{-g} \left(\frac{1}{\kappa^2} R - \frac{1}{4} F^2 \right) + \frac{2}{\kappa^2} \int_{\partial \mathcal{M}} \sqrt{-h} (K - \hat{K}), \tag{5}$$

$$ds^{2} = -f(r)dt^{2} + \frac{dr^{2}}{f(r)} + r^{2}d\Omega_{2}^{2}, \qquad f(r_{h}) = 0,$$
(6)

$$f(r) = 1 - \frac{2M}{r} + \frac{q^2 + p^2}{4r^2}, \qquad F = \frac{1}{\kappa} \left(\frac{q}{r^2} dt \wedge dr + p\Omega_2\right).$$
(7)

$$I_{\rm E,ren} = -\int_{\mathcal{M}} \sqrt{g} \left(\frac{1}{\kappa^2} R - \frac{1}{4} F^2 \right) - \frac{2}{\kappa^2} \int_{\partial \mathcal{M}} \sqrt{h} (K - \hat{K})$$
(5)

$$ds_{\rm E}^2 = f(r)d\tau^2 + \frac{dr^2}{f(r)} + r^2 d\Omega_2^2, \qquad f(r_h) = 0,$$
(6)

$$f(r) = 1 - \frac{2M}{r} + \frac{q^2 + p^2}{4r^2}, \qquad F = \frac{1}{\kappa} \left(\frac{q}{r^2} dt \wedge dr + p\Omega_2\right).$$
(7)

$$I_{\rm E,ren} = -\int_{\mathcal{M}} \sqrt{g} \left(\frac{1}{\kappa^2} R - \frac{1}{4} F^2 \right) - \frac{2}{\kappa^2} \int_{\partial \mathcal{M}} \sqrt{h} (K - \hat{K})$$
(5)

$$ds_{\rm E}^2 = f(r)d\tau^2 + \frac{dr^2}{f(r)} + r^2 d\Omega_2^2, \qquad f(r_h) = 0,$$
(6)

$$f(r) = 1 - \frac{2M}{r} + \frac{q^2 + p^2}{4r^2}, \qquad F = \frac{1}{\kappa} \left(\frac{q}{r^2} dt \wedge dr + p\Omega_2\right).$$
(7)



・ロ> < 団> < 三> < 三> < 三・< 三・< へへ

$$I_{\rm E,ren} = -\int_{\mathcal{M}} \sqrt{g} \left(\frac{1}{\kappa^2} R - \frac{1}{4} F^2 \right) - \frac{2}{\kappa^2} \int_{\partial \mathcal{M}} \sqrt{h} (K - \hat{K})$$
(5)

$$ds_{\rm E}^2 = f(r)d\tau^2 + \frac{dr^2}{f(r)} + r^2 d\Omega_2^2, \qquad f(r_h) = 0,$$
(6)

$$f(r) = 1 - \frac{2M}{r} + \frac{q^2 + p^2}{4r^2}, \qquad F = \frac{1}{\kappa} \left(\frac{q}{r^2} dt \wedge dr + p\Omega_2\right).$$
(7)



▶ Regularity Condition on ds_E^2 : Time periodicity → Hawking Temperature

$$f(r)d\tau^2 + \frac{dr^2}{f(r)}\Big|_{r\sim r_h + \epsilon} \sim \rho^2 d\psi^2 + d\rho^2 \tag{8}$$

$$I_{\rm E,ren} = -\int_{\mathcal{M}} \sqrt{g} \left(\frac{1}{\kappa^2} R - \frac{1}{4} F^2 \right) - \frac{2}{\kappa^2} \int_{\partial \mathcal{M}} \sqrt{h} (K - \hat{K})$$
(5)

$$ds_{\rm E}^2 = f(r)d\tau^2 + \frac{dr^2}{f(r)} + r^2 d\Omega_2^2, \qquad f(r_h) = 0,$$
(6)

$$f(r) = 1 - \frac{2M}{r} + \frac{q^2 + p^2}{4r^2}, \qquad F = \frac{1}{\kappa} \left(\frac{q}{r^2} dt \wedge dr + p\Omega_2\right).$$
(7)



Regularity Condition on ds_E^2 : Time periodicity \rightarrow Hawking Temperature

$$f(r)d\tau^2 + \frac{dr^2}{f(r)}\Big|_{r\sim r_h + \epsilon} \sim \rho^2 d\psi^2 + d\rho^2 \tag{8}$$

• Regularity Condition on A (F = dA): A should be regular at the horizon $A_t(r_h) = 0 \rightarrow A_t(r) = \frac{1}{\kappa} \left(\frac{q}{r} - \frac{q}{r_h} \right), \qquad A^2(r_h) = \frac{p^2}{\kappa^2 r_h^2} \cot^2(\theta) \quad (9)$

