# Baryogenesis & Dark Matter in Axionic Higgs Portal

Recent and on-going works in collaboration with

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- I. Introduction
- II. Axionic Higgs portal
  - Axion
  - Dark matter
  - Baryogenesis
- III. Summary

# What is physics beyond the SM?

Before LHC

Models to explain the EW hierarchy problem

H-----H 
$$\delta m_H^2 \sim \frac{(\text{UV cutoff scale})^2}{16 \, \pi^2}$$

 $\rightarrow$  TeV particles with sizable couplings to the SM

e.g. SUSY, extra dimensions, composite Higgs, ...

 $\rightarrow$  (often) WIMP as cold dark matter



#### I. Introduction- BSM

#### LHC results so far

- No significant deviations from the SM
- No clear signals for BSM

#### Direct & indirect dark matter searches so far

• No evidence of WIMPs





#### I. Introduction- BSM

#### What is physics beyond the SM?

After LHC

Another viable direction

- New light particles feebly coupled to the SM



- Cosmologically important
- Solution to the SM problems?

#### I. Introduction- BSM

## What is physics beyond the SM?

Portal frameworks

Interaction between the SM and dark sector

- Dark Higgs:  $\lambda S^2 H^{\dagger} H$
- Sterile neutrino: yLHN
- Dark photon:  $\epsilon F'_{\mu\nu}B^{\mu\nu}$
- Axion:  $\frac{a}{f}\tilde{F}_{\mu\nu}F^{\mu\nu}$

## Axion/Axion-like-particle

Pseudo Nambu-Goldstone boson of spontaneously broken U(1)

Periodic:  $\phi \equiv \phi + 2\pi f$ 

where axion decay constant f = U(1) breaking scale

Mass and couplings

- Interaction: suppressed by f
- Mass: controlled by shift symmetry,  $\phi \rightarrow \phi + \text{constant}$

$$m_{\phi} = \frac{(\text{shift symmetry breaking scale})^2}{f}$$

 $\rightarrow$  Feebly interacting light particle for large f

# **Axion portal**

- Perturbative shift symmetry,  $\phi \rightarrow \phi + \text{constant}$ 
  - 3 types of interaction
  - Yukawa, derivative, and anomalous coupling

$$m_{\psi}e^{ic_{1}\frac{\phi}{f}}\bar{\psi}\psi + c_{2}\frac{\partial_{\mu}\phi}{f}\bar{\psi}\gamma^{\mu}\gamma_{5}\psi + \frac{c_{3}}{16\pi^{2}}\frac{\phi}{f}F\tilde{F}$$

- Physics: combinations of  $c_i$  invariant under chiral field redefinitions

 Potential to be probed by cosmological, astrophysical and laboratory observations

#### I. Introduction- Axion

#### **Experiments**



## Examples

QCD axion for CP conservation in QCD

Neutron EDM bound on  $\bar{\theta}G\tilde{G}$ 



Peccei, Quinn 1977

Axion anomalously coupled to gluons

- Natural solution to the strong CP problem:  $\bar{\theta} \propto \langle \phi \rangle = 0$
- Dark matter: misalignment, topological defects

#### I. Introduction- Axion

#### Examples

Freese, Frieman, Olinto 1990

Natural inflation

Inflation: initial conditions for the Big Bang cosmology

Very flat potential from axion

$$V = \Lambda^4 \left( 1 - \cos\left(\frac{\phi}{f}\right) \right)$$
 with  $f \ge M_{Pl}$ 



#### II. Axionic Higgs Portal- Framework

#### **Portal interaction**

• Axion portal ( $\phi \equiv \phi + 2\pi f$ )

$$m_{\psi}e^{ic_{1}\frac{\phi}{f}}\bar{\psi}\psi + c_{2}\frac{\partial_{\mu}\phi}{f}\,\bar{\psi}\gamma^{\mu}\gamma_{5}\psi + \frac{c_{3}}{16\pi^{2}}\frac{\phi}{f}F\tilde{F}$$

Another type of axion portal coupling

$$M^2 \cos\left(\frac{\phi}{f}\right) |H|^2$$

- Feeble interaction with the SM via the Higgs field
- Much interest after "Cosmological relaxation of the Higgs mass"

Graham, Kaplan, Rajendran 2015

#### II. Axionic Higgs Portal- Framework

#### **UV completion**

- Model with perturbative shift symmetry
  - Hidden QCD
  - Vector-like lepton doublets  $L + L^c$  and singlets  $N + N^c$



# **Axion properties**

Scalar potential

$$V = \lambda |H|^4 + \mu^2 |H|^2 - M^2 \cos\left(\frac{\phi}{f}\right) |H|^2 - \Lambda^4 \cos\left(\frac{\phi}{f} + \alpha\right)$$

$$\uparrow$$
closed Higgs loop:  $\Lambda^4 \ge \frac{1}{16\pi^2} M^2 \Lambda_{cut}^2$ 

- Axion properties depend on  $f, M, \Lambda, \alpha$
- Axion evolution is tied to EW symmetry breaking!