• Electric potential $\Phi_E^{(1)}$: conjugate variable of Q

$$I_{\rm E, ren} = \beta F, \qquad F = E - TS - \Phi_E^{(1)}Q$$
 (10)

$$\frac{\partial I_{\rm ren}}{\partial \beta} = E - \Phi_E^{(1)} Q, \qquad \Phi_E^{(1)} = \frac{\partial E}{\partial Q} \Big|_{r_b} = \frac{1}{\kappa} \frac{q}{r_h}$$
(11)

• Electric potential $\Phi_E^{(2)}$

$$\Phi_E^{(2)} = A_t(r_h) - A_t(\infty) = \frac{1}{\kappa} \frac{q}{r_h}$$
(12)

•
$$\Phi_E^{(1)}$$
 and $\Phi_E^{(2)}$ are agreed : $\Phi_E^{(1)} = \Phi_E^{(2)} = \Phi_E$
• First law

$$dE = TdS + \Phi_e dQ + \Phi_m dP \tag{13}$$

where

$$Q_m = \int *F, \qquad Q_m = \int F \tag{14}$$

▲□▶ ▲□▶ ▲ 三▶ ▲ 三 ● ● ●

Thermodynamics of charged TB/TN-AdS space

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ □ のへぐ

Charged TB/TN-AdS spacetime

$$I_E = -\frac{1}{16\pi G_4} \int d^4 x \sqrt{g} \left(R + \frac{6}{l^2} - F^{\mu\nu} F_{\mu\nu} \right) - \frac{1}{8\pi G_4} \int d^3 x \sqrt{h} K$$
(15)
$$I_{\rm ren.} = I_E + \frac{1}{8\pi G_4} \int d^3 x \sqrt{h} \left(\frac{2}{l} + \frac{l}{2} R_3 \right),$$
(16)

$$ds_{\rm E}^2 = f_E(r)(d\tau + 2\chi\lambda(\theta)d\phi)^2 + \frac{dr^2}{f_E(r)} + (r^2 - \chi^2)(d\theta^2 + Y(\theta)^2 d\phi^2), \quad (17)$$

$$f_E = \frac{l^{-2}(r^2 - \chi^2)^2 + (\kappa - 4l^{-2}\chi^2)(r^2 + \chi^2) - 2Mr + P^2 + Q^2}{r^2 - \chi^2},$$
 (18)

$$A_{\rm E} = \frac{1}{2\chi} h_E(r) \left(d\tau + 2\chi \lambda(\theta) d\phi \right), \qquad h_{\rm E} = \frac{2iQ\chi r - P(r^2 + \chi^2)}{r^2 - \chi^2}$$
(19)

- ► TB-AdS solution when $r = r_b \neq \chi$, $f_E(r_b) = 0$
- ► TN-AdS solution when $r = r_n = \chi$, $f_E(r_n) = 0$

Electric and magnetic charge

Electric charge

$$Q_e[\xi_t] \equiv \frac{1}{\omega} \int_{\partial \Sigma_\kappa} *F = \lim_{r \to \infty} \frac{Q\left(r^2 - s^2\right) - 2Prs}{r^2 + s^2} = Q, \qquad (20)$$

Magnetic charge

$$Q_m[\xi_\phi] \equiv \frac{1}{\omega} \int_{\partial \Sigma_\kappa} F = \lim_{r \to \infty} \frac{P\left(r^2 - s^2\right) + 2Qrs}{r^2 + s^2} = P$$
(21)

< ロ > < 団 > < 国 > < 国 > < 国 > < 国 > < 国 > < < 回 > < < つ <

Regularity conditions on dyonic TB-AdS

• Regularity condition on
$$ds_E^2$$
 at $r = r_b$

$$T = \frac{1}{4\pi} f'_E(r_b) = \frac{l^2(\kappa r_b - M_b) + 2r_b \left(r_b^2 - 3\chi^2\right)}{2\pi l^2 \left(r_b^2 - \chi^2\right)}$$
(22)

• Regularity condition on A at $r = r_b$

$$A_{\rm E} = \frac{1}{2\chi} h_E(r) \left(d\tau + 2\chi\lambda(\theta)d\phi \right), \quad h_{\rm E} = \frac{2iQ\chi r - P(r^2 + \chi^2)}{r^2 - \chi^2}, \quad (23)$$
$$A_{\rm E,\tau}(r_b) = 0 \quad \rightarrow \quad h_E(r_b) = 0 \quad \rightarrow \quad P = \frac{2iQ\chi r_b}{r_b^2 + \chi^2} \quad (24)$$
Electric and magnetic Potential of dyonic TB-AdS

Electric $\Phi_E^{(1)}$ and magnetic $\Phi_M^{(1)}$ potential : conjugate variable of Q and P

$$I_{\rm E, ren} = \beta F, \qquad F = E - TS - \Phi_E^{(1)}Q - \Phi_M^{(1)}P$$
 (25)

$$\frac{\partial I_{\rm ren}}{\partial \beta} = E - \Phi_E^{(1)} Q - \Phi_M^{(1)} P, \tag{26}$$

$$\Phi_E^{(1)} = \frac{\partial E}{\partial Q}\Big|_{r_b, P}, \qquad \Phi_M^{(1)} = \frac{\partial E}{\partial P}\Big|_{r_b, Q}$$
(27)

Electric
$$\Phi_E^{(2)}$$
 and magnetic $\Phi_M^{(2)}$ potential

$$\Phi_E^{(2)} = A_\tau(r_h) - A_\tau(\infty), \qquad \Phi_M^{(2)} = \Phi_M^{(2)}(r_h) - \Phi_M^{(2)}(\infty)$$
(28)

where

$$\Phi_M^{(2)}(r) = \int^r dr' B(r'), \qquad B(r) = \frac{1}{\sqrt{g}} \epsilon^{tr\theta\phi} F_{\theta\phi}$$
(29)

• $\Phi_E^{(1)} \neq \Phi_E^{(2)}$ and $\Phi_M^{(1)} \neq \Phi_M^{(2)}$

• Once imposing the regularity condition, $\Phi_E^{(1)} = \Phi_E^{(2)}$

Thermodynamics of dyonic TB-AdS ($\kappa \neq 1$)

temperature and entropy

$$T = \frac{1}{4\pi} f'_E(r_b) = \frac{l^2(\kappa r_b - M_b) + 2r_b \left(r_b^2 - 3\chi^2\right)}{2\pi l^2 \left(r_b^2 - \chi^2\right)},$$
$$S \equiv \left(\beta \frac{\partial}{\partial \beta} - 1\right) I_E = \frac{\pi \omega \left(r_b^2 - \chi^2\right)}{2G},$$

by using thermodynamic relations

$$\Phi_{\rm E}^{(1)} = \frac{\omega}{4\pi G} \frac{Q\left(r_b^2 + \chi^2\right) r_b + 2iPr_b^2\chi}{\left(r_b^2 - \chi^2\right)^2}, \qquad \Phi_{\rm M}^{(1)} = \frac{\omega}{4\pi G} \frac{2iQr_b^2\chi - P\left(r_b^2 + \chi^2\right) r_b}{\left(r_b^2 - \chi^2\right)^2}$$
$$E^{(1)} = \frac{r_b\omega}{8\pi G} \left(\kappa + \frac{r_b^2 - 3\chi^2}{l^2} + \frac{\left(Q^2 - P^2\right)\left(r_b^2 + \chi^2\right) + 4iPQr_b\chi}{\left(r_b^2 - \chi^2\right)^2}\right),$$

by using conventional method

$$\Phi_{\rm E}^{(2)} = \frac{\omega}{4\pi G} \frac{Qr_b + iP\chi}{(r_b^2 - \chi^2)}, \qquad \Phi_{\rm M}^{(2)} = \frac{\omega}{4\pi G} \frac{iQ\chi - Pr_b}{(r_b^2 - \chi^2)}.$$
$$E^{(2)} = \frac{\omega}{8\pi G} \left(\kappa r_b + \frac{r_b^3 - 3r_b\chi^2}{l^2} + \frac{(P^2r_b - 2iPQ\chi)\left(r_b^2 + \chi^2\right) + Q^2r_b\left(r_b^2 - 3\chi^2\right)}{(r_b^2 - \chi^2)^2}\right)$$

Thermodynamics of dyonic TB-AdS ($\kappa \neq 1$)

imposing the regularity condition

$$E^{(1)} = E^{(2)} = E = \frac{\omega}{8\pi G} \left(\kappa r_b + \frac{r_b \left(r_b^2 - 3\chi^2 \right)}{l^2} + \frac{Q^2 r_b}{r_b^2 + \chi^2} \right), \quad (30)$$

$$\Phi_{\rm E}^{(1)} = \Phi_{\rm E}^{(2)} = \Phi_{\rm E} = \frac{\omega}{4\pi G} \frac{Qr_b}{\left(r_b^2 + \chi^2\right)},\tag{31}$$

$$\Phi_{\rm M}^{(1)} = 0 \neq \Phi_{\rm M}^{(2)} = -\frac{\omega}{4\pi G} \frac{iQ\chi}{\left(r_b^2 + \chi^2\right)}$$
(32)

thermodynamic relations are satisfied

$$dE = TdS + \Phi_{\rm E}dQ,\tag{33}$$

$$F = E - TS - \Phi_{\rm E}Q \tag{34}$$

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ ● □ ● ● ● ●

upon using $\Phi_{M}^{(1)} = \Phi_{M} = 0.$

Thermodynamics of dyonic TB-AdS ($\kappa = 1$)