#### **Experimental constraints**

- Axion-Higgs mixing after EWSB for  $\alpha \neq 0$ 
  - Stringent constraints for axion at sub-MeV to multi-GeV from rare K and B meson decays, and beam-dump experiments
- Further constraints if anomalously couples to SM gauge bosons



**Cosmological roles of axion** 

$$V = \lambda |H|^4 + \mu^2 |H|^2 - \frac{M^2}{f} \cos\left(\frac{\phi}{f}\right) |H|^2 - \frac{\Lambda^4}{f} \cos\left(\frac{\phi}{f} + \alpha\right)$$

#### **Axion-driven EW phase transition**

for *M* above the weak scale

# Cosmologically stable axion

for tiny  $\alpha$  or huge f

#### Freeze-in dark matter

- Alternative to WIMP dark matter (freeze-out)
- Never in thermal equilibrium: feebly coupled to SM
  - Produced via thermal freeze-in McDonald 2001, Choi, Roszkowski 2005, Petraki, Kusenko 2007



Hall, Jedamzik, March-Russell, West 2009

2-2 scattering, decay of thermal particles

• Relic abundance:  $\Omega_{\chi} \propto \lambda^2 m_{\chi}$ 

- Observed DM abundance if  $\lambda \sim 10^{-12}$  and  $m_{\chi} \sim 100 {\rm GeV}$ 

• Need an explanation for  $\lambda \ll 1!$ 

Gravitino, axino in SUSY (many works) Clockwork FIMP, Mohan and Sengupta 2018

#### II. AHP- Model for Freeze-in Axion Dark Matter

#### **Axionic Higgs portal**

**KSJ**, Im 1907.07383

Coupled to the SM ONLY via Higgs portal

- CP conserving minimum  $\phi = 0$ 
  - No ALP-Higgs mixing
  - Stable axion due to accidental Z<sub>2</sub> symmetry  $\phi \rightarrow -\phi$
  - Feebly coupled to SM thermal bath for large f
  - $\rightarrow$  Natural framework for freeze-in DM

## Axion dark matter

Axion properties

- Mass mainly from closed Higgs loops:  $m_{\phi} \simeq \frac{\Lambda^2}{f}$ 

- Interactions with the SM

$$\frac{\lambda_{h\phi}}{4}h^2\phi^2 + \frac{\lambda_{h\phi}\nu}{2}h\phi^2 \text{ with } \lambda_{h\phi} = \left(\frac{M}{f}\right)^2$$

- Never in thermal equilibrium for  $\lambda_{h\phi} < 10^{-7}$ 

 If thermalized, it overcloses the universe in most of parameter space satisfying the bound on DM scattering with nuclei

## Axion dark matter

- Freeze-in production via  $h \rightarrow \phi \phi$  (dominant if open), and  $hh \rightarrow \phi \phi$ 
  - Correct dark matter density

$$\lambda_{h\phi} \simeq 10^{-10} \times \left(\frac{m_{\phi}}{_{3 \text{MeV}}}\right)^{-\frac{1}{2}} \text{ and } m_{\phi} \simeq 1 \text{MeV} \times \left(\frac{\Lambda}{_{10^{3} \text{GeV}}}\right)^{\frac{4}{5}}$$
  
 $\lambda_{h\phi} \simeq 10^{-11} \text{ and } m_{\phi} \simeq 380 \text{GeV} \times \left(\frac{\Lambda}{_{10^{9} \text{GeV}}}\right)$ 

• Coherent oscillations: negligible if  $T_{\rm osc} \gg 10^6 \times m_{\phi}$ 

## **UV completion**

- Non-perturbative Higgs portal via hidden QCD
  - Lepton doublets  $L + L^c$  and singlets  $N + N^c$
  - Generically,  $\alpha \neq 0$
- Severe constraint on the lifetime ( $\alpha$ ) from gamma ray observations





#### II. AHP- Model for Electroweak Baryogenesis

#### **Electroweak phase transition**

Last period affecting baryon asymmetry



B+L violation by rapid EW sphaleron transition in symmetric phase





- Baryogenesis
- 1) Nonzero B-L above EW scale: Leptogenesis, Affleck-Dine, ...
- 2) B+L generation at EW scale and sphaleron decoupling
  - → EW baryogenesis