Thermodynamic quantities

$$T = \frac{1}{4\pi} f'_E(r_b) = \frac{1}{8\pi\chi}, \quad S = \frac{\pi\omega}{2G} \left[r_b^2 + \chi^2 - \frac{24\chi^3 r_b}{l^2} + \frac{8Q^2\chi^3 r_b}{\left(r_b^2 + \chi^2\right)^2} \right]$$

by using thermodynamic relations

$$\begin{split} \Phi_{\rm E}^{(1)} &= \frac{\omega}{4\pi G} \frac{r_b \left(2iP\chi r_b + Q(r_b^2 + \chi^2)\right)}{\left(\chi^2 - r_b^2\right)^2}, \qquad \Phi_{\rm M}^{(1)} = -\frac{\omega}{4\pi G} \frac{r_b \left(P(r_b^2 + \chi^2) - 2iQ\chi r_b\right)}{\left(\chi^2 - r_b^2\right)^2}.\\ E^{(1)} &= \frac{\omega}{8\pi G} \left[M_b + \frac{r_b^2 + \chi^2}{4\chi} - \frac{3\chi^2 r_b + r_b^3}{l^2} + \frac{\left(Q^2 - P^2\right)\left(-3\chi^4 r_b - 6\chi^2 r_b^3 + r_b^5\right)}{\left(r_b^2 - \chi^2\right)^3} + \frac{16iPQ\chi^3 r_b^2}{\left(\chi^2 - r_b^2\right)^3}\right], \end{split}$$

by using conventional method

$$\begin{split} \Phi_{\rm E}^{(2)} &= \frac{\omega}{4\pi G} \frac{(Qr_b + iP\chi)}{(r_b^2 - \chi^2)}, \quad \Phi_{\rm M}^{(2)} &= \frac{\omega}{4\pi G} \frac{(-Pr_b + iQ\chi)}{(r_b^2 - \chi^2)} \\ E^{(2)} &= \frac{\omega}{8\pi G} \left[M_b + \frac{r_b^2 + \chi^2}{4\chi} - \frac{r_b \left(r_b^2 + 3\chi^2\right)}{l^2} + \frac{1}{(r_b^2 - \chi^2)^3} \left(P^2 r_b (r_b^4 + 6r_b^2 \chi^2 + \chi^4) + 2iPQ\chi (-3r_b^4 - 6r_b^2 \chi^2 + \chi^4) + Q^2 r_b (r_b^4 - 10r_b^2 \chi^2 + \chi^4) \right) \right]. \end{split}$$

Thermodynamics of dyonic TB-AdS ($\kappa = 1$)

imposing the regularity condition

$$E = \frac{\omega}{8\pi G} \left[M_b + \frac{r_b^2 + \chi^2}{4\chi} - \frac{r_b \left(r_b^2 + 3\chi^2\right)}{l^2} + \frac{Q^2 \left(3\chi^2 r_b + r_b^3\right)}{\left(r_b^2 + \chi^2\right)^2} \right], \quad (35)$$

$$S = \frac{\omega}{24\chi^3 r_b} \left[\frac{8Q^2 \chi^3 r_b}{r_b} \right] = \frac{8Q^2 \chi^3 r_b}{r_b}$$

$$S = \frac{\alpha}{4G} \left[r_b^2 + \chi^2 - \frac{24\chi + b}{l^2} + \frac{34\chi + b}{\left(r_b^2 + \chi^2\right)^2} \right],$$
(36)

$$\Phi_{\rm E} = \frac{\omega}{4\pi G} \frac{Qr_b}{\left(r_b^2 + \chi^2\right)} \tag{37}$$

thermodynamic relations are satisfied

$$dE = TdS + \Phi_{\rm E}dQ \tag{38}$$

$$F = E - TS - \Phi_{\rm E}Q \tag{39}$$

• Heat Capacity ($P = -2i\chi g$)

$$C_{P,Q} \equiv -\beta \frac{\partial S}{\partial \beta} = \frac{\omega l^2 r_b \chi^2}{2Ga_1} \left[\frac{\left(r_b^2 - \chi^2\right)^2 \left(l^2 r_b - 12\chi^3\right)}{l^2 \chi^2} + \frac{4\chi \left(Q^2 - P^2\right) \left(9r_b^4 + 14r_b^2 \chi^2 + \chi^4\right) + 16iPQr_b \left(r_b^4 + 8r_b^2 \chi^2 + 3\chi^4\right)}{\left(r_b^2 - \chi^2\right)^2} \right]$$

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ ● □ ● ● ● ●

Thermodynamics of charged TB-AdS ($\kappa = 1$)



Figure: The horizon radius of TB-AdS for $\kappa = 1$ vs χ for fixed values of Q and g with G = l = 1.

▲□▶ ▲□▶ ▲ □▶ ▲ □▶ ▲ □ ● のへで

Thermodynamics of charged TB-AdS ($\kappa = 1$)



Figure: Entropy of TB-AdS for $\kappa = 1$ vs χ for fixed values of Q and g with G = l = 1

Thermodynamics of charged TB-AdS ($\kappa = 1$)



Figure: Heat capacity for TB-AdS for $\kappa = 1$ vs χ for fixed Q and g with G = l = 1

Thermodynamics of charged TN-AdS ($\kappa = 1$)

► Taub-"NUT"-AdS solution when $r_+ = \chi$ where $f_E(r_+) = 0$

$$f_n(r) = \frac{l^{-2}(r-\chi)^2(r+3\chi)\chi + \kappa(r-\chi)\chi - (P^2 + Q^2)}{\chi(r+\chi)}, \qquad (40)$$

$$l^{-2}(r_{+} - \chi)^{2}(r_{+} + 3\chi)\chi + \kappa(r_{+} - \chi)\chi - P^{2} - Q^{2} = 0$$
(41)
 $\rightarrow P^{2} + Q^{2} = 0 \quad (P = iQ) \text{ is required}$

Hawking temperature

$$T_n = \frac{1}{4\pi} f'_n(\chi) = \frac{\kappa}{8\pi\chi}$$

• entropy, energy, and gauge field potential ($P = 2iv\chi \rightarrow Q = 2v\chi$)

$$S = \left(\beta \frac{\partial}{\partial \beta} - 1\right) I_{\text{ren.}} = \frac{2\pi\omega\chi^2}{G} \left(1 + 2v^2 - \frac{6\chi^2}{l^2}\right),$$
$$E = \partial_\beta I_{\text{ren.}} = \frac{\omega\chi}{2G} \left(1 + 2v^2 - \frac{4\chi^2}{l^2}\right),$$
$$\Phi = \frac{i\omega}{2G} \frac{1}{2\chi} \left(h_{\text{E}}(\chi) - h_{\text{E}}(\infty)\right) = \frac{\omega}{2G}v$$

First law and free energy

$$dE = T_n dS + \Phi dQ, \qquad F = E - T_n S - \Phi Q.$$

Thermodynamics of charged TN-AdS ($\kappa \neq 1$)

For $\kappa=0$ case,

- zero Hawking temperature
- energy and gauge potential

$$E = \frac{\omega}{2G} \left(\frac{Q^2}{4\chi} - \frac{\chi^3}{l^2} \right), \qquad \Phi = \frac{Q\omega}{4G\chi}$$
(42)

First law and free energy

$$dE = \Phi dQ, \qquad F = E - \Phi Q. \tag{43}$$

For $\kappa = -1$ case, $f_n(r)$ becomes negative near $r_+ = \chi$ and so we exclude this case.

< ロ > < 団 > < 三 > < 三 > < 三 > < 三 < の < (~)

Near Horizon Geometry of the extremal TB-AdS

The extremal limit

 $T_{\rm H} \rightarrow 0.$

At the zero temperature limit, the metric is factorized as follows

$$f_{\rm E} = \frac{1}{r^2 l^2} (r + r_0 + \alpha) (r + r_0 - \alpha) (r - r_0)^2$$

where

$$r_{0} = \sqrt{\chi^{2} + \frac{1}{6}l\left(\sqrt{\kappa^{2}l^{2} + 12\left(Q^{2} + P^{2}\right)} - \kappa l\right)},$$

$$\alpha = \sqrt{4\chi^{2} - \frac{1}{3}l\left(\sqrt{\kappa^{2}l^{2} + 12\left(Q^{2} + P^{2}\right)} + 2\kappa l\right)}.$$

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

Near Horizon Geometry of the extremal TB-AdS

The extremal limit

 $T_{\rm H} \rightarrow 0.$

At the zero temperature limit, the metric is factorized as follows

$$f_{\rm E} = \frac{1}{r^2 l^2} (r + r_0 + \alpha) (r + r_0 - \alpha) (r - r_0)^2$$

where

$$r_{0} = \sqrt{\chi^{2} + \frac{1}{6}l\left(\sqrt{\kappa^{2}l^{2} + 12\left(Q^{2} + P^{2}\right)} - \kappa l\right)},$$

$$\alpha = \sqrt{4\chi^{2} - \frac{1}{3}l\left(\sqrt{\kappa^{2}l^{2} + 12\left(Q^{2} + P^{2}\right)} + 2\kappa l\right)}.$$