Lots of works since 1985

#### II. AHP- Model for Electroweak Baryogenesis

#### **EW baryogenesis**



- Requirements
  - Strong first-order phase transition to avoid washout

SM: crossover if  $m_h > 75$  GeV

- Sufficient CP violation beyond SM



#### **Conventional scenarios**

Strong first-order phase transition

e.g. thermal or effective Higgs cubic term, log potential

higher dim operator with low cutoff

- $\rightarrow$  New particles coupled to *H* or sizable modification of Higgs sector
- Non-local baryogenesis
  - CP violation in front of wall, B violation away from wall
- Probe of EWBG
  - LHC (direct searches) and EDM experiments

c.f. severe constraint from electron EDM:  $|d_e| < 1.1 \times 10^{-29} e \cdot cm$ 

ACME II 2018

#### II. AHP- Model for Electroweak Baryogenesis

#### **Axionic Higgs portal**

• Free energy:  $V = V_0 + (\text{thermal effect})$ 

$$V_{0} = \lambda |H|^{4} + \left[ \mu^{2} - M^{2} \cos\left(\frac{\phi}{f}\right) \right] |H|^{2} - \Lambda^{4} \cos\left(\frac{\phi}{f} + \alpha\right)$$
  
$$\Delta V_{\text{thermal}} = c_{H} T^{2} |H|^{2} \quad \text{and} \quad \text{and}$$

Free energy in terms of 3 positive parameters

$$\alpha, \ \epsilon \equiv \frac{\sqrt{\lambda}\Lambda^2}{M^2}, \ r \equiv \frac{\Lambda^2}{\sqrt{\lambda}v_0^2}$$

with  $\lambda$  and  $\mu^2$  fixed by  $m_h = 125 \text{GeV}$  and  $v_0 = 246 \text{GeV}$ 

# **EWPT in axionic Higgs portal**

- For large *f*, tiny thermal and quantum corrections to *V* from axion
- Free energy  $V = V(h, \phi/f)$ : Insensitive to f
- Tunneling mainly along the light axion direction for  $f \gg \text{TeV}$ 
  - Higgs field can be replaced by solving  $\partial_h V = 0$
- Phase transition
  - Two degenerate minima at  $T_c$ : lower than in the SM
  - Bubble nucleation at  $T_n$
  - Barrier disappears at  $T_2$

#### **EWPT in axionic Higgs portal**



# **Approximate scaling behaviors**

• Euclidean action of O(3) symmetric critical bubble

 $S_3 \propto f^3$ 

 $\Rightarrow$  Bubble radius  $\propto f$ , and  $T_n$  close to  $T_2$ 

c.f. conventional models: thin wall,  $T_n$  close to  $T_c$ 

 Relatively smooth phase transition (no substantial expansion of bubbles), but rapid ALP evolution

## **EW phase transition**

• Case with 
$$r \equiv \frac{\Lambda^2}{\sqrt{\lambda}v_0^2} = 1.2$$



- First-order PT in red region (wider at small *r*)
- Red region close to blue line
- → EWPT at very low T
   PT pattern: insensitive to *f*

# Implications

- Delayed EW phase transition
  - Late-time entropy production to dilute preexisting relics (e.g. dark matter)

$$\Delta = 10^4 \left(\frac{T_{\rm reh}}{40 {\rm GeV}}\right)^3 \left(\frac{T_n}{2 {\rm GeV}}\right)^{-3}$$

Temperature after phase transition

$$T_{reh} = 40 \text{GeV}\left(\frac{\Delta V^{1/4}}{100 \text{GeV}}\right)$$

- Strong first order phase transition
  - EW baryogenesis

## **Axionic EW baryogenesis**

Strong first order phase transition driven by axion via

$$M^2 \cos\left(\frac{\phi}{f}\right) |H|^2$$

- New direction in EWBG
  - Free from EDM and LHC constraints for  $f \gg \text{TeV}$
  - Axion searches to reveal the connection between
    - EW phase transition and baryogenesis

## **EW baryogenesis**

- CP violation depending on the bubble wall profile
- 1) Axion-dependent top quark Yukawa

$$(y_t + x_t e^{i\phi/f})ht_L t_R$$

axion variation along the wall  $\rightarrow$  B number chemical potential

2) Anomalous axion coupling to EW gauge bosons

$$\frac{1}{16\pi^2} \frac{\phi}{f} W_{\mu\nu} \widetilde{W}^{\mu\nu}$$

 $\frac{d\phi}{dt} \rightarrow B$  number chemical potential

#### II. AHP- Model for Electroweak Baryogenesis- Baryon Asymmetry

#### **EW baryogenesis**

Condition for EWBG



#### II. AHP- Model for Electroweak Baryogenesis- Baryon Asymmetry

## 1) Non-local EWBG KSJ, Jung, Shin 1806.02591

CP violation from axion-dependent top quark mass



- Baryon asymmetry
  - CP violation  $x_t$ , wall width  $L_w$ , wall velocity  $v_w$
  - Sizable diffusion effect for  $L_w T_n \leq 100 \rightarrow$  upper bound on f