Taking the near horizon expansion,

$$\tau \to \frac{\tilde{\tau}}{\epsilon}, \qquad r \to r_0 + \epsilon$$

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

Near Horizon Geometry of the extremal TB-AdS

The extremal limit

 $T_{\rm H} \rightarrow 0.$

At the zero temperature limit, the metric is factorized as follows

$$f_{\rm E} = \frac{1}{r^2 l^2} (r + r_0 + \alpha) (r + r_0 - \alpha) (r - r_0)^2$$

where

$$r_{0} = \sqrt{\chi^{2} + \frac{1}{6}l\left(\sqrt{\kappa^{2}l^{2} + 12\left(Q^{2} + P^{2}\right)} - \kappa l\right)},$$

$$\alpha = \sqrt{4\chi^{2} - \frac{1}{3}l\left(\sqrt{\kappa^{2}l^{2} + 12\left(Q^{2} + P^{2}\right)} + 2\kappa l\right)}.$$

Taking the near horizon expansion,

$$\tau \to \frac{\tilde{\tau}}{\epsilon}, \qquad r \to r_0 + \epsilon$$

$$ds_{\text{ext}}^{2} \sim \frac{\tilde{r}^{2}}{l^{2}} \frac{(4r_{0}^{2} - \alpha^{2})}{r_{0}^{2} - \chi^{2}} d\tilde{\tau}^{2} + \frac{l^{2}}{\tilde{r}^{2}} \frac{r_{0}^{2} - \chi^{2}}{(4r_{0}^{2} - \alpha^{2})} d\tilde{r}^{2} + (r_{0}^{2} - \chi^{2})(d\theta^{2} + Y(\theta)^{2}d\phi^{2}),$$

$$\sim AdS_{2} \times S^{2} (H^{2} \text{ or } R^{2})$$

▲□▶▲□▶▲□▶▲□▶ ▲□ ● ④ ● ●

- $\blacktriangleright \kappa \neq 1$ case : We can take the zero temperature limit.
- $\triangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

- $\kappa \neq 1$ case : We can take the zero temperature limit.
- $\blacktriangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

where

1. physical value of r_b is only defined by finite ranges of χ

- $\blacktriangleright \kappa \neq 1$ case : We can take the zero temperature limit.
- $\blacktriangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

▲□▶▲□▶▲□▶▲□▶ ■ のへで

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

- $\kappa \neq 1$ case : We can take the zero temperature limit.
- $\blacktriangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

- $\kappa \neq 1$ case : We can take the zero temperature limit.
- $\triangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

<□▶ < □▶ < □▶ < □▶ < □▶ < □▶ < □ < ○ < ○

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

[Conclusion1.] No zero temperature limit for $\kappa = 1$ case

3. Time periodicity is not defined on AdS_2

- $\blacktriangleright \kappa \neq 1$ case : We can take the zero temperature limit.
- $\triangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

<□▶ < □▶ < □▶ < □▶ < □▶ < □▶ < □ < ○ < ○

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

- 3. Time periodicity is not defined on AdS_2
- 4. Intrinsically we cannot give time periodicity

- $\kappa \neq 1$ case : We can take the zero temperature limit.
- $\blacktriangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

<□▶ < □▶ < □▶ < □▶ < □▶ < □▶ < □ < ○ < ○

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

- 3. Time periodicity is not defined on AdS_2
- 4. Intrinsically we cannot give time periodicity
- 4. No Misner string at the horizon

- $\blacktriangleright \kappa \neq 1$ case : We can take the zero temperature limit.
- $\blacktriangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

<□▶ < □▶ < □▶ < □▶ < □▶ < □▶ < □ < ○ < ○

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

- 3. Time periodicity is not defined on AdS_2
- 4. Intrinsically we cannot give time periodicity
- 4. No Misner string at the horizon
- 5. Then we can give up $\Delta \tau = 4\chi$

- $\kappa \neq 1$ case : We can take the zero temperature limit.
- $\triangleright \kappa = 1$ case : We imposed the following condition at finite temperature

$$T_H = \frac{1}{2\pi} f'(r_b) = \frac{1}{8\pi\chi}$$
(44)

where

- 1. physical value of r_b is only defined by finite ranges of χ
- 2. but zero temperature limit requires $\chi \to \infty$

[Conclusion1.] No zero temperature limit for $\kappa = 1$ case

- 3. Time periodicity is not defined on AdS_2
- 4. Intrinsically we cannot give time periodicity
- 4. No Misner string at the horizon
- 5. Then we can give up $\Delta \tau = 4\chi$

[Conclusion2.] zero temperature limit for $\kappa = 1$ case without $\Delta \tau = 4\chi$

The mass parameter and charges are related as

$$M = \frac{\left(2\kappa l^2 - 12\chi^2 + l\Delta\right)\sqrt{6\chi^2 - l^2\kappa + l\Delta}}{3\sqrt{6}l^2},$$
$$\Delta = \sqrt{\kappa^2 l^2 + 12(Q^2 + P^2)}$$

The free energy and energy are

$$F_{ext} = -\frac{\omega r_b}{2G} \left(\frac{\kappa \chi^2}{r_b^2 - \chi^2} + \frac{r_b^2 + 3\chi^2}{l^2} \right),$$
$$E_{ext} = \frac{\omega r_b^3}{2G} \left(\frac{\kappa}{r_b^2 - \chi^2} + \frac{2}{l^2} \right)$$

The first law and free energy are satisfied

$$F_{ext} = E_{ext} - \Phi_{\rm E}Q|_{ext},$$
$$dE_{ext} = \Phi_{\rm E}dQ|_{ext}$$

▲□▶ ▲□▶ ▲ 三▶ ▲ 三▶ ● 三 ● のへで

Inspired by the idea¹ that

- $\blacktriangleright \ \omega \equiv \frac{1}{T}$ where ω is the conjugate variable to the charges Q+J
- BPS black holes behaves similar to the AdS-Schwarzschild black holes

¹Sunjin Choi, Joonho Kim, Seok Kim and June Nahmgoong, "Comments on deconfinement in AdS/CFT", arXiv:1811.08646

Inspired by the idea¹ that

- $\omega \equiv \frac{1}{T}$ where ω is the conjugate variable to the charges Q + J
- ▶ BPS black holes behaves similar to the AdS-Schwarzschild black holes Here we identify Φ_E as "temperature".

$$\Phi_{\rm E} = \frac{\omega}{2G} \frac{Q_{ext} r_b}{(r_b^2 + \chi^2)}, \qquad Q_{ext}^2 = \left(r_b^2 + \chi^2\right)^2 \left(\frac{\kappa}{r_b^2 - \chi^2} + \frac{3}{l^2}\right)$$

¹Sunjin Choi, Joonho Kim, Seok Kim and June Nahmgoong, "Comments on deconfinement in AdS/CFT", arXiv:1811.08646

Inspired by the idea¹ that

• $\omega \equiv \frac{1}{T}$ where ω is the conjugate variable to the charges Q + J

▶ BPS black holes behaves similar to the AdS-Schwarzschild black holes Here we identify Φ_E as "temperature".

$$\Phi_{\rm E} = \frac{\omega}{2G} \frac{Q_{ext} r_b}{(r_b^2 + \chi^2)}, \qquad Q_{ext}^2 = \left(r_b^2 + \chi^2\right)^2 \left(\frac{\kappa}{r_b^2 - \chi^2} + \frac{3}{l^2}\right)$$

 \blacktriangleright horizon radius r_b vs inverse "temperature"



¹Sunjin Choi, Joonho Kim, Seok Kim and June Nahmgoong, "Comments on deconfinement in AdS/CFT", arXiv:1811.08646

Inspired by the idea¹ that

 $\blacktriangleright \omega \equiv \frac{1}{T}$ where ω is the conjugate variable to the charges Q + J

BPS black holes behaves similar to the AdS-Schwarzschild black holes Here we identify $\Phi_{\rm E}$ as "temperature".

$$\Phi_{\rm E} = \frac{\omega}{2G} \frac{Q_{ext} r_b}{(r_b^2 + \chi^2)}, \qquad Q_{ext}^2 = \left(r_b^2 + \chi^2\right)^2 \left(\frac{\kappa}{r_b^2 - \chi^2} + \frac{3}{l^2}\right)$$

- horizon radius r_b vs inverse "temperature"
- entropy S vs inverse "temperature"



deconfinement in AdS/CFT", arXiv:1811.08646 ◆□▶ ◆□▶ ◆ 三▶ ◆ 三▶ ◆ □ ◆ ○ ◇ ◇ ◇

Now let us define the heat capacity as follows

$$C \equiv T \frac{\partial S}{\partial T} \rightarrow C \equiv \Phi_{\rm E} \frac{\partial S}{\partial \Phi_{\rm E}}$$

then

$$C = \Phi_{\rm E} \frac{\partial S}{\partial \Phi_{\rm E}} = \frac{\pi \omega r_b^2 \left(r_b^2 - \chi^2 \right) \left(3r_b^2 + \kappa l^2 - 3\chi^2 \right)}{G \left(-6\chi^2 r_b^2 + 3r_b^4 - \kappa l^2 \chi^2 + 3\chi^4 \right)}$$