#### 1) Non-local EWBG

• Correct baryon asymmetry for  $3\text{TeV} \le f \le 10\text{TeV}$ 



(axion mass: GeV to 20 GeV)

#### II. AHP- Model for Electroweak Baryogenesis- Baryon Asymmetry

#### 2) Local EWBG KSJ, Jung, Shin 1811.03294

CP violation from axion-dependent EW @-term

 $\frac{\phi}{f}W\widetilde{W} \rightarrow \frac{d\phi}{dt}$  = chemical potential for baryon number

- Simultaneous B and CP violations across thick walls
  - B generation through EW anomaly

$$\frac{dn_B}{dt} = \frac{3}{2} \frac{\Gamma_{\text{sph}}}{T} \frac{d}{dt} \frac{\phi}{f} - \frac{39}{4} \frac{\Gamma_{\text{sph}}}{T^3} n_B$$

$$\uparrow$$
sphaleron-induced washout

- Wash-out: need to suppress axion oscillations quickly

via thermal dissipation (axion-Higgs mixing)

# 2) Local EWBG

• Correct baryon asymmetry for  $10^5 \text{GeV} \le f \le 10^7 \text{GeV}$ 

r=1.1 (A=130 GeV)  $\eta/\eta_{obs}$ 1.0 Not 1st order 0.8 3 Not strong 1st order  $(v_n < T_n)$ 0.6 2 w 0.4 1 1/10 0.2  $f = 10^{6} {
m GeV}$ Metastable 0 Not global minimum 0.0 0.0 0.5 1.0 1.5 2.0 a

(axion mass: MeV to GeV)

#### II. AHP- Model for Electroweak Baryogenesis- Baryon Asymmetry

#### How to probe axion-driven EWBG

- Axion at MeV-GeV (local) or GeV-20GeV (non-local)
  - window without strong theoretical interests so far





axion-Higgs mixing: rare B-meson decays, beam dump

#### II. AHP- Model for Cogenesis

#### Axionic cogenesis KSJ, Sang Hui Im 2012.xxxx

• Axionic Higgs portal with  $M \ll m_Z$ 

$$V = \lambda |H|^4 + \mu^2 |H|^2 - \frac{M^2 \cos\left(\frac{\phi}{F}\right)}{|H|^2} - \frac{\Lambda^4 \cos\left(\frac{\phi}{F} + \alpha\right)}{|H|^2}$$

axion evolution is triggered by EW symmetry breaking



CP violation from anomalous coupling to EW gauge bosons

$$\frac{1}{16\pi^2} \frac{\phi}{f} W_{\mu\nu} \widetilde{W}^{\mu\nu}$$

axion evolution  $\frac{d}{dt} \frac{\phi}{f}$ : chemical potential of baryon number

# **Axionic cogenesis**

- Spontaneous baryogenesis via slow shift of potential minimum  $\frac{\phi}{F}$ 
  - Correct baryon asymmetry for  $F > 10^7 \times f$
  - Clockwork mechanism



collective rotations of N axions

- Dark matter from axion coherent oscillations
  - Cosmologically stable due to tiny mass
  - Photophobic axion to avoid astrophysical constraints
    - $(f < 10^7 \text{GeV for sufficient thermal dissipation})$

### Axion coupling to dark sector

- Axion-driven EWBG with CP violation from
  - axion-dependent top quark Yukawa
  - anomalous axion coupling to EW gauge bosons
- Dark sector
  - Dark photon dark matter:  $\frac{\phi}{f} \tilde{F}'_{\mu\nu} F'^{\mu\nu}$
  - Fermionic dark matter:  $m_{\chi}e^{ic\frac{\varphi}{f}}\bar{\chi}\chi$





Dark

Sector

φ

SM

#### III. Summary

#### Axion

- Feebly interacting light particle with properties controlled by perturbative shift symmetry
- Strong CP problem, dark matter, inflation, ...

## **Axionic Higgs portal**

- Electroweak hierarchy: cosmological relaxation
- Matter-antimatter asymmetry via EW baryogenesis
- Dark matter via freeze-in or misalignment