◆□ ▶ ◆□ ▶ ◆ □ ▶ ◆ □ ▶ ● □ ● ● ● ●

Now let us define the heat capacity as follows

$$C \equiv T \frac{\partial S}{\partial T} \rightarrow C \equiv \Phi_{\rm E} \frac{\partial S}{\partial \Phi_{\rm E}}$$

then

$$C = \Phi_{\rm E} \frac{\partial S}{\partial \Phi_{\rm E}} = \frac{\pi \omega r_b^2 \left(r_b^2 - \chi^2\right) \left(3r_b^2 + \kappa l^2 - 3\chi^2\right)}{G \left(-6\chi^2 r_b^2 + 3r_b^4 - \kappa l^2\chi^2 + 3\chi^4\right)}$$

$$C$$

$$\frac{300}{100}$$

$$\frac{0.1 \quad 0.2 \quad 0.3 \quad 0.4 \quad 0.5 \quad 0.6 \quad \Phi_E^{-1}$$

$$-100$$

$$-200$$

< ロ > < 団 > < 巨 > < 巨 > < 巨 < つ < つ

Summary and Future works



▲□▶▲□▶▲≡▶▲≡▶ ■ めへぐ

For all κ, thermodynamics are tested for non-charged/charged TB/TN-AdS space.

For all κ, thermodynamics are tested for non-charged/charged TB/TN-AdS space.

◆□ ▶ ◆□ ▶ ◆ □ ▶ ◆ □ ▶ ● □ ● ● ● ●

The regularity condition which comes from $A_t(r_+) = 0$ is necessary to satisfy the first law and free energy.

- For all κ, thermodynamics are tested for non-charged/charged TB/TN-AdS space.
- The regularity condition which comes from $A_t(r_+) = 0$ is necessary to satisfy the first law and free energy.
- For the extremal case, the thermodynamic like behaviours are observed by identifying the gauge potential to temperature.

- For all κ, thermodynamics are tested for non-charged/charged TB/TN-AdS space.
- The regularity condition which comes from $A_t(r_+) = 0$ is necessary to satisfy the first law and free energy.
- For the extremal case, the thermodynamic like behaviours are observed by identifying the gauge potential to temperature.
- (Recently, there have been studied about thermodynamics of Lorentzian Taub-NUT-spacetimes.)
▲□▶▲□▶▲≡▶▲≡▶ ■ 少々⊙

• More investigation on extremal limit for $\kappa = 1$

- More investigation on extremal limit for $\kappa = 1$
- (C2) understanding physical mechanism that the entropy changes from finite temperature to zero temperature for κ = 1 case.

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

- More investigation on extremal limit for $\kappa = 1$
- (C2) understanding physical mechanism that the entropy changes from finite temperature to zero temperature for κ = 1 case.

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

(C2) resolution for Misner string at extremal case $\kappa = 1$

- More investigation on extremal limit for $\kappa = 1$
- (C2) understanding physical mechanism that the entropy changes from finite temperature to zero temperature for $\kappa = 1$ case.
- (C2) resolution for Misner string at extremal case $\kappa = 1$
- Holographic aspect. If the boundary field theory calculation can match the entropy of BPS Taub-NUT-AdS spacetime (in progress)

▲□▶▲□▶▲□▶▲□▶ □ のQ@

- More investigation on extremal limit for $\kappa = 1$
- (C2) understanding physical mechanism that the entropy changes from finite temperature to zero temperature for κ = 1 case.
- (C2) resolution for Misner string at extremal case $\kappa = 1$
- Holographic aspect. If the boundary field theory calculation can match the entropy of BPS Taub-NUT-AdS spacetime (in progress)
- Thermodynamics for rotating-Taub-NUT-AdS spacetime.

- More investigation on extremal limit for $\kappa = 1$
- (C2) understanding physical mechanism that the entropy changes from finite temperature to zero temperature for κ = 1 case.
- (C2) resolution for Misner string at extremal case $\kappa = 1$
- Holographic aspect. If the boundary field theory calculation can match the entropy of BPS Taub-NUT-AdS spacetime (in progress)

▲□▶▲□▶▲□▶▲□▶ □ のQ@

- Thermodynamics for rotating-Taub-NUT-AdS spacetime.
- Holographic aspect of rotating-Taub-NUT-AdS spacetime